

A measurement of a relativistic pair-plasma beam instability at the HiRadMat Facility (CERN)

Jack Halliday^{1,7*}, Charlie Arrowsmith¹, F. Miniati¹, P. J. Bilbao², P. Simon^{3,4}, E. Andersen⁵, A. F. A. Bott¹, S. Burger³, H. Chen⁶, F. D. Cruz², T. Davenne⁷, A. Dyson¹, I. Efthymiopoulos³, D. H. Froula⁸, A. Goillot³, J. T. Gudmundsson^{9,10}, D. Haberberger⁸, T. Hodge^{1,11}, B. T. Huffman¹, S. Iaquinta¹, E. E. Los¹, G. Marshall¹¹, B. Reville¹², P. Rousiadou¹³, S. Sarkar¹, A. A. Schekochihin¹, L. O. Silva², R. Simpson⁶, V. Stergiou^{1,3,14}, R. M. G. M. Trines⁷, T. Vieu¹², S. Zhang¹, N. Charitonidis³, R. Bingham^{7,15}, G. Gregori¹

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*jack.halliday@stfc.ac.uk – jackhalliday.github.io/

¹University of Oxford, ²Instituto Superior Técnico, Lisboa, ³European Organization for Nuclear Research (CERN), ⁴GSI Helmholtzzentrum für Schwerionenforschung GmbH, ⁵University of Bergen, ⁶Lawrence Livermore National Laboratory, ⁷STFC Rutherford Appleton Laboratory, ⁸University of Rochester Laboratory for Laser Energetics, ⁹University of Iceland, ¹⁰KTH Royal Institute of Technology, Stockholm, ¹¹AWE, ¹²Max-Planck-Institut für Kernphysik, Heidelberg, ¹³University of Ioannina, ¹⁴National Technical University of Athens, ¹⁵University of Strathclyde.

Mass symmetry in pair plasmas leads to different plasma behaviour to traditional plasmas, and they can enrich outflows from extreme astrophysical objects





Pair plasma processes play a fundamental role in energy dissipation and radiative emission...

...but models have never been tested in the laboratory.



High-yield pair beams can be produced using ultra-relativistic protons accelerated by the Super Proton Synchrotron at CERN.



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HiRadMat

High-Radiation to Materials

A laboratory testbed for relativistic electron positron plasma studies has been developed at the HiRadMat facility, driven using the Super Proton Synchrotron at CERN



For an ambient plasma, a meter-scale inductively coupled argon plasma discharge was constructed, and the plasma parameters characterized using a Langmuir probe.





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No evidence of filamentation is observed in the transverse beam profile, imaged using luminescence screens.



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Faraday rotation using a magneto-optic crystal achieves a magnetic field sensitivity of B~5 mT, but no significant fields are observed.



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Suppressed beam instability due to the finite thermal spread of the beam is the most plausible explanation, which is confirmed by 3D particle-in-cell simulations.



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Assuming the instrument sensitivity as the upper bound, we estimate an upper bound for the growth rate of the magnetic field

FLUKA + Ampere's Law $\Rightarrow B_0 = 0.78 \text{ mT}$ **FLUKA** simulated 3 field profile Faraday sensitivity $\Rightarrow B_{exp} \leq 5 \text{ mT}$ $B_{\phi}(r)$ (mT) T $\langle \Gamma_{\rm exp} \rangle \lesssim t_{\rm prop}^{-1} \ln \left(\frac{\langle B_{\rm exp} \rangle}{B_0} \right) \approx 0.7 \, {\rm ns}^{-1}$ 0 Faraday Propagation = 2.7 ns 0.5 0.0 1.0 1.5 probe r (cm) (Annual)

Scaling the growth rate and saturated magnetic field amplitude





Theoretical growth rate:

$$\Gamma_{\rm fastest} = \sqrt{2/3} \frac{\omega_p}{\Delta \theta} \left(\frac{n_{\pm}}{2n_p \gamma_{\pm}} \right)^{\frac{2}{3}} \qquad \left\{ \Delta \theta \gg \left(\frac{n_{pm}}{2n_p \gamma_{\pm}} \right)^{\frac{1}{3}} \right\}$$

Scaling relation:

$$\Gamma_{\rm blz,sc}[{\rm s}^{-1}] \le 3 \times 10^{-11} \left(\frac{\Gamma_{\rm exp}}{0.7 \text{ ns}^{-1}} \right)$$

	Pair beam			Plasma	
Parameter	$n_{\pm} (\mathrm{cm}^{-3})$	$\langle \gamma_{\pm} \rangle$	$\Delta heta$	$n_{ m p}({ m cm}^{-3})$	$ u_{ m e}/\omega_{ m p}$
Experiment Typical blazar jet	5×10^{10} 10^{-23}	$\frac{10^3}{10^5}$	$0.025 \\ 10^{-4}$	10^{12} 2 × 10 ⁻⁷	10^{-3} 10^{-13}

Pair density, mean Lorentz factor and transverse momentum spread of beam, and electron density and collisionality of plasma.



