

# Radiatively Driven Experiments on the MAGPIE Pulsed-Power Generator

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V. Valenzuela Villaseca<sup>5,2</sup>, S. Merlini<sup>2</sup>, T. Mundy<sup>2,6</sup>, S. N. Bland<sup>2</sup>, S. V. Lebedev<sup>2</sup>

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<sup>5</sup> Princeton University; <sup>6</sup> Sandia National Labs

Imperial College  
London



MAGPIE

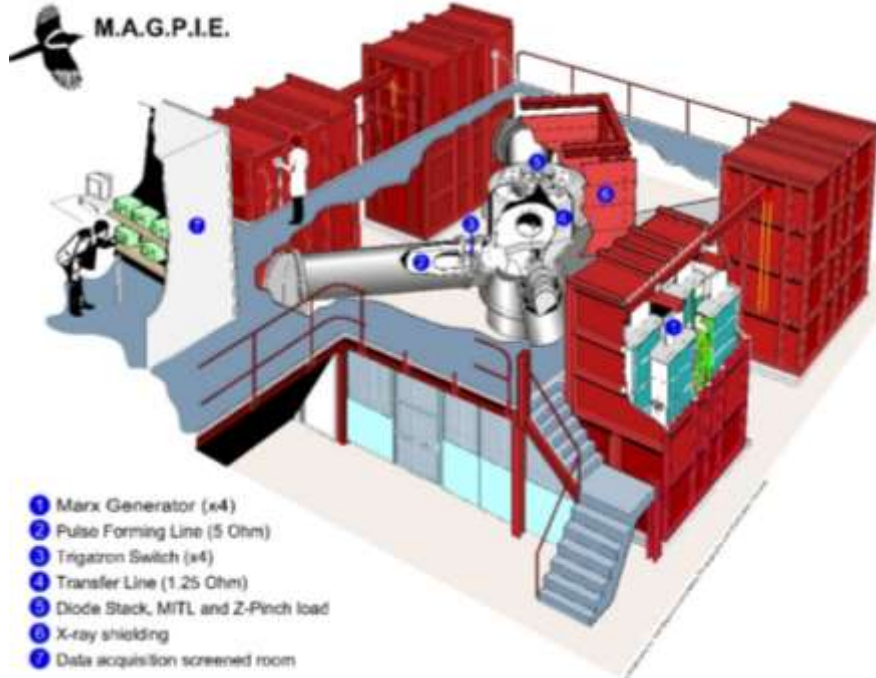


University of Nevada, Reno

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# Experimental facility and diagnostics

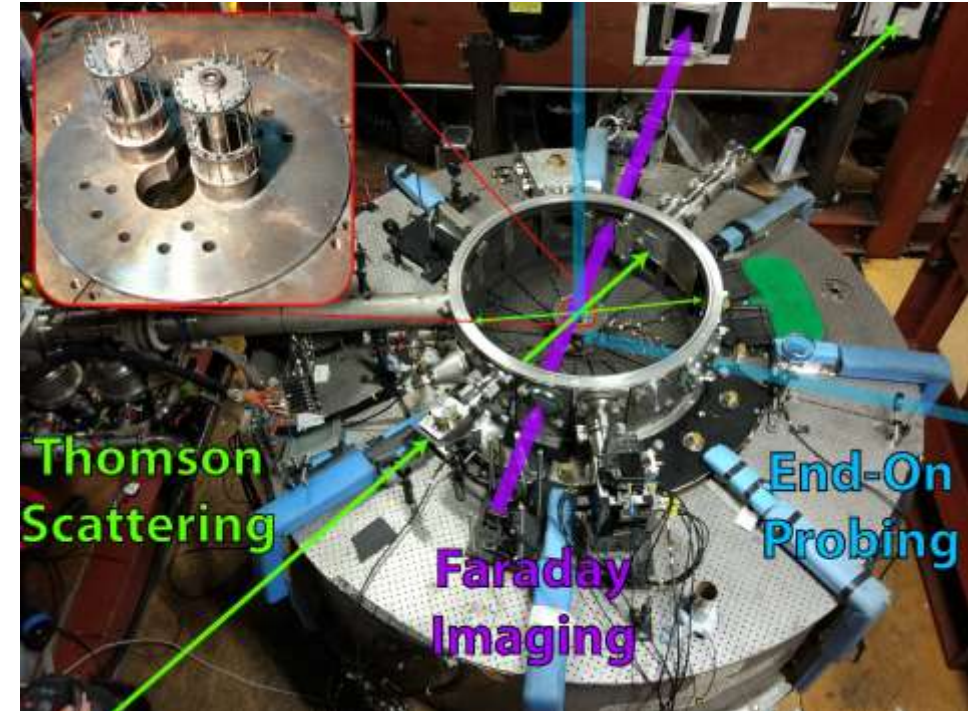
1.4 MA, 240 ns, 1 TW, 250 kJ



~ 30 kJ delivered to a load

Plasma scales:  $\begin{cases} L \sim 10 \text{ mm} \\ \tau \sim 400 \text{ ns} \end{cases}$

Load region



$\langle B_y \rangle$

Faraday rotation

$\vec{V}_{fl}, \vec{V}_d, ZT_e, T_i$

Thomson scattering

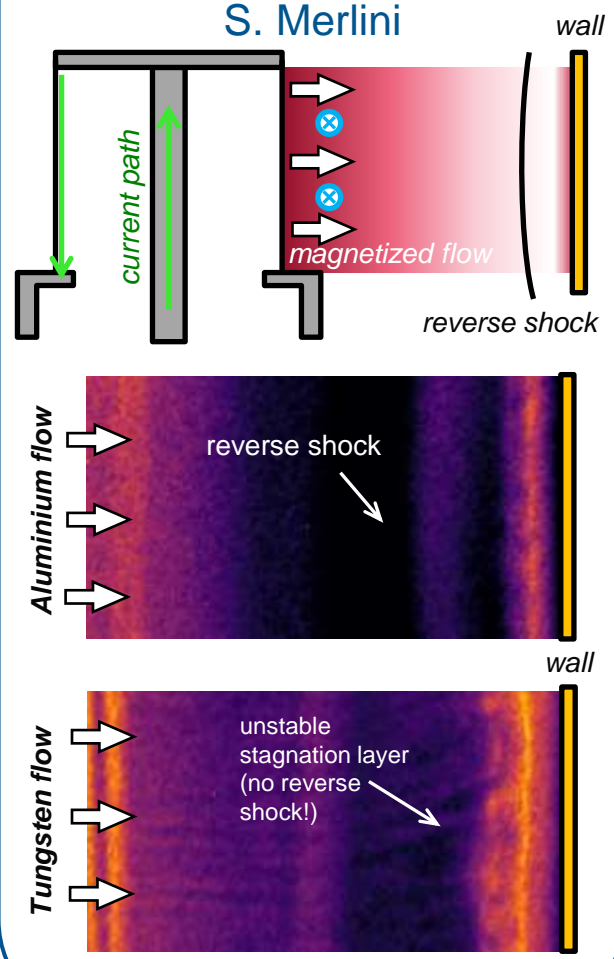
$n_e L$

Imaging interferometry

# Current driven laboratory astrophysics on MAGPIE

## Instabilities in magnetized shock experiments

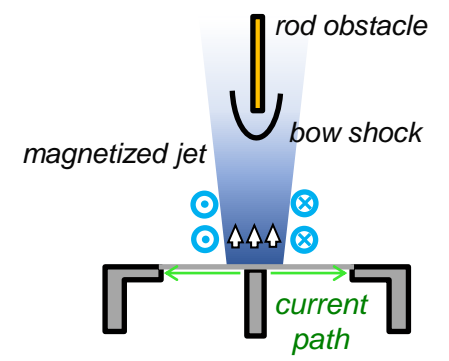
S. Merlini



S. Merlini+ arXiv:2306.01847

## Shocks in magnetic tower jet experiments

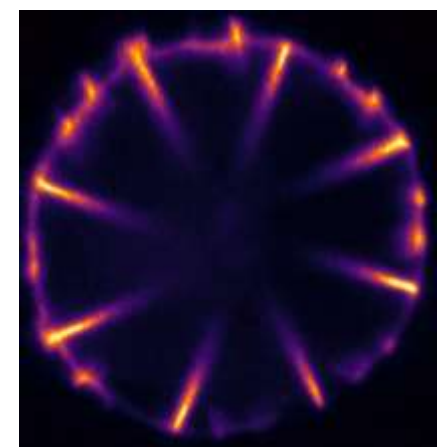
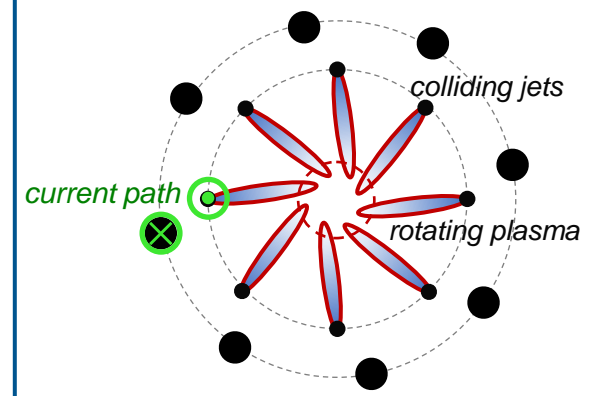
D. Russell / F. Suzuki Vidal



D. R. Russell+ PRL 2022

## Differentially rotating plasmas

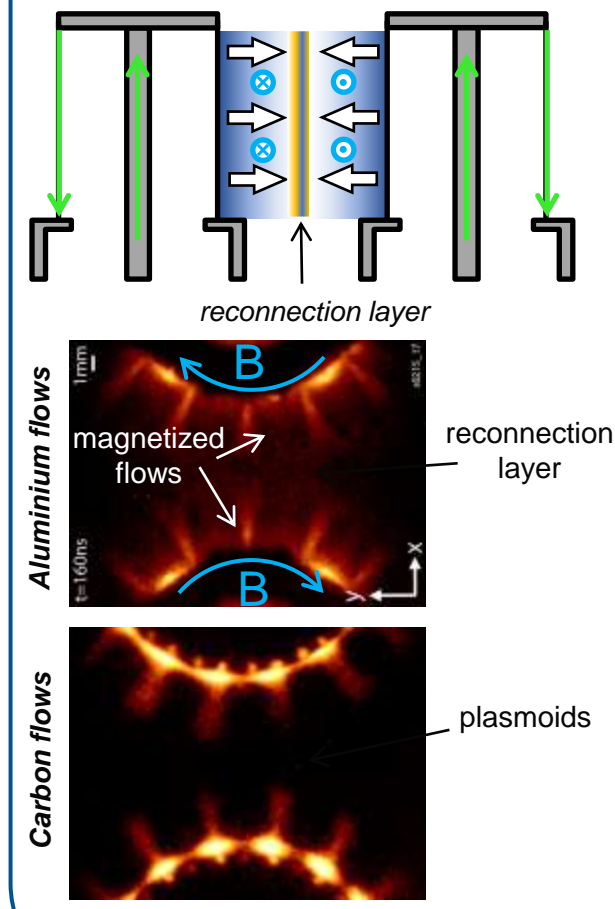
V. Valenzuela-Villaseca



V. Valenzuela-Villaseca+ PRL 2023

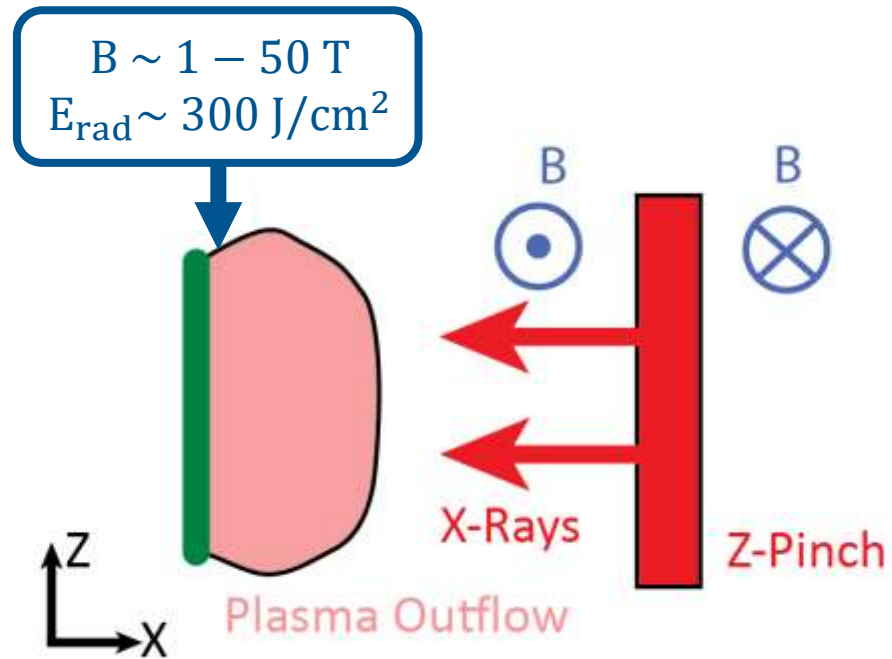
## Magnetic reconnection experiments

L. Suttle / J. Hare

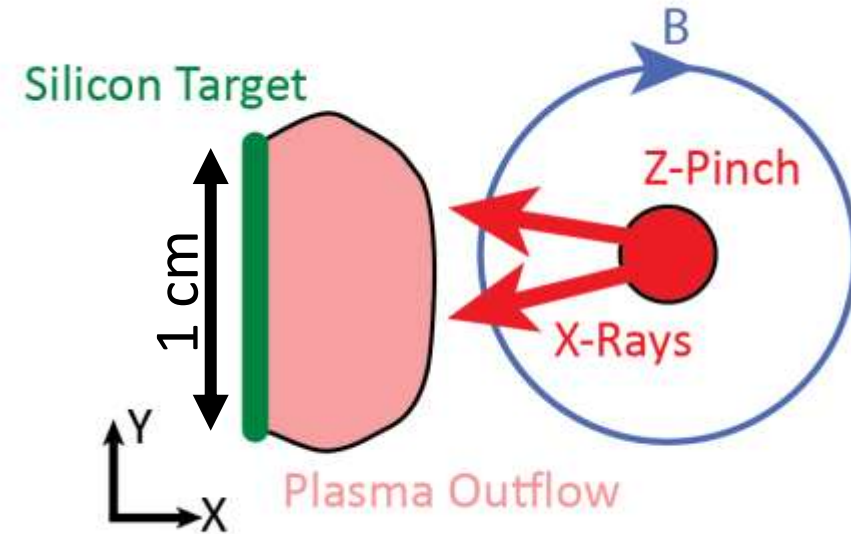


L. G. Suttle+ PRL 2016

# Overview of experimental setup



**Side-On** (X-Z plane) view of the experiment



**End-On** (X-Y plane) view of the experiment

- X-Rays from aluminium wire array Z-Pinch
- $T_B \sim 10$  eV on target
- Ablated silicon plasma expands into magnetic field ( $B \sim 10$  T)
- Target size  $\sim 1$  cm<sup>2</sup>, irradiated uniformly



# Overview & Motivation

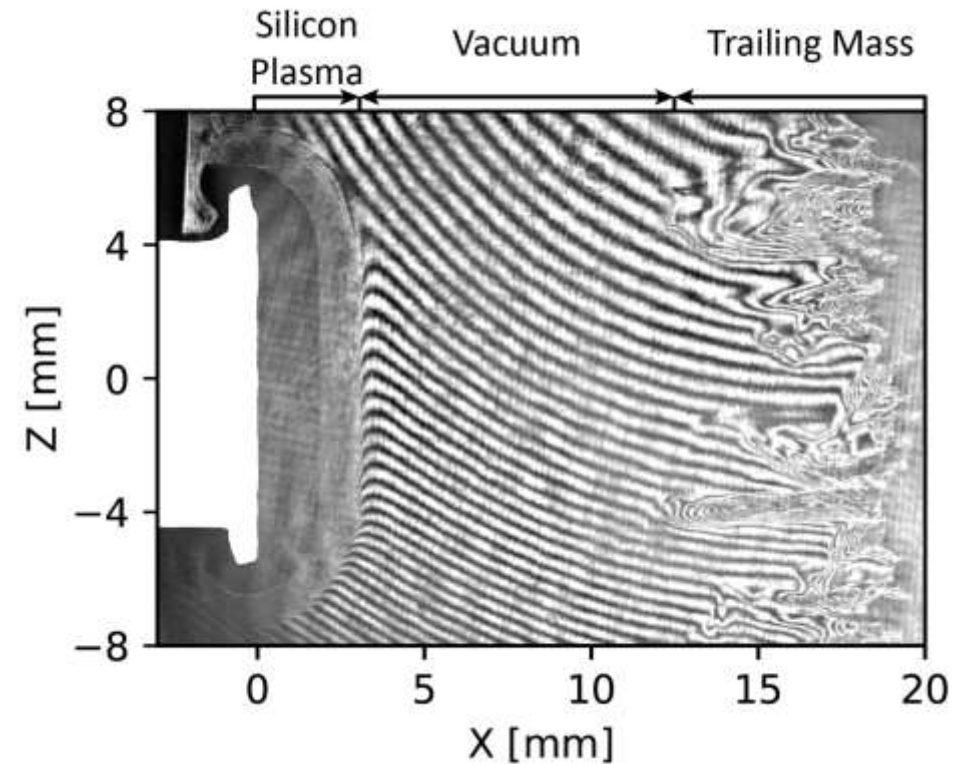
Wire arrays at 1 MA

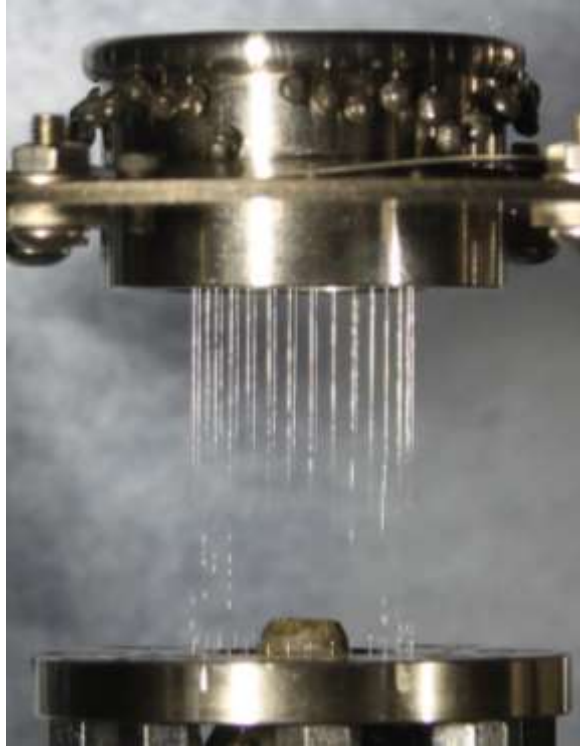
Laser probing measurements [Halliday+ PoP 2022]  
characterize hydrodynamics

**Single target experiments:** Cold plasma, large system size – platform is a great testbed for problems in atomic kinetics

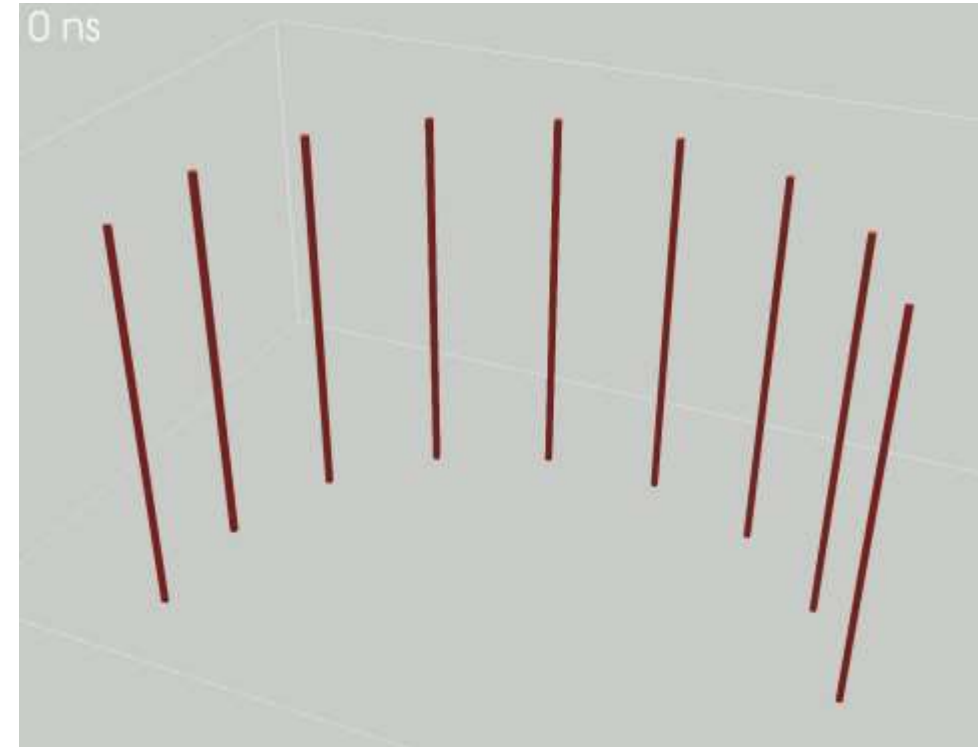
**Two target experiments:** Potential to study radiative instabilities using colliding flows

**Motivation:** Understand how satellites are damaged during nuclear weapon tests in the upper atmosphere.



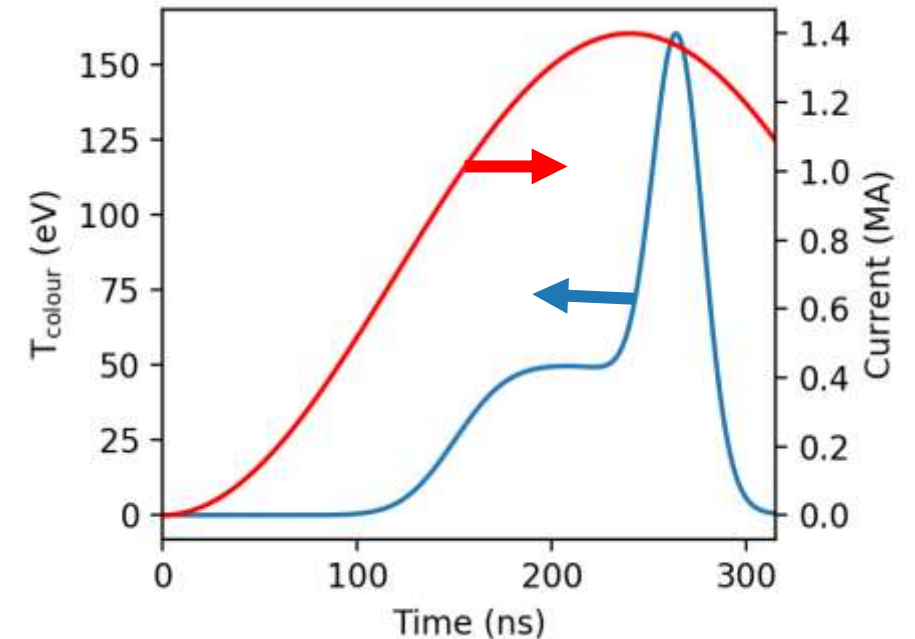


A 32-wire aluminium array used in MAGPIE experiments



Mass density from Gorgon (MHD) simulation

- **Precursor:**
  - Longer pulse
  - Colder spectral character ( $T_c \sim 30$  eV)
  - Radiates  $\sim 400$  J in total
- **Implosion:**
  - Emitted radiation  $\sim 15$  kJ over  $\sim 30$  ns
  - Non-thermal: forest of L shell lines
  - Some K-Shell radiation also
  - Estimate  $T_c \sim 150$  eV

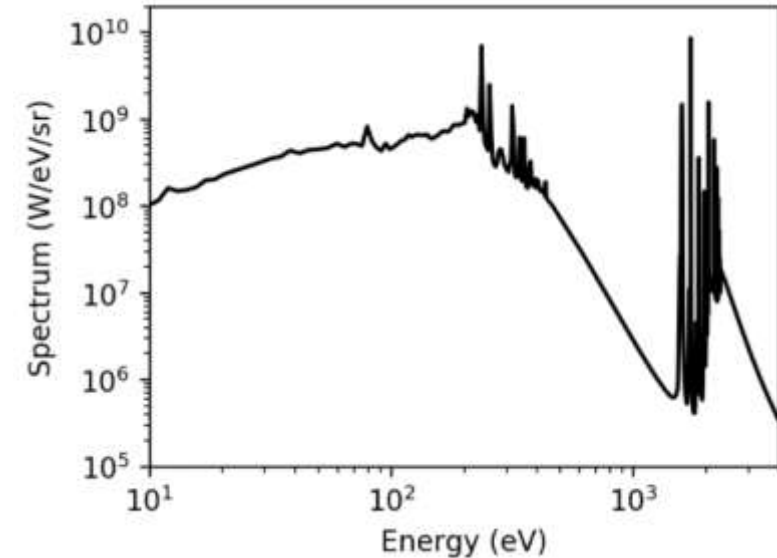


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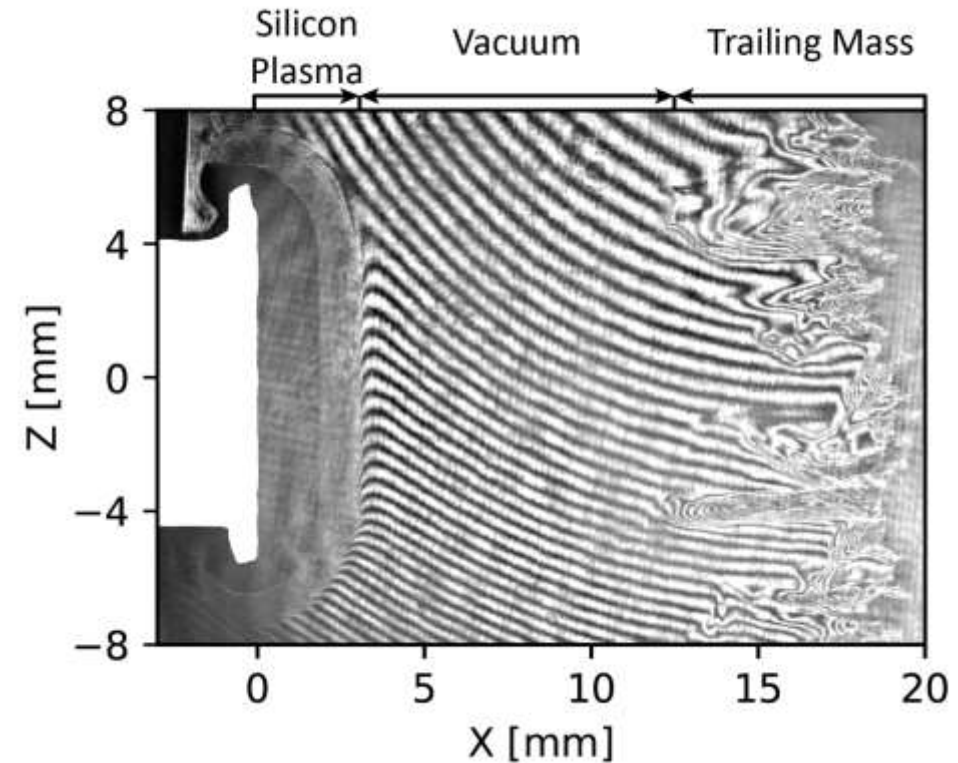
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Laser probing measurements [Halliday+ PoP 2022]  
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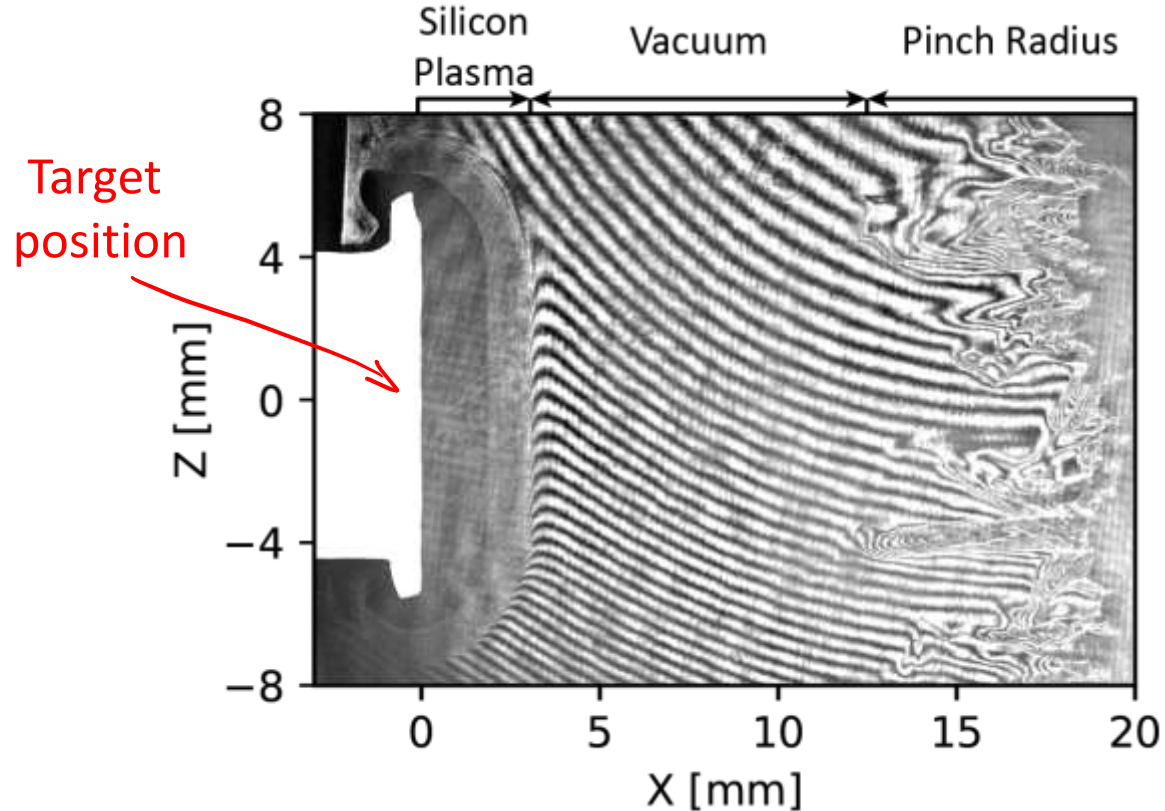
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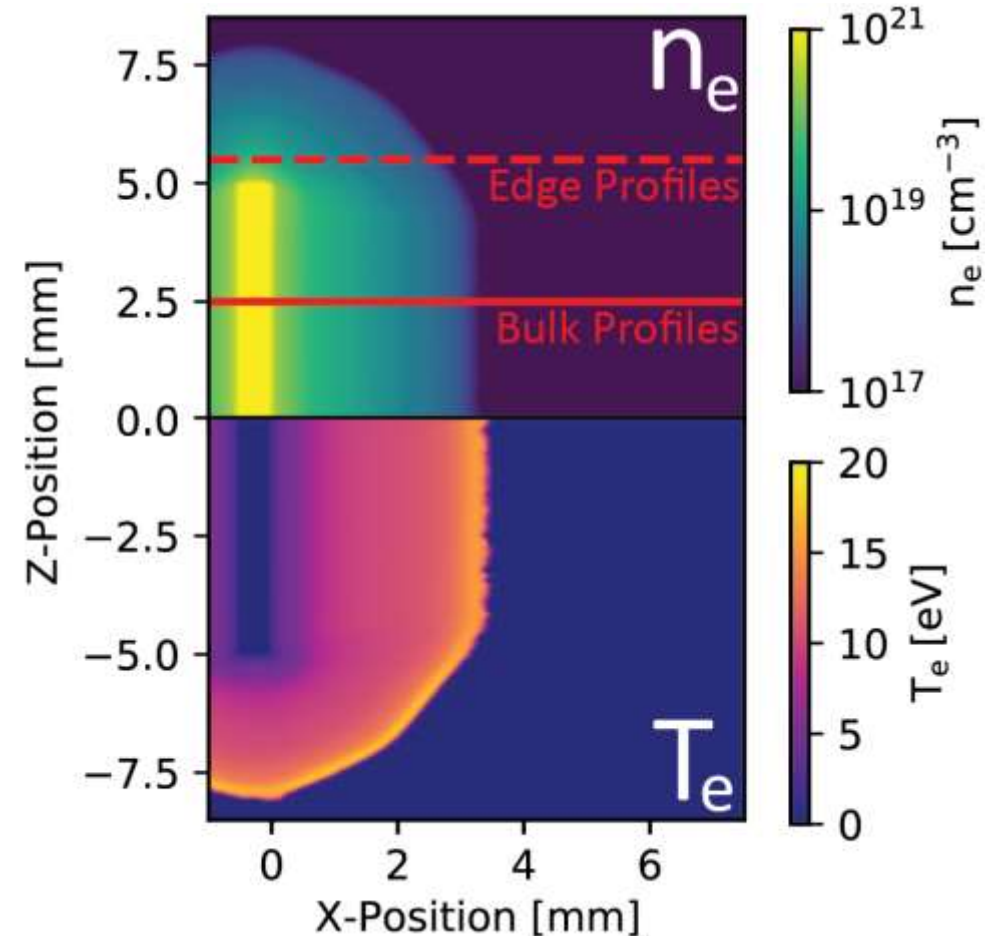
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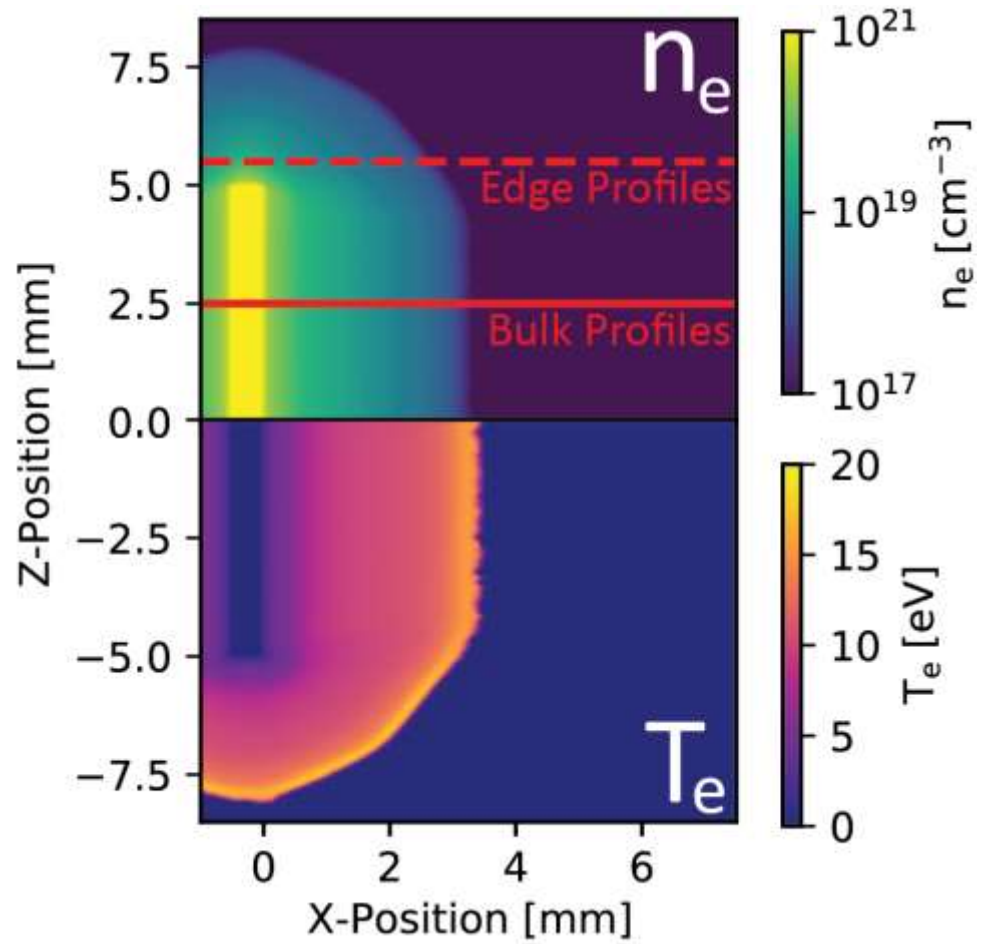
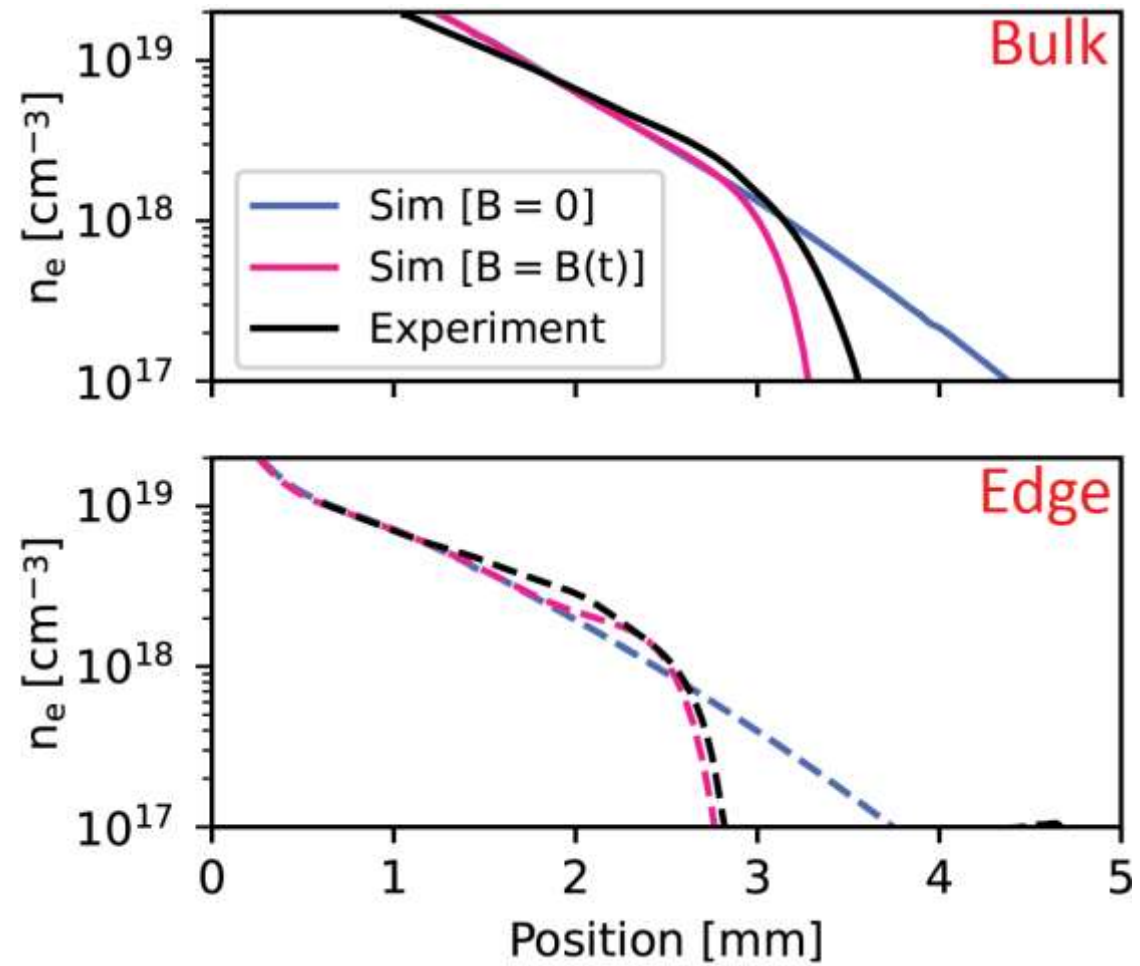
# Plasma outflow is extremely simple



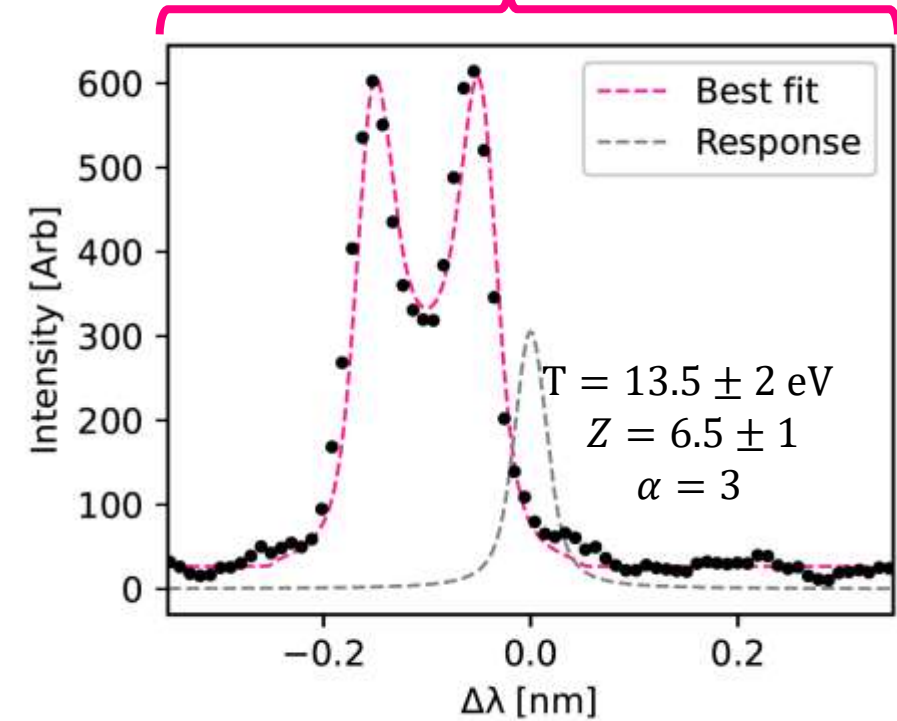
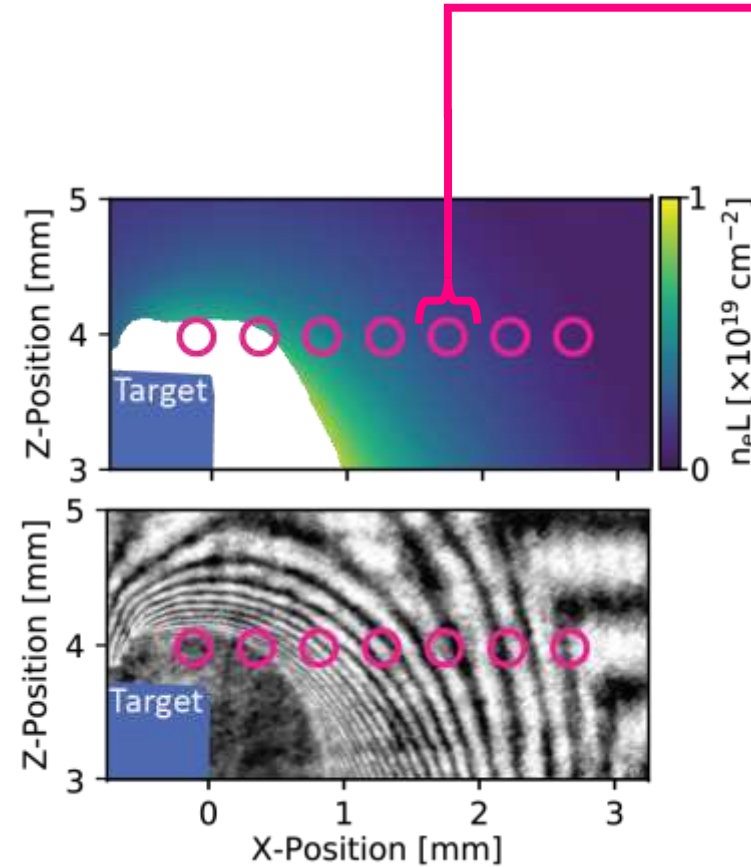
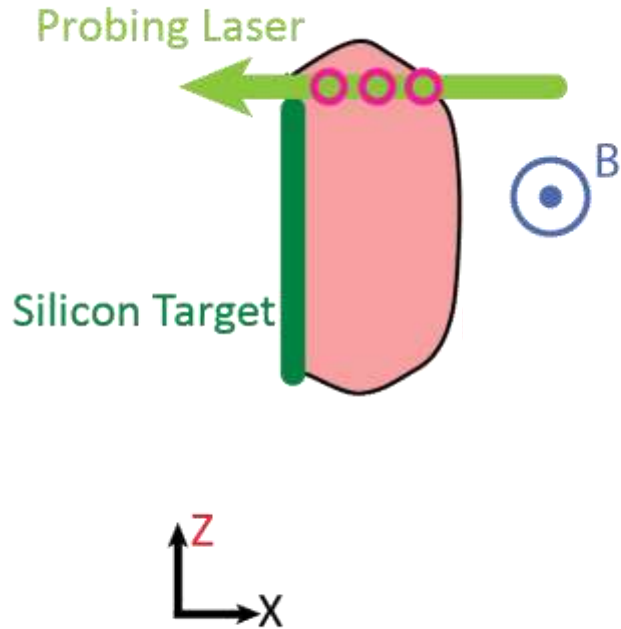
- Smooth  $\sim 1D$  expansion profile confirmed by orthogonal laser probing
- Density profiles well reproduced by R-MHD simulations performed with Chimera



# R-MHD simulations were performed with Chimera



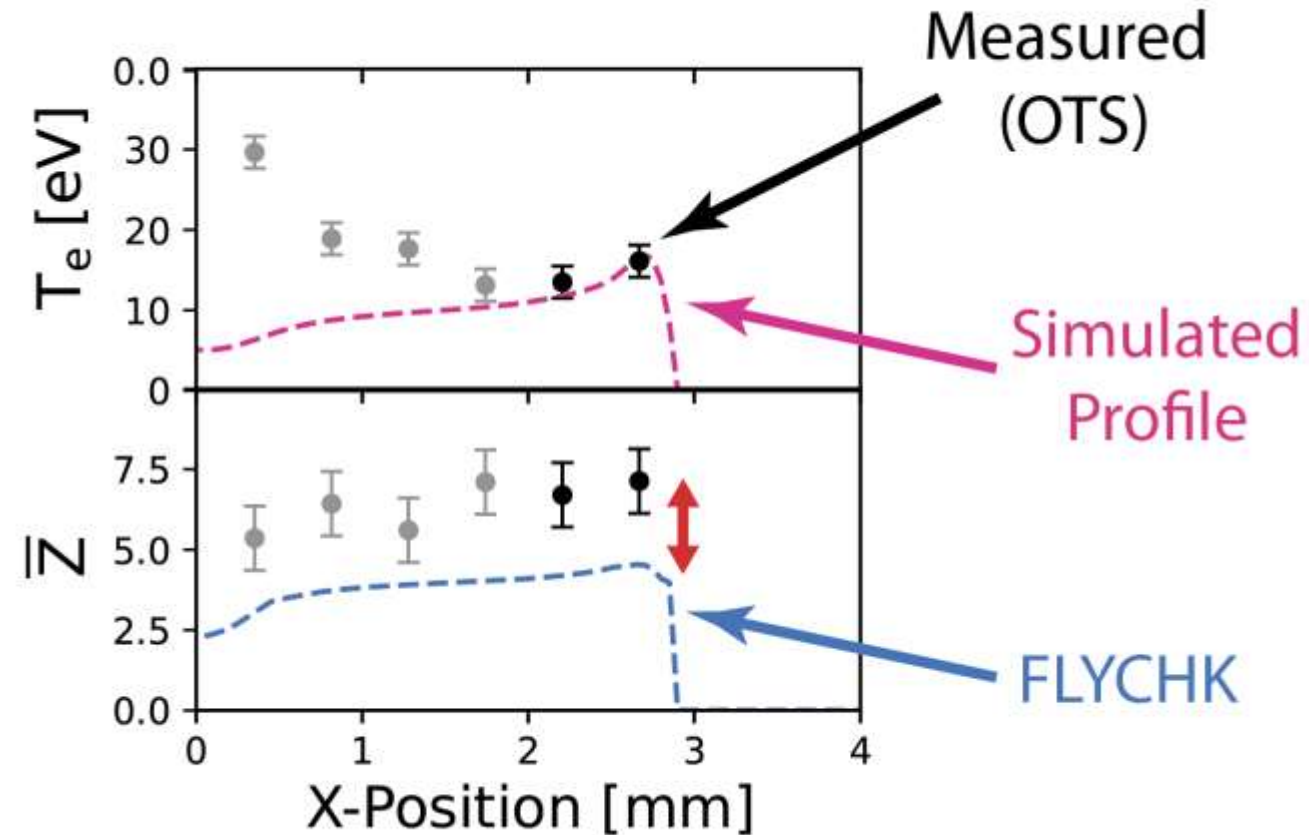
# Thomson scattering [localised diagnosis of $T, V, Z$ ]





# Measured $\bar{Z}$ is higher than the CR equilibrium value

- Data points are experimental measurements from IAW Thomson scattering
- Pink curve indicates results from R-MHD simulations
- Blue curve derived by applying simulated  $n_e$  and  $T_e$  values to FLYCHK simulations (no external radiation field)
- **Significant disagreement in  $\bar{Z}$  values**





# Charge state distribution altered by driving radiation?

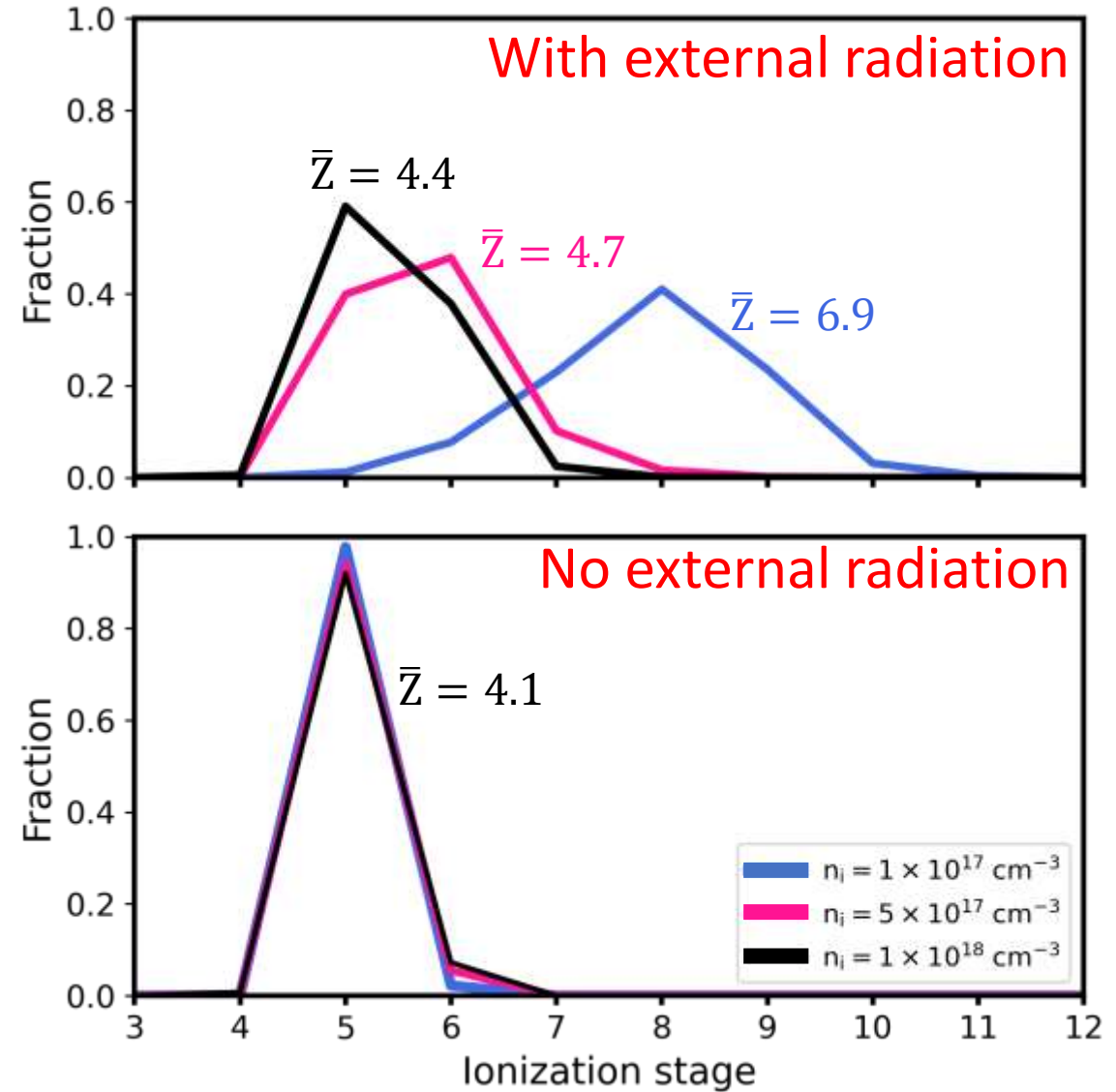
- PrismSPECT results with:

$$T_e = 10 \text{ eV}, \quad n_i = \begin{matrix} 1 \times 10^{17} \text{ cm}^{-3} \\ 5 \times 10^{17} \text{ cm}^{-3} \\ 1 \times 10^{18} \text{ cm}^{-3} \end{matrix}$$

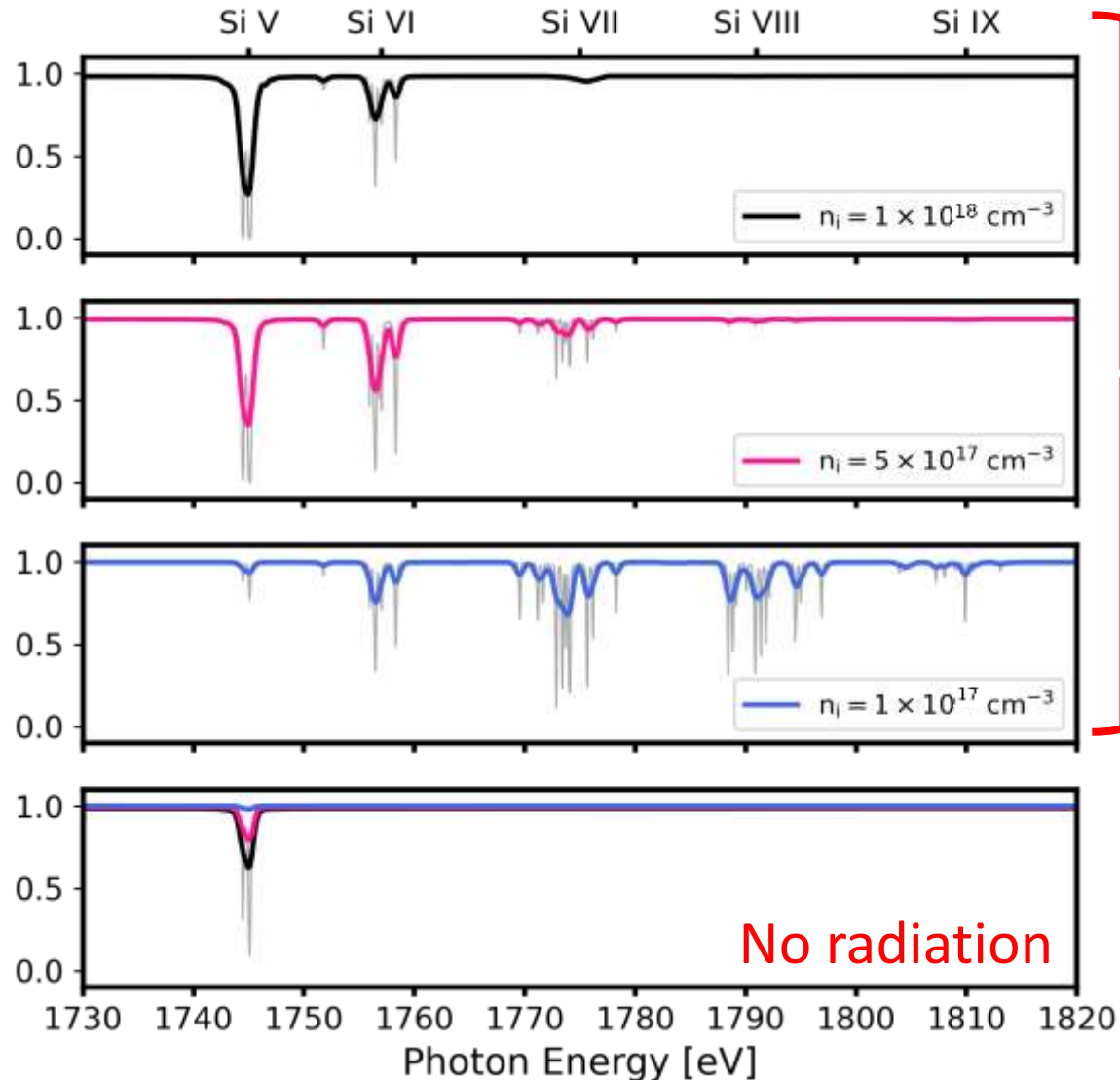
- Applied external radiation field (approximates pinch at peak emission):

$$T_c = 150 \text{ eV}, \quad T_B = 10 \text{ eV}$$

- Steady-state, nLTE simulation
- Driving radiation perturbs charge state distribution**



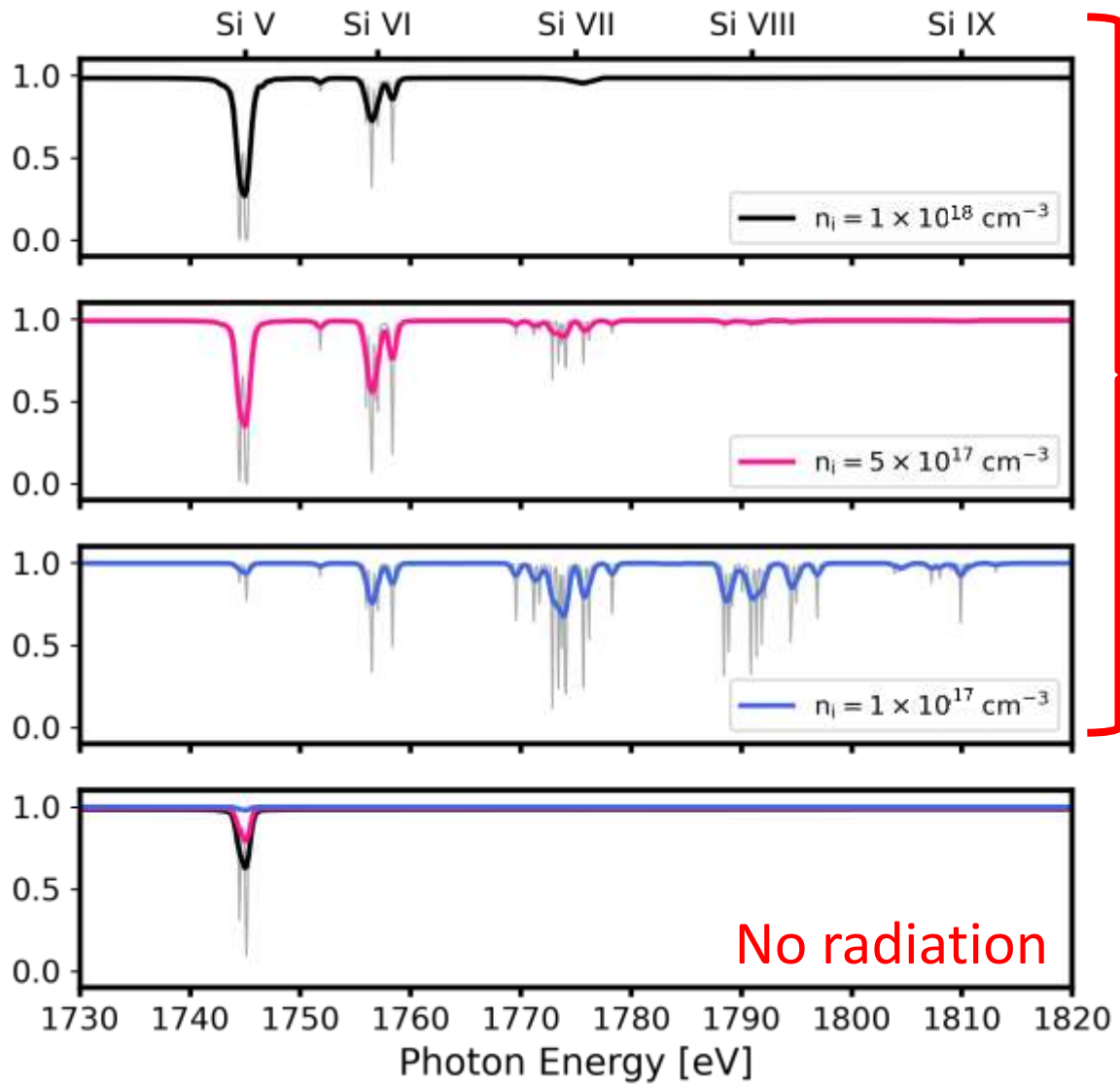
# Diagnose silicon K-Shell absorption features



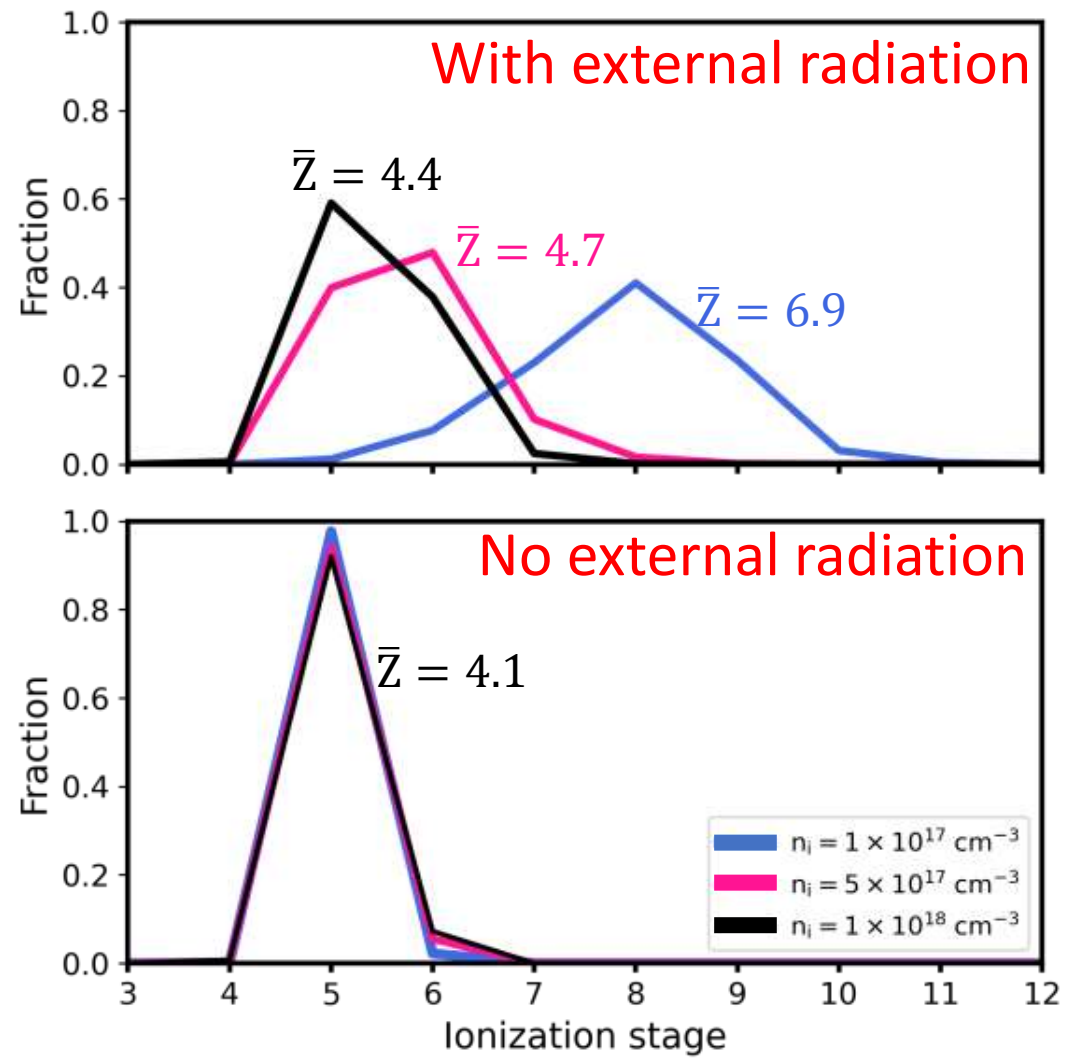
With external radiation

- Transmission spectra are dominated by  $n = 1 \rightarrow 2$  absorption features
- Absorption features for different ionisation stages spectrally separated
- Relative intensity provides diagnostic of charge state distribution
- Instrumental broadening is applied to spectra

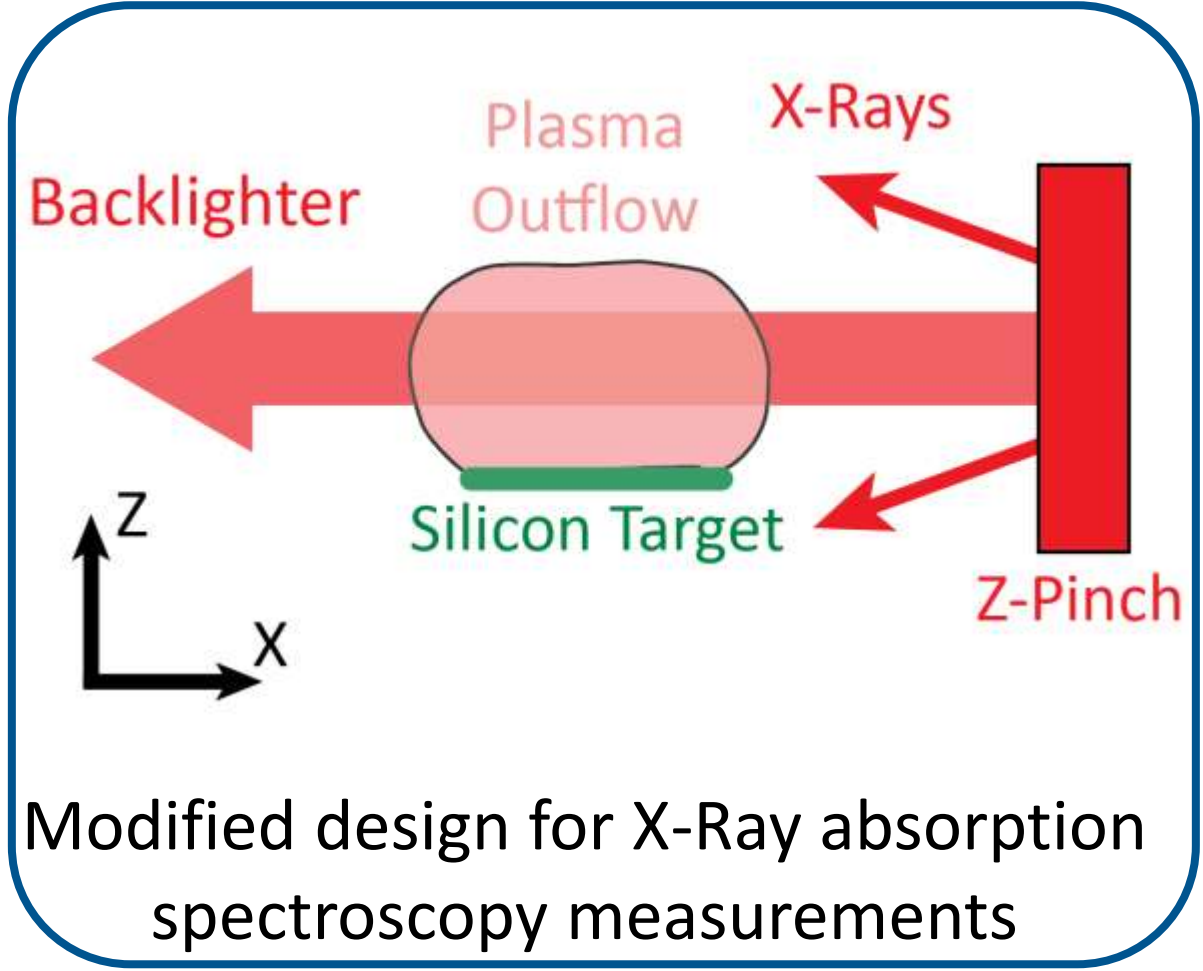
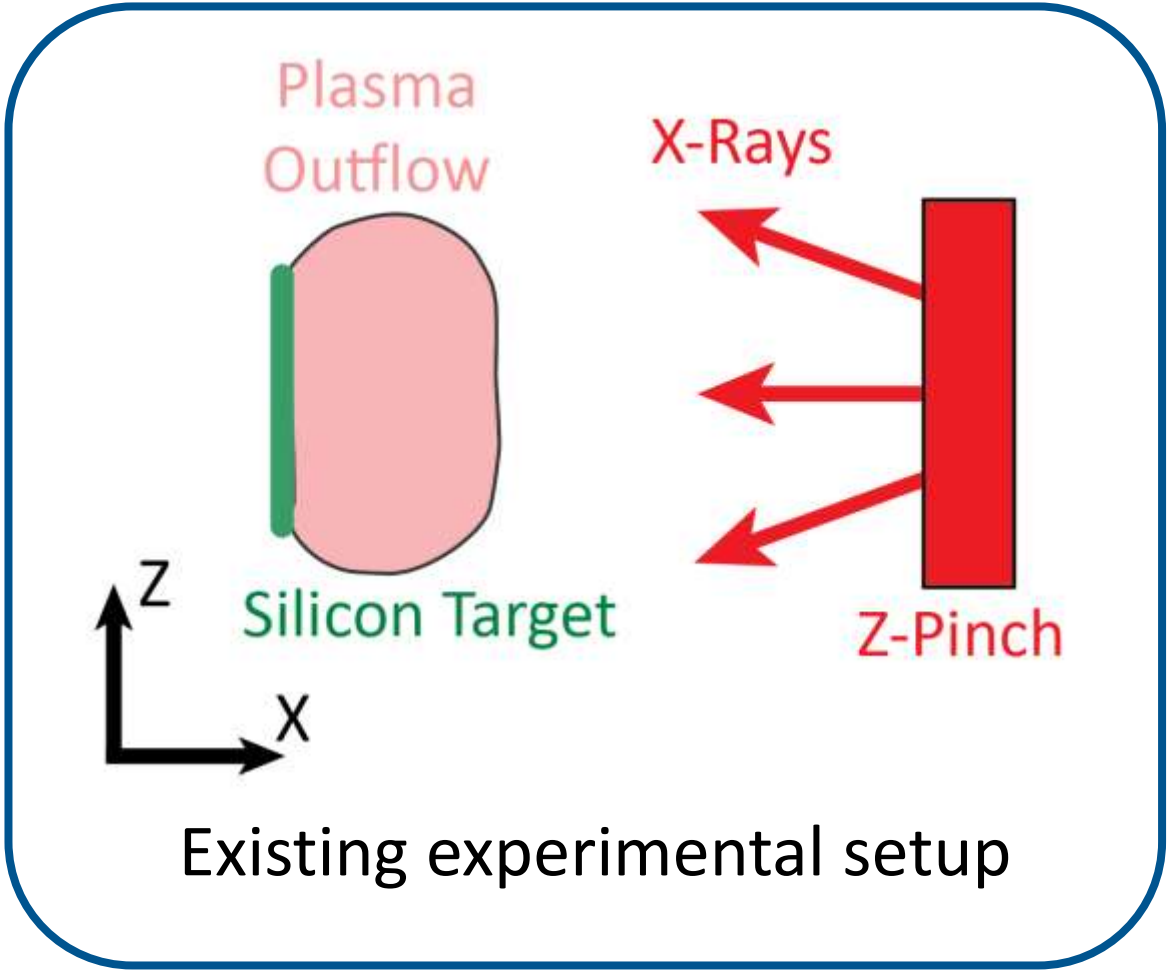
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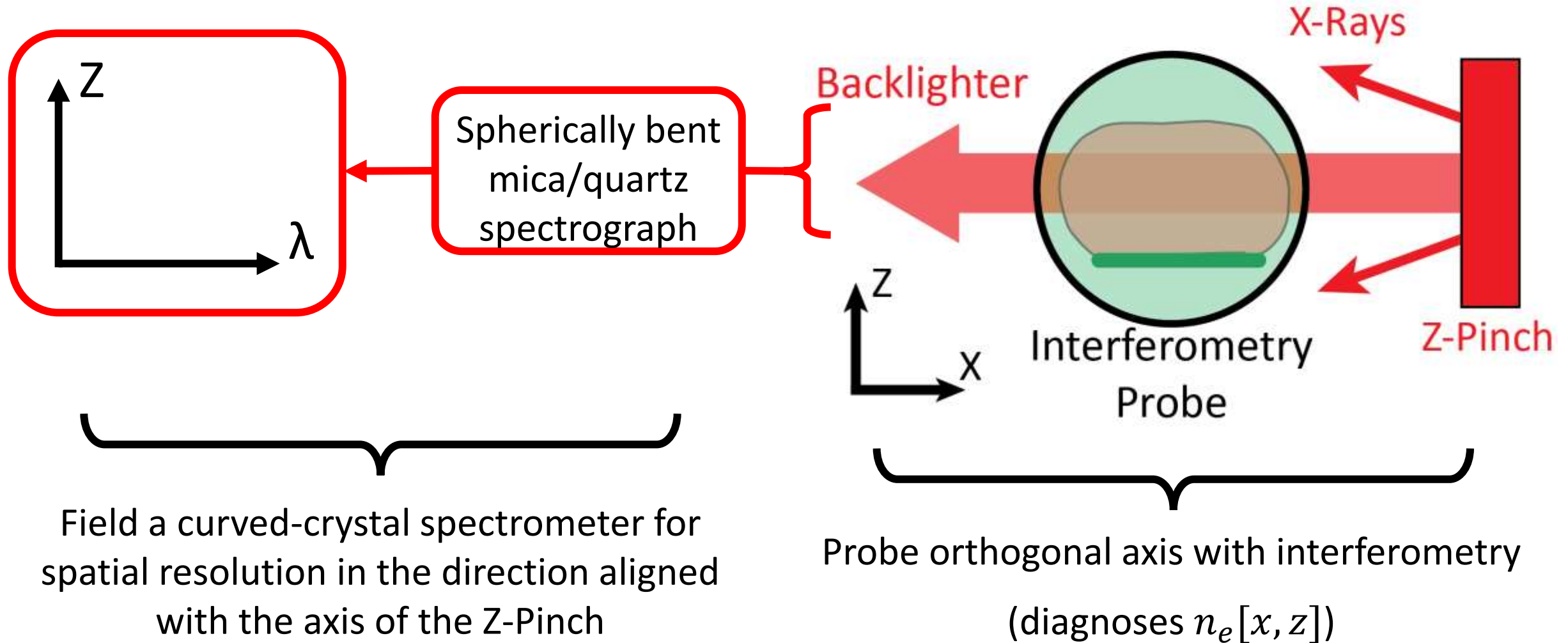
With external radiation



# Modified setup for X-Ray absorption measurements

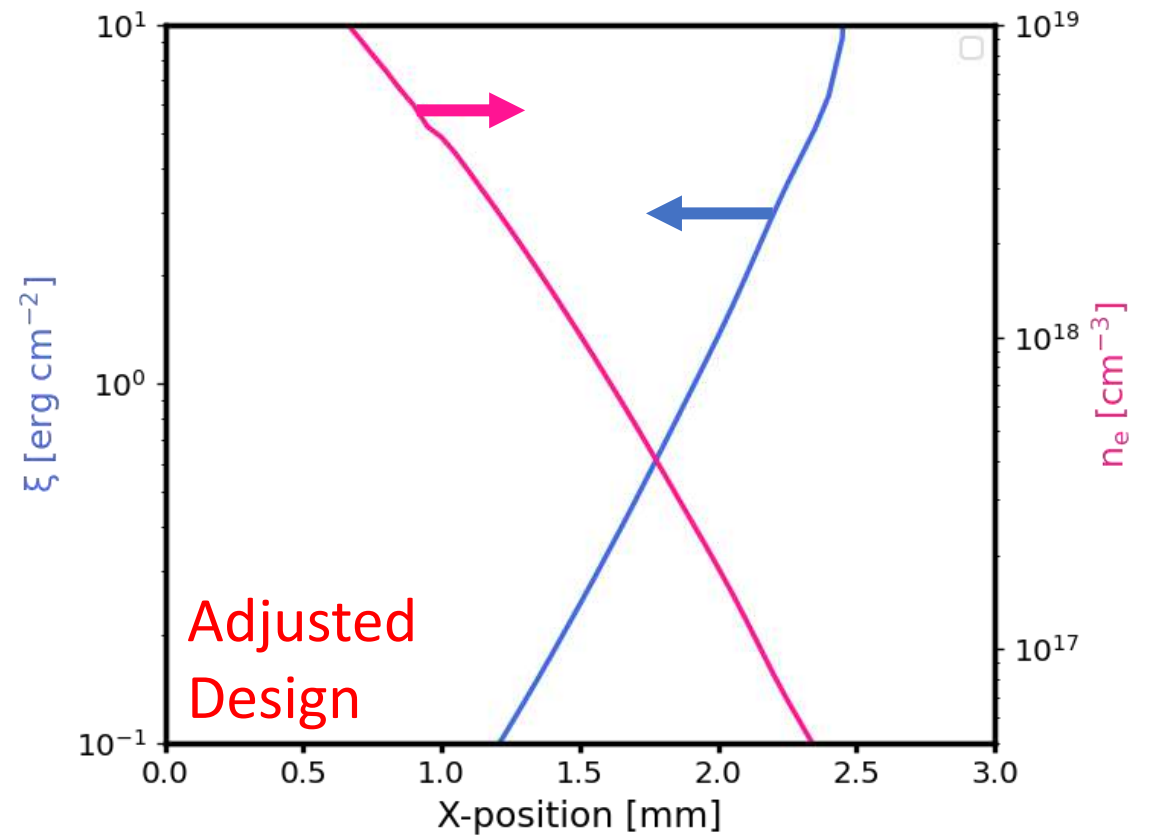
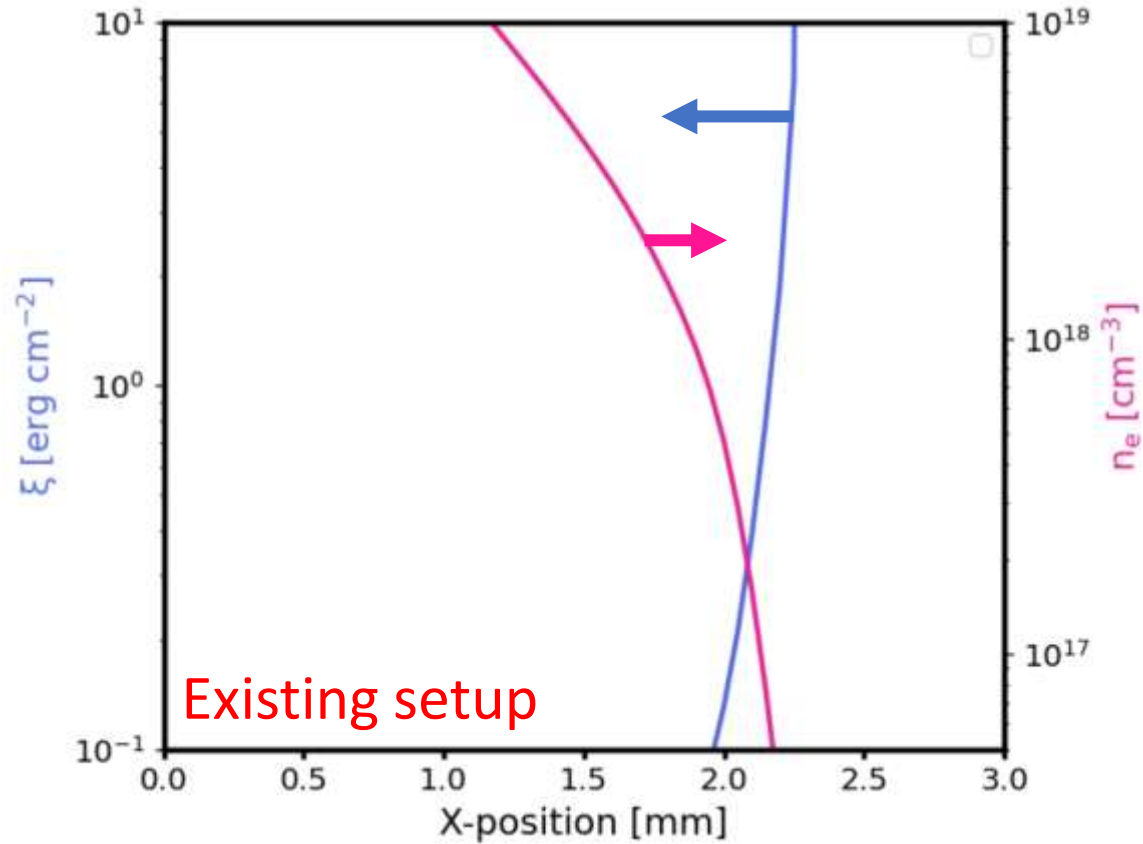


# Spatially resolve spectra to sample range of $n_e$ values





# Potential to access a photoionization relevant regime?

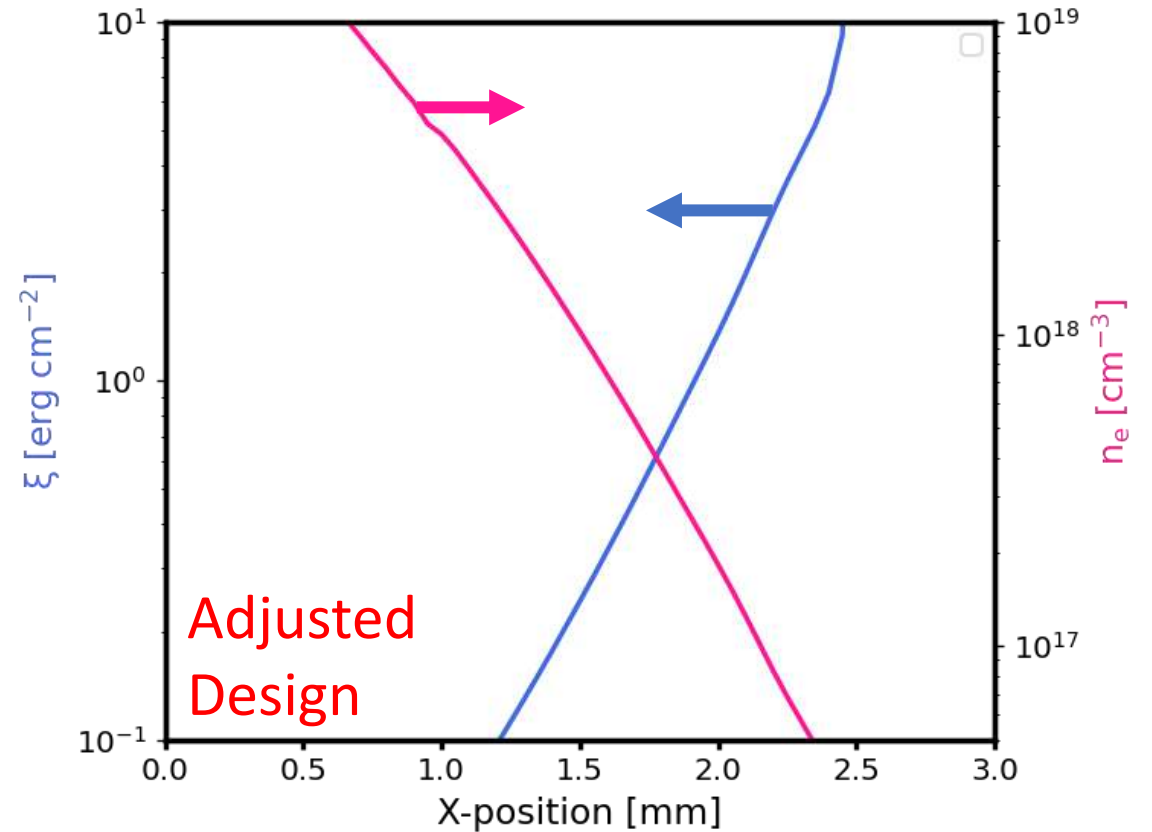


Photoionization parameter,  $\xi = \int d\nu F_\nu / n_e$

# Potential to access a photoionization relevant regime?



**Relevance – accretion disk physics**



Photoionization parameter,  $\xi = \int dv F_\nu / n_e$

# Applications in EUV Photolithography

- Minimum feature size scales with  $\lambda$
- Current state-of-the-art is EUV lithography ( $\lambda \sim 13 \text{ nm} \Rightarrow \varepsilon \sim 900 \text{ eV}$ )
- Source uses laser to heat  $\sim \text{mm}$  scale tin droplets (50 kHz rep-rate!)
- Similar density and temperature conditions to 1 MA Z-Pinch driven experiments
- Experimental data would be helpful to validate atomic kinetics modelling [J. Sheil+ EUVXRAY 2022]



Photo credit: ASML Holdings 2023

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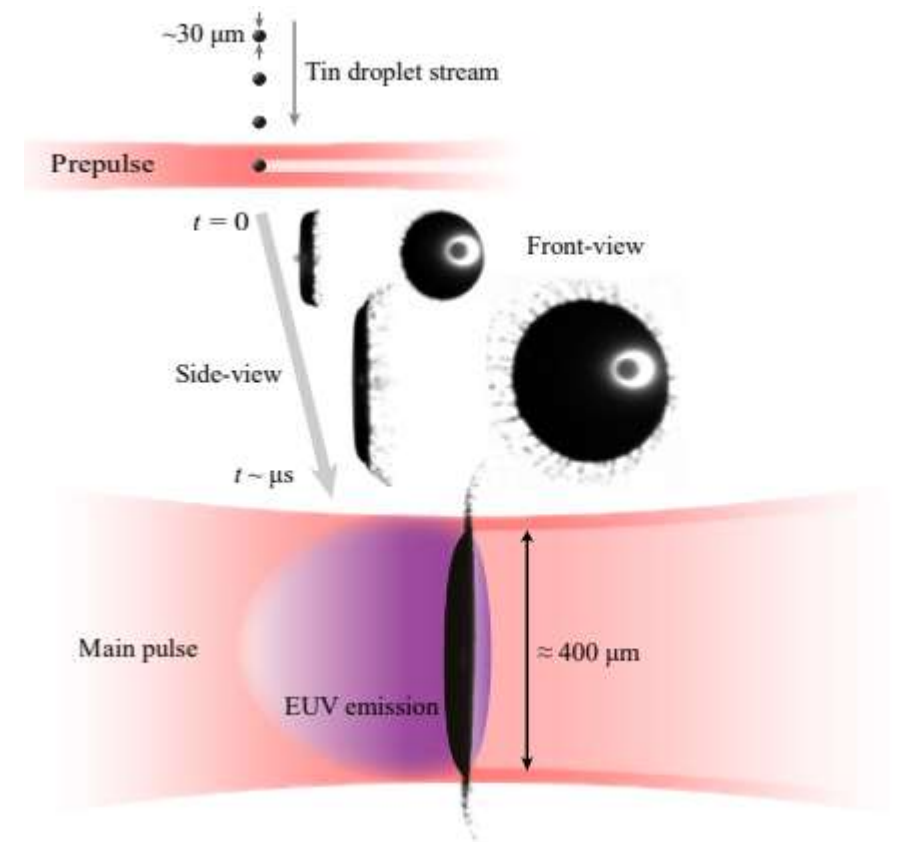


Photo credit: B. Liu PhD Thesis 2022

# Overview & Motivation

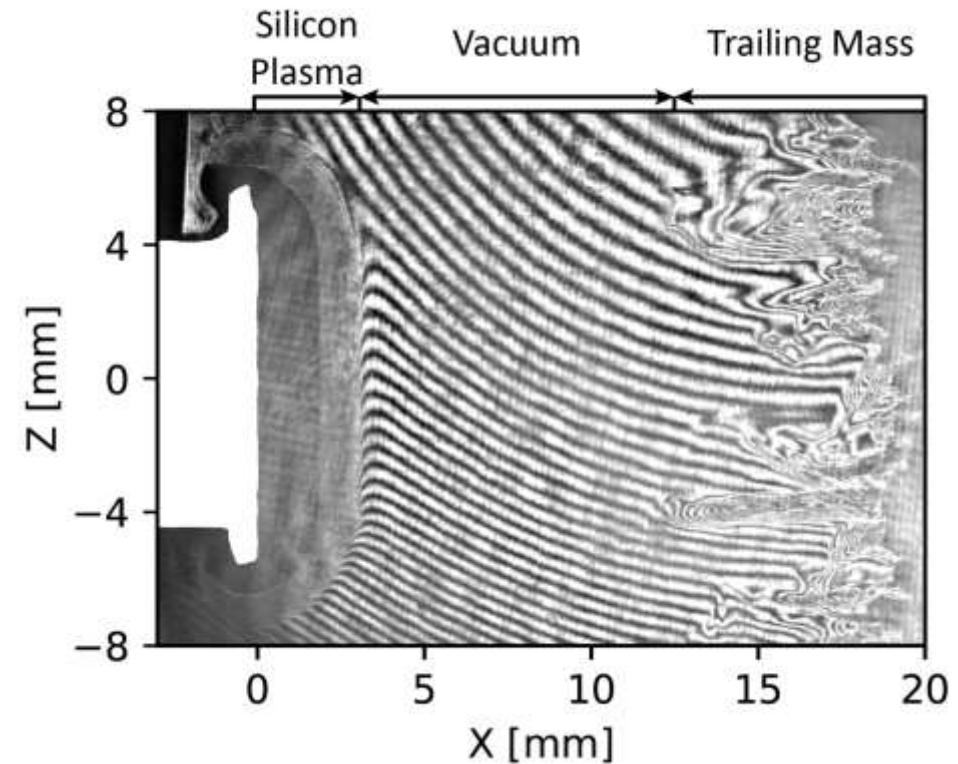
Wire arrays at 1 MA

Laser probing measurements [Halliday+ PoP 2022] characterize hydrodynamics

**My work:** Cold plasma, large system size – platform is a great testbed for problems in atomic kinetics (also radiation hydrodynamics)

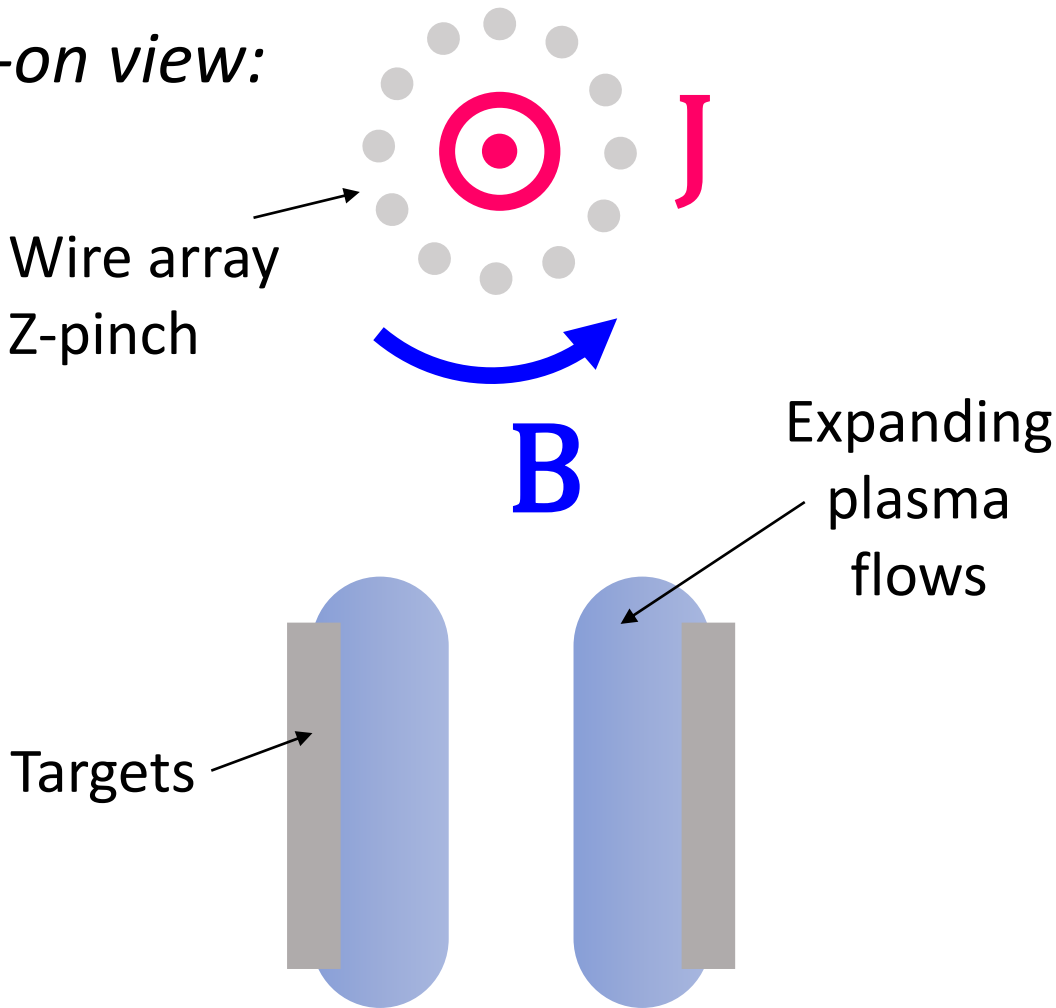
**Work from MAGPIE:** Potential to study radiative instabilities using colliding flows

**Motivation:** Understand how satellites are damaged during nuclear weapon tests in the upper atmosphere.

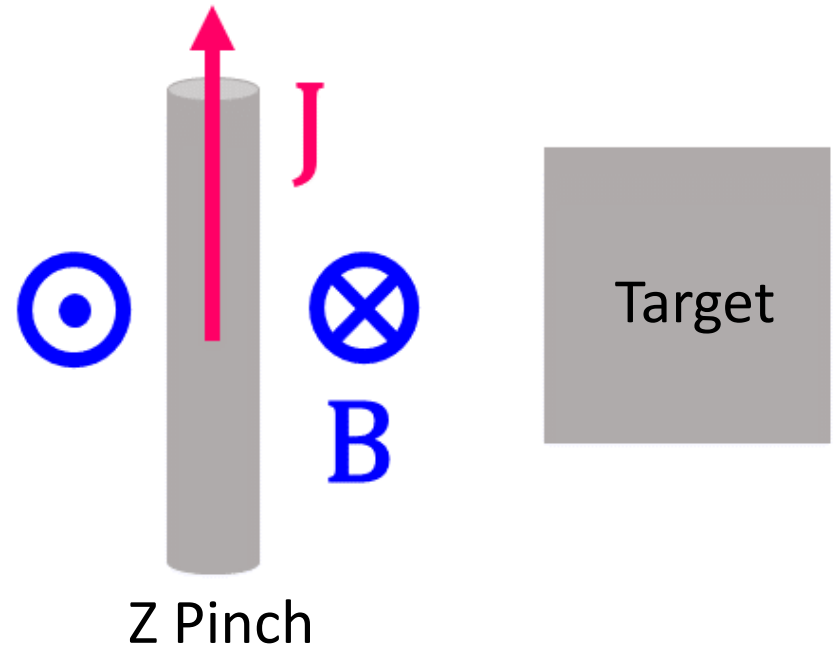




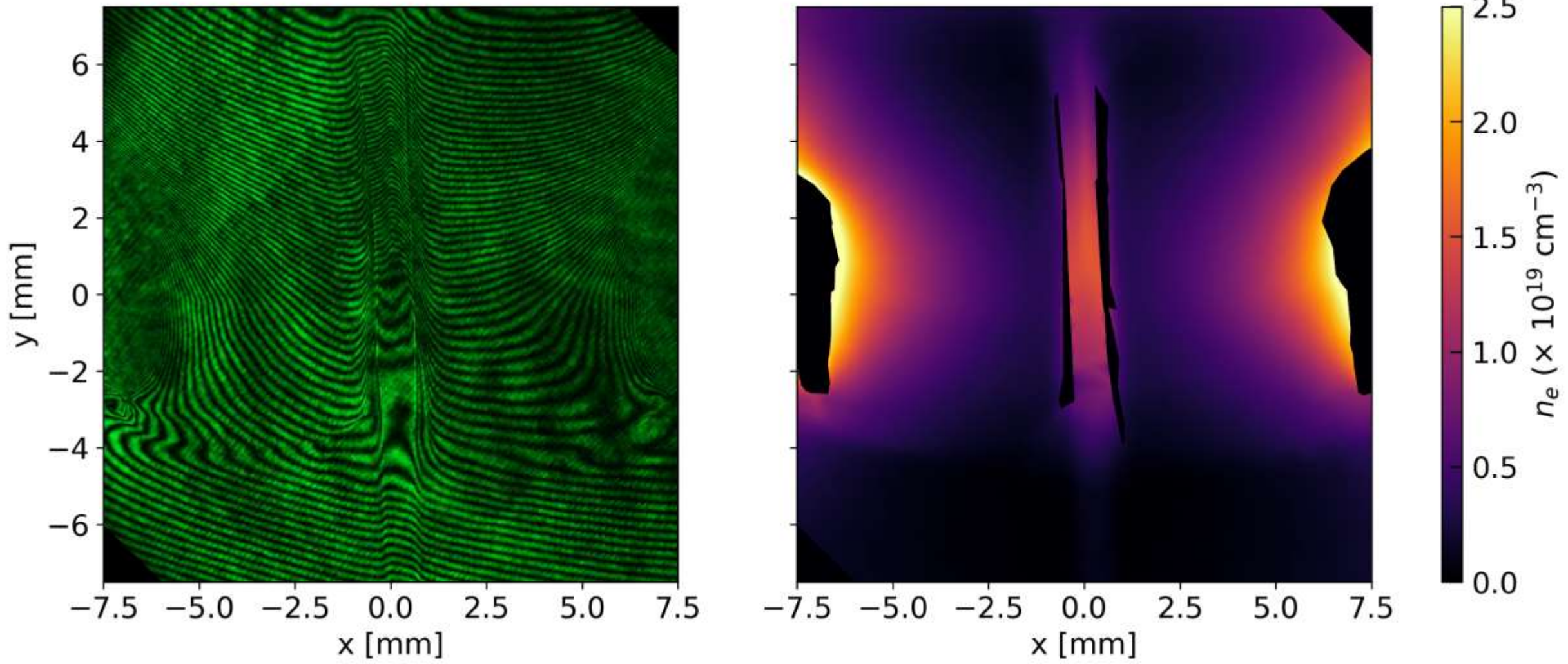
*End-on view:*



*Side-on view:*

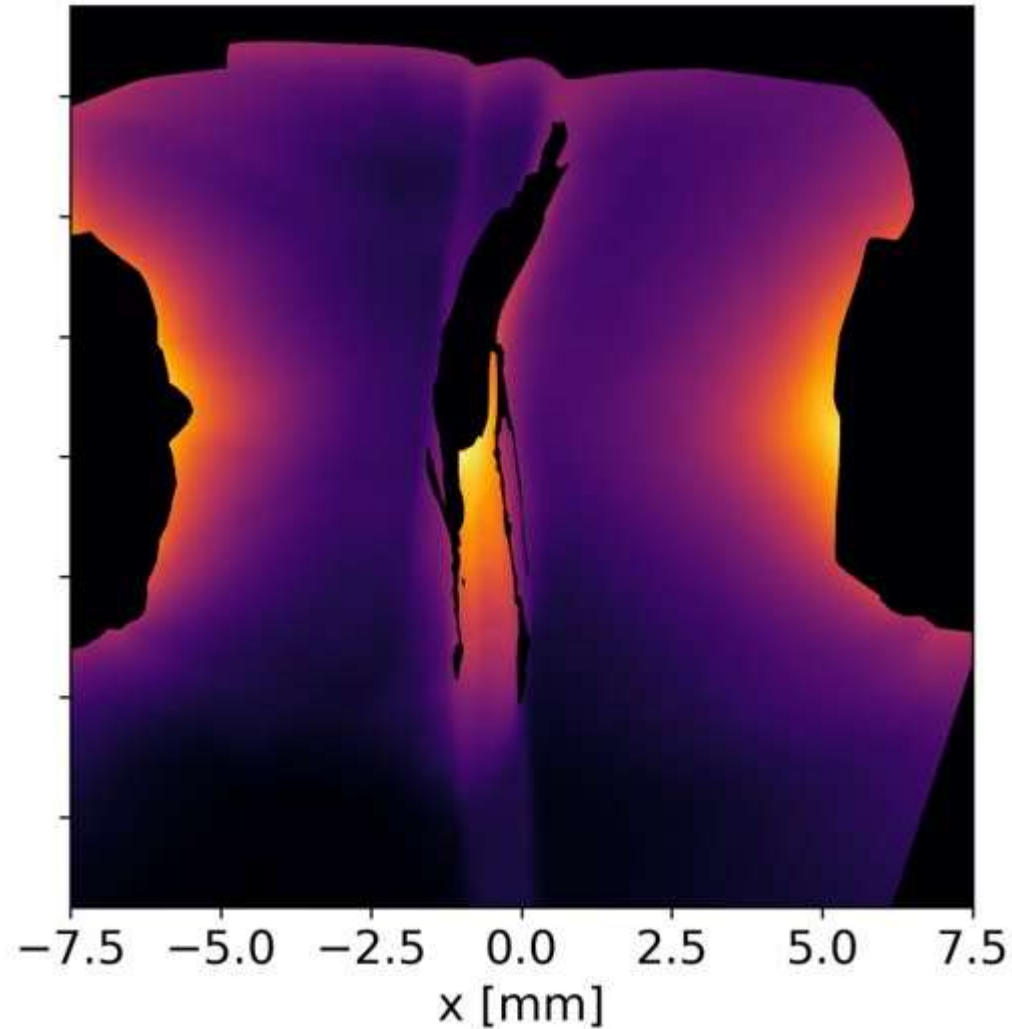


# A collimated, dense stagnation layer is formed

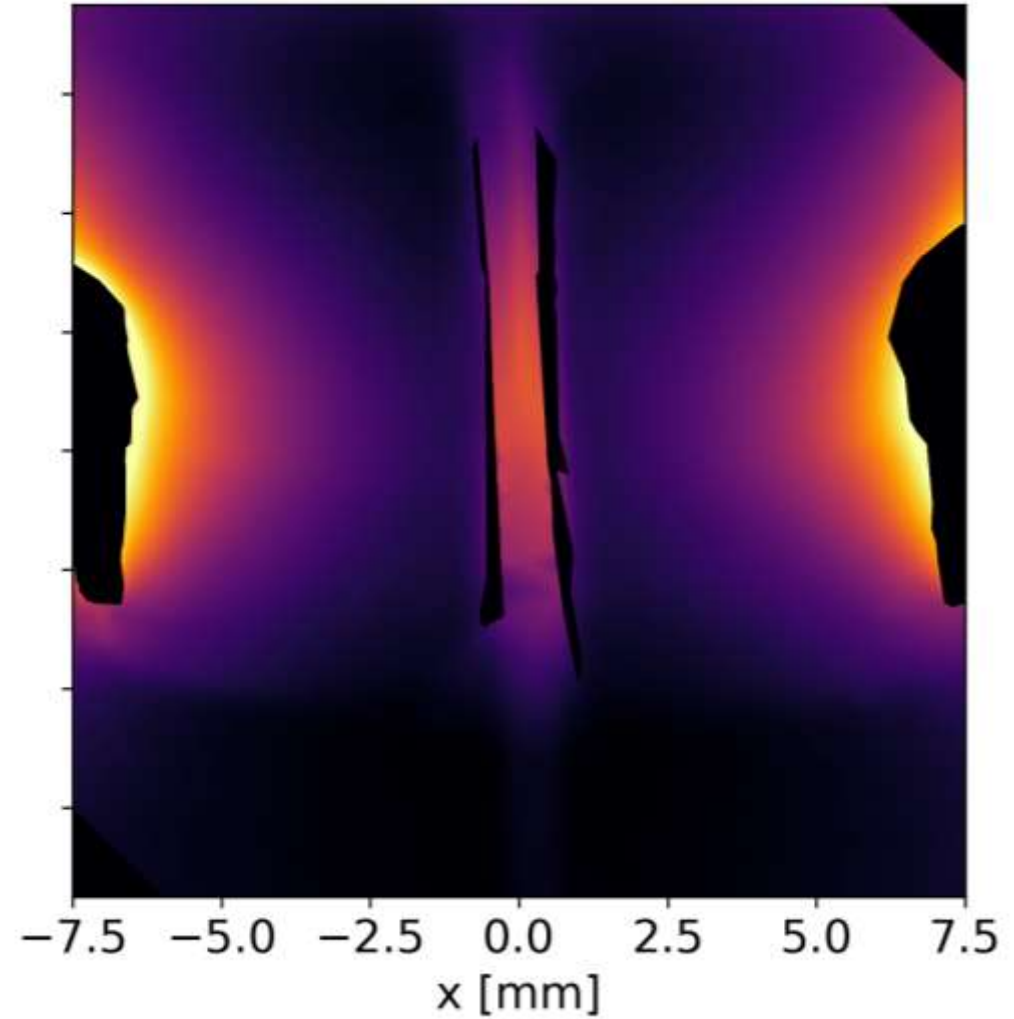


# The morphology is sensitive to plasma composition

Silicon

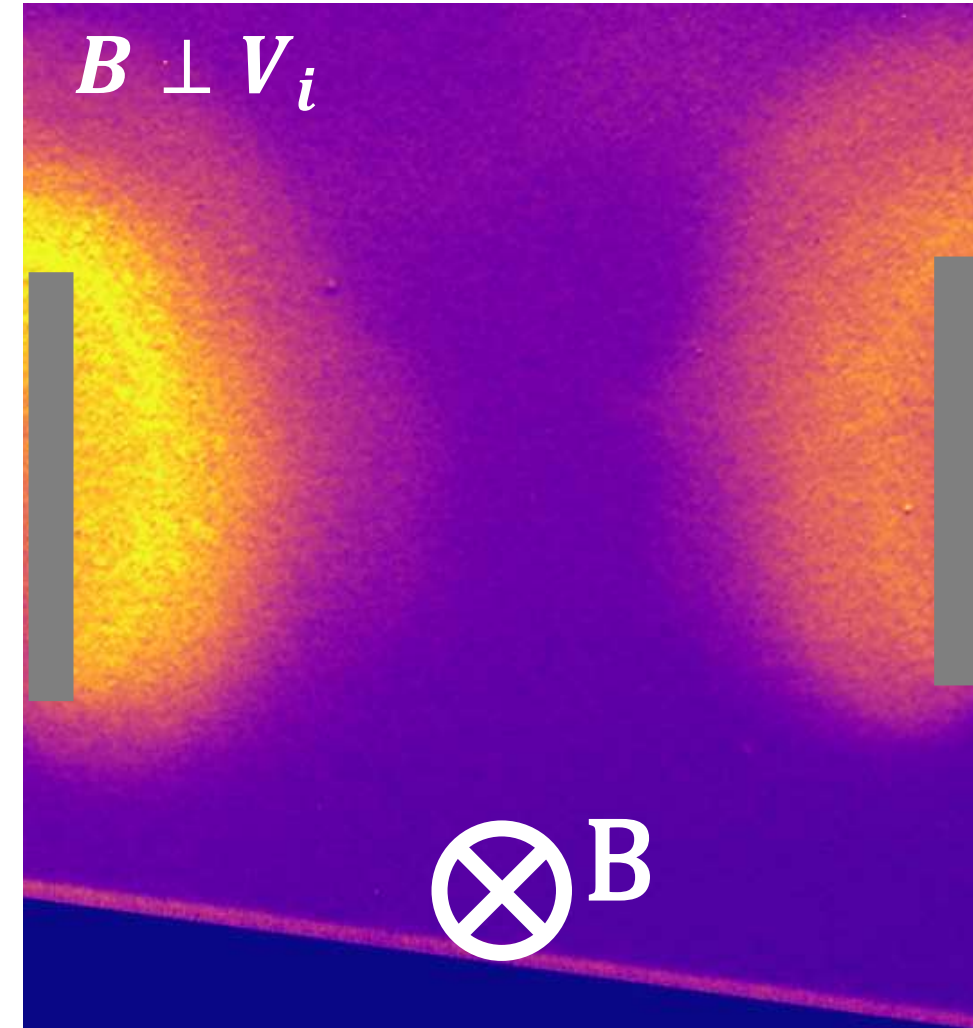
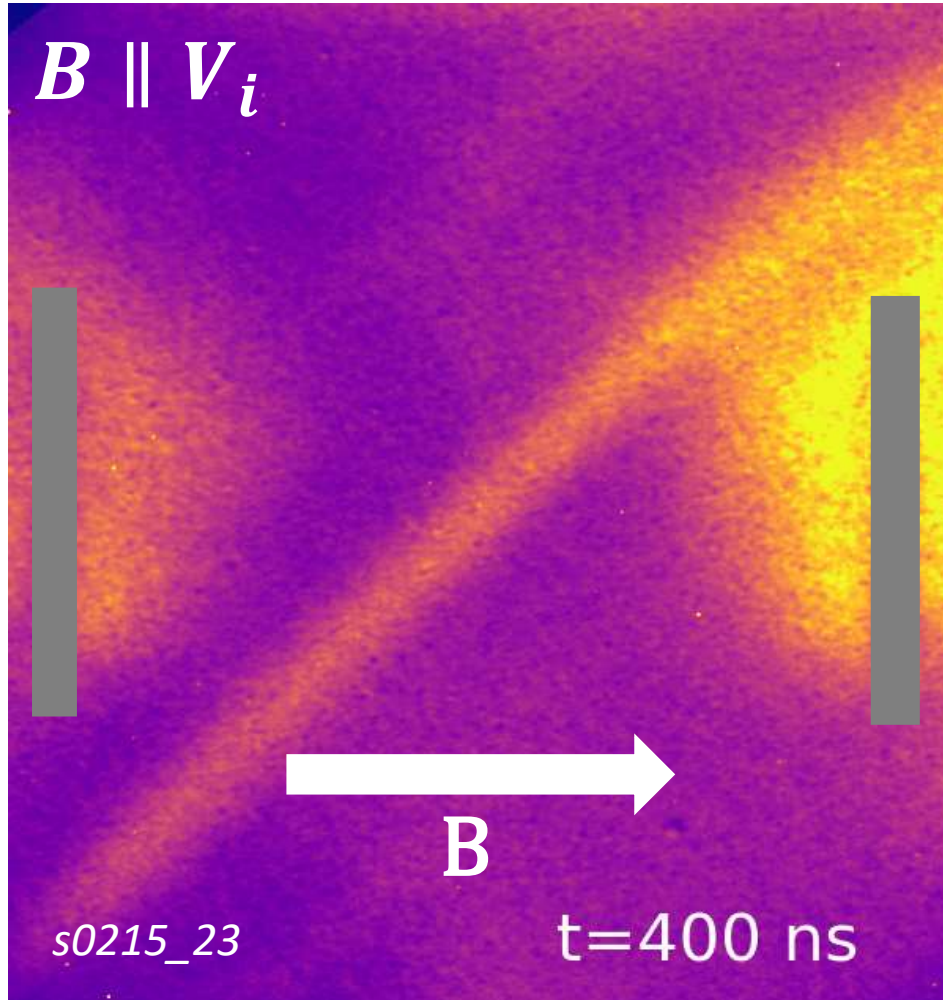


Carbon

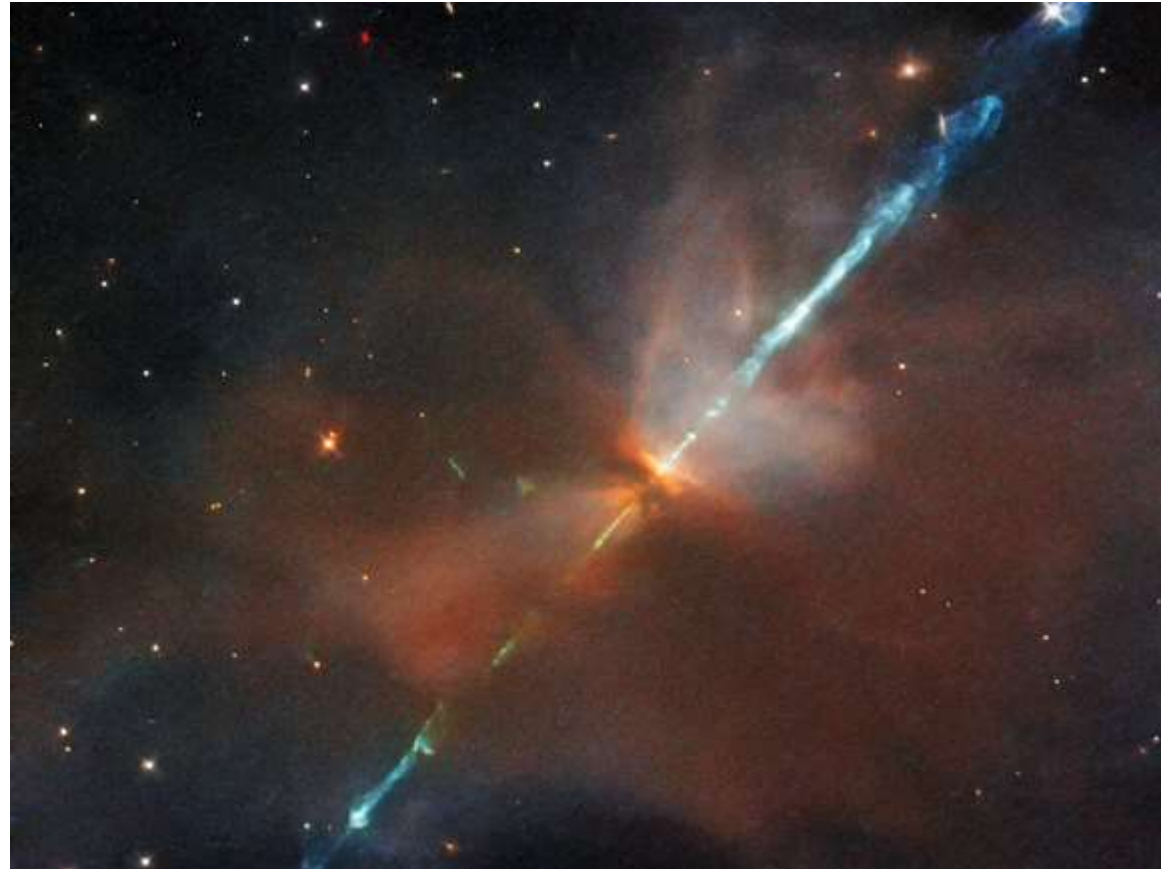




# Layer formation is suppressed by an orthogonal B field



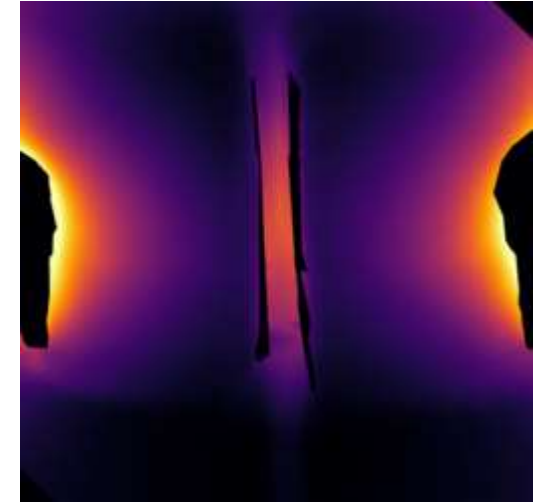
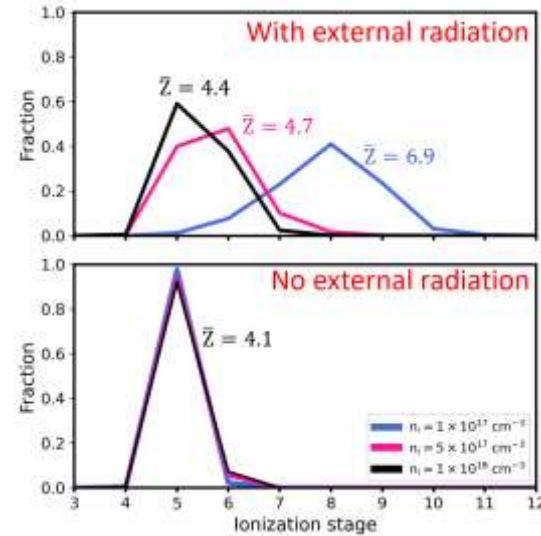
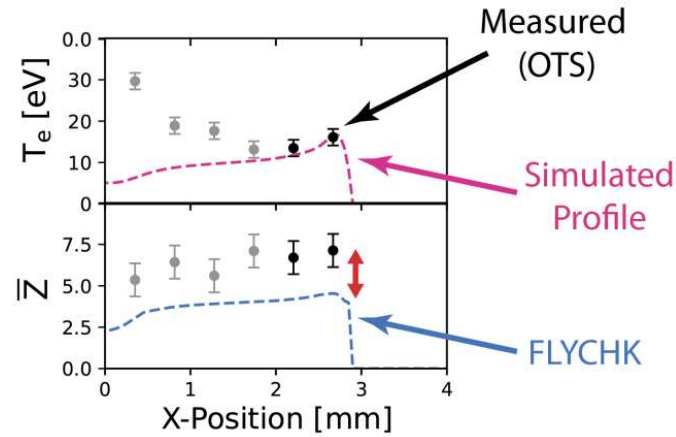
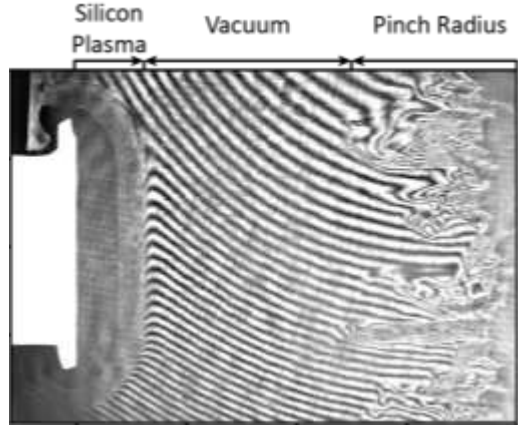
Structure in proto-stellar jets are believed to be driven by radiative instabilities.



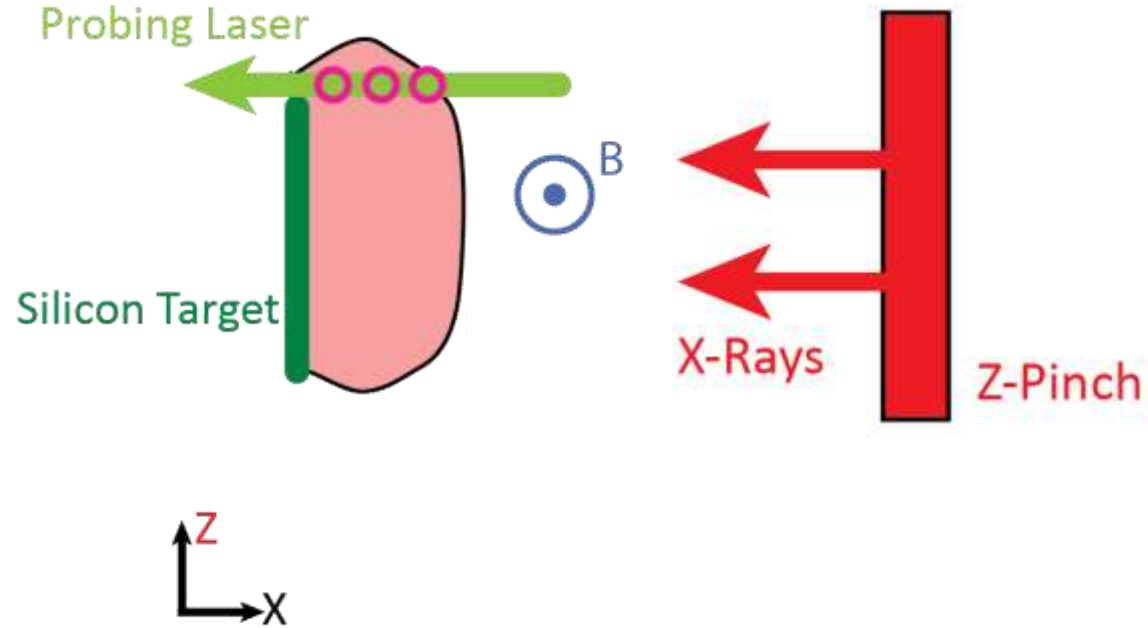
Credit ESA/NASA



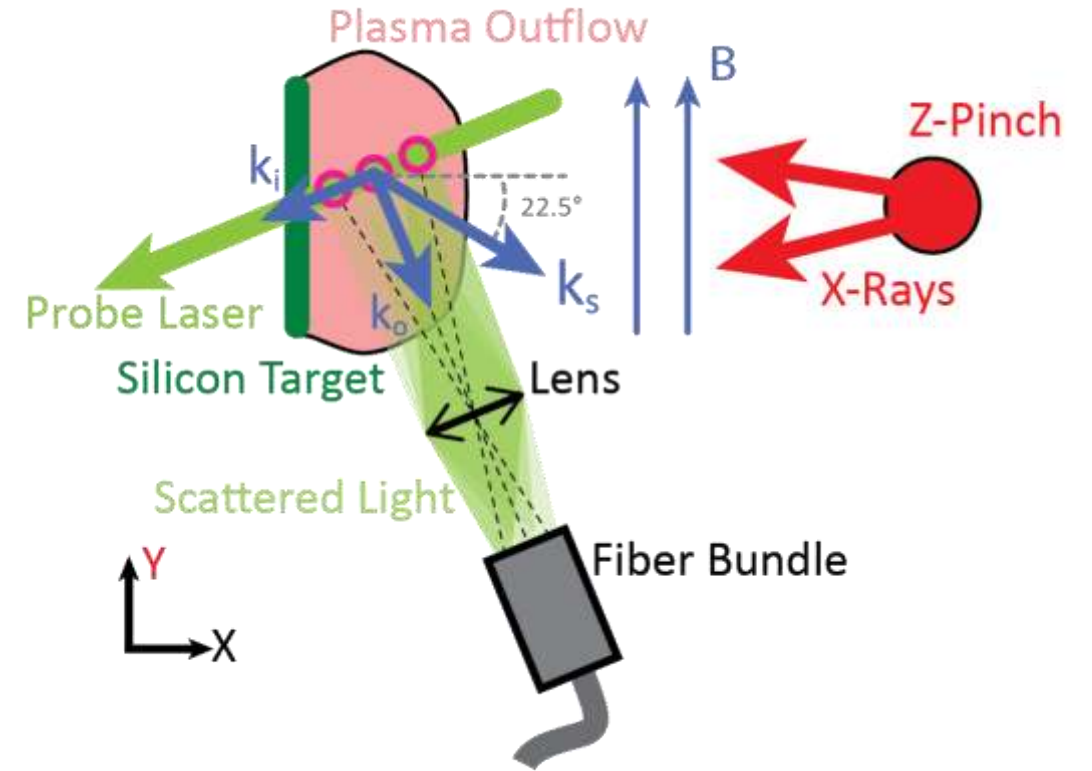
# Conclusions



- Hydrodynamic features are simple & well- characterized [Halliday+ PoP 2022]
  - Thomson measurements hint that driving radiation changes  $\bar{Z}$
- Novel testbed for atomic-kinetics models – relevant to accretion disk physics
  - Potential to study radiative instabilities relevant to stellar jets

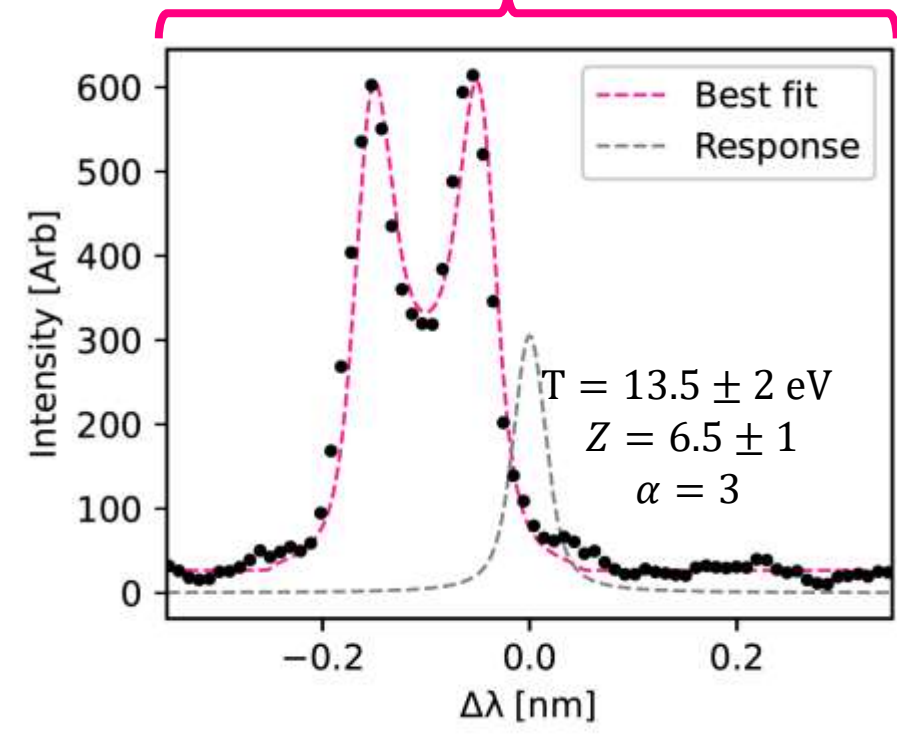
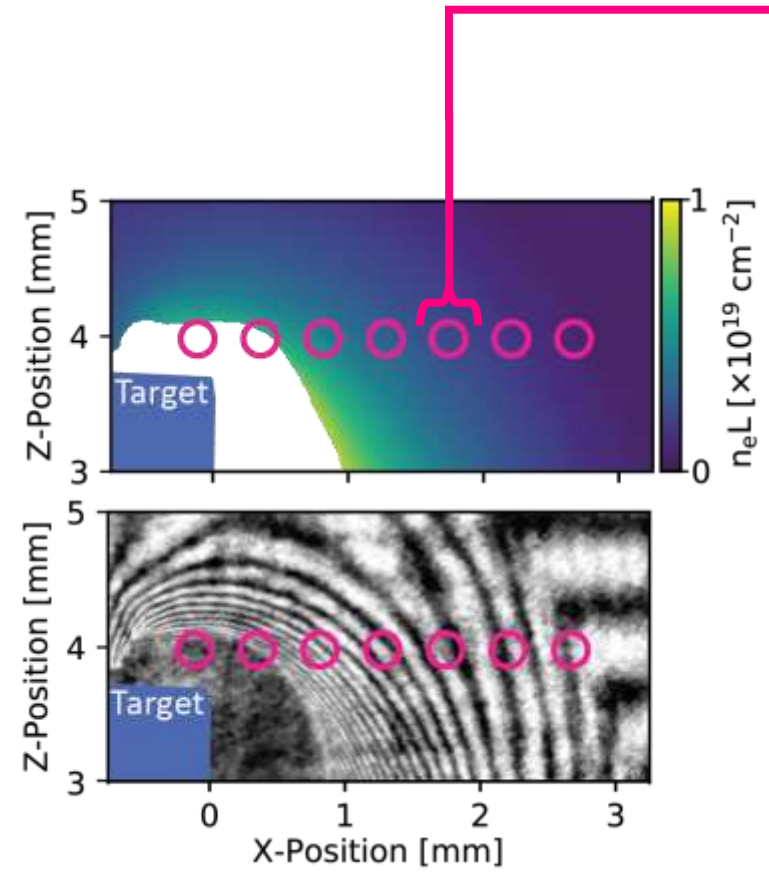
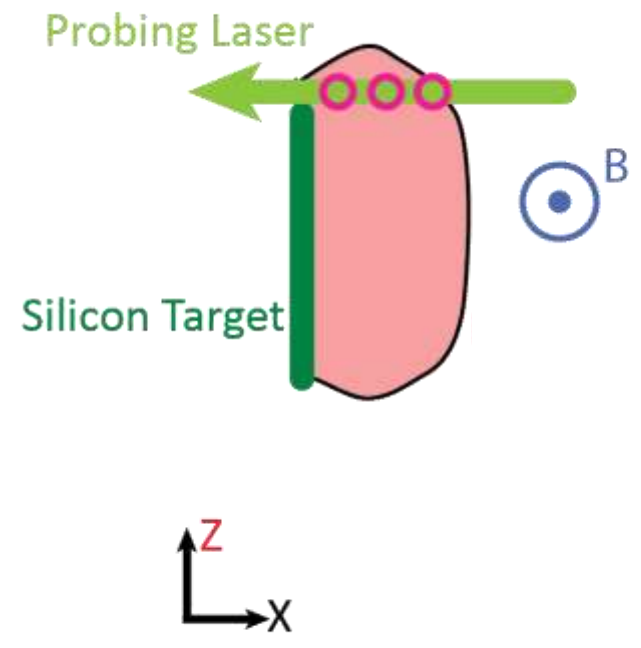


Side-On (X-Z plane) view of the experiment



End-On (X-Y plane) view of the experiment

# Thomson scattering [localised diagnosis of $T, V, Z$ ]



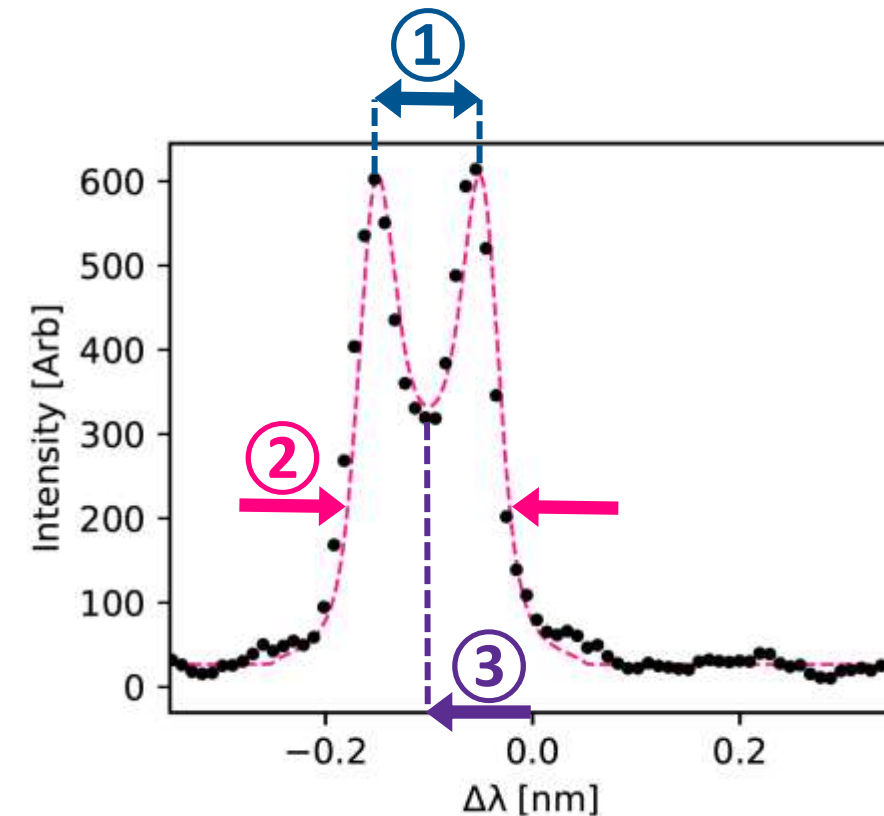
# Analysis of Ion-Acoustic Thomson Scattering Data

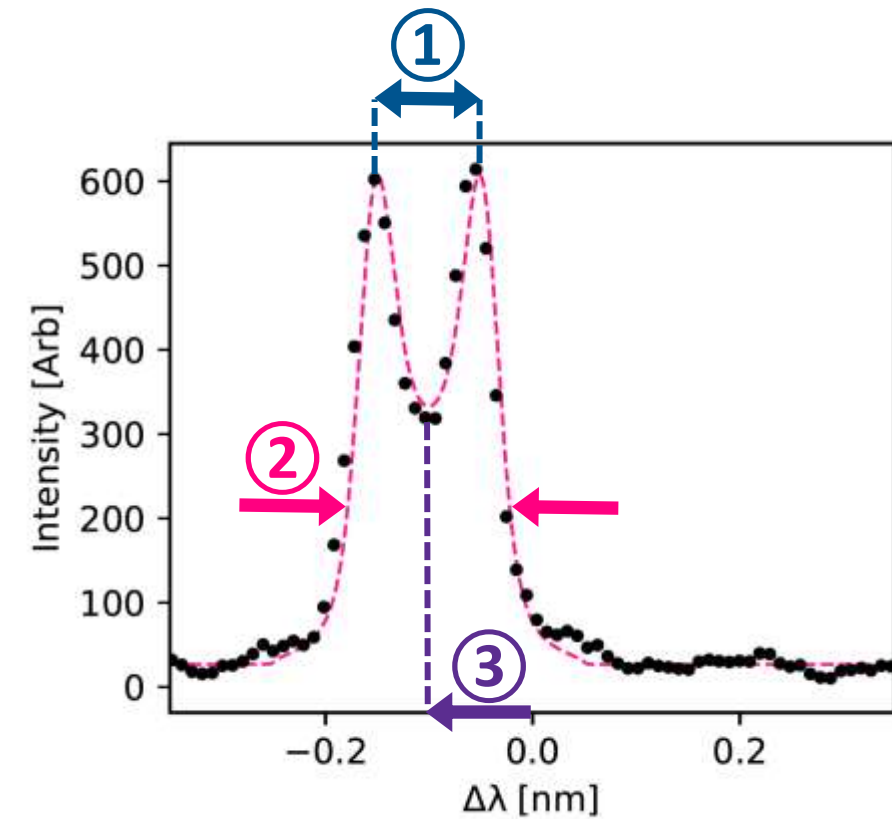
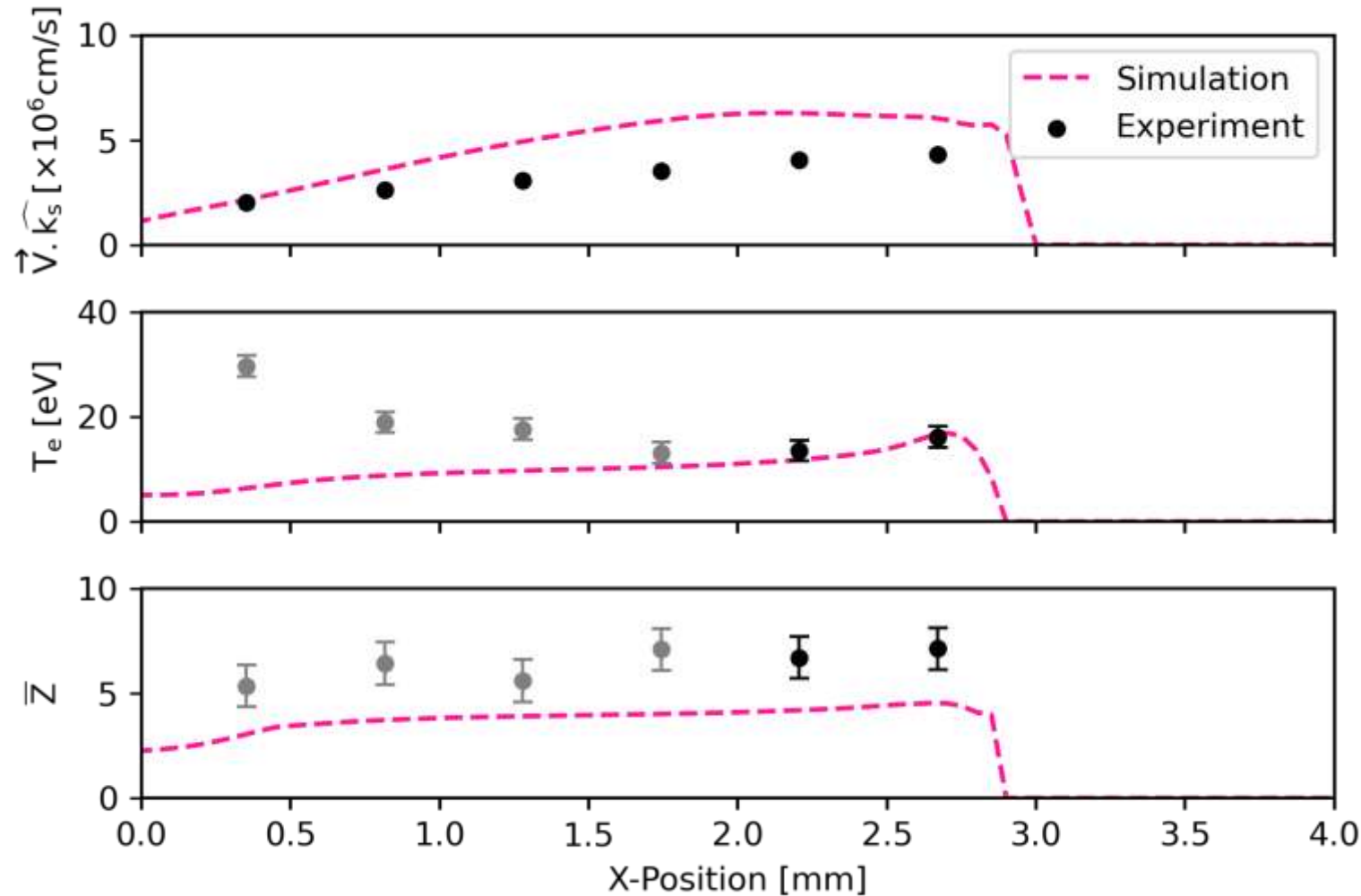
- ①: Ion Acoustic peak **separation** depends on  $\bar{Z} \times T_e$
- ②: Feature **width** depends on  $n_e$ ,  $T_i$ , and spectral response
- ③: Doppler **shift** from probe wavelength depends on  $\vec{V} \cdot \hat{k}_s$

Enforced  $T_e = T_i$ , and allowed  $\bar{Z}$  to vary ( $\tau_{ei} \lesssim 1$  ns).

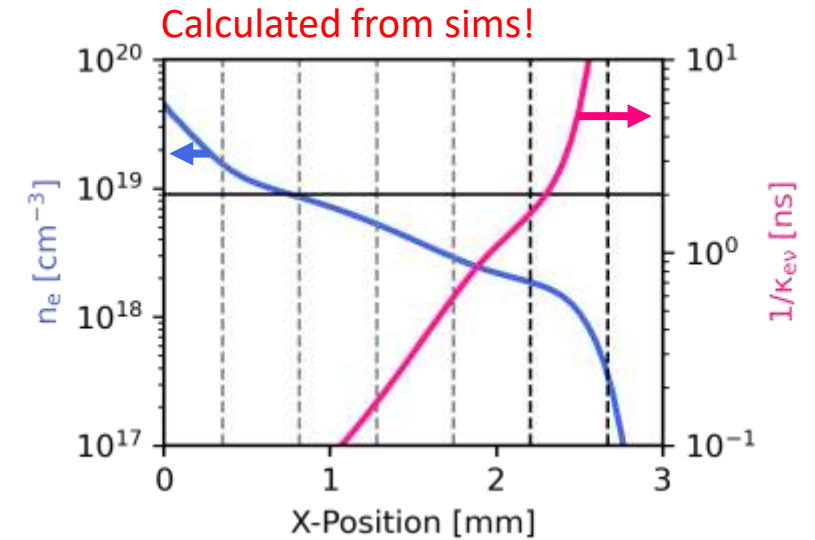
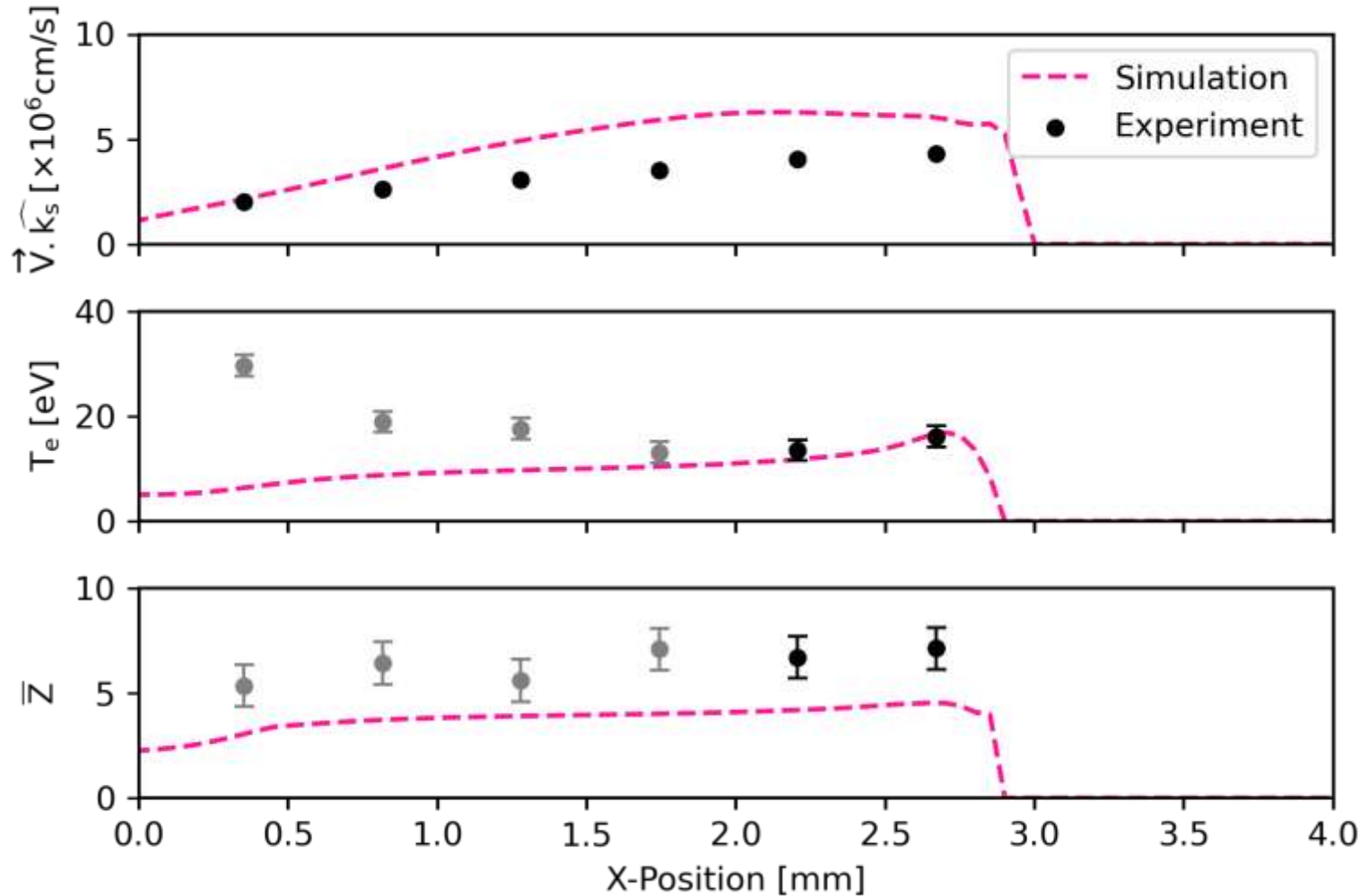
Convolved calculated spectra with measured spectral response.

Constrained value of  $n_e$  from (near simultaneous) interferometry.





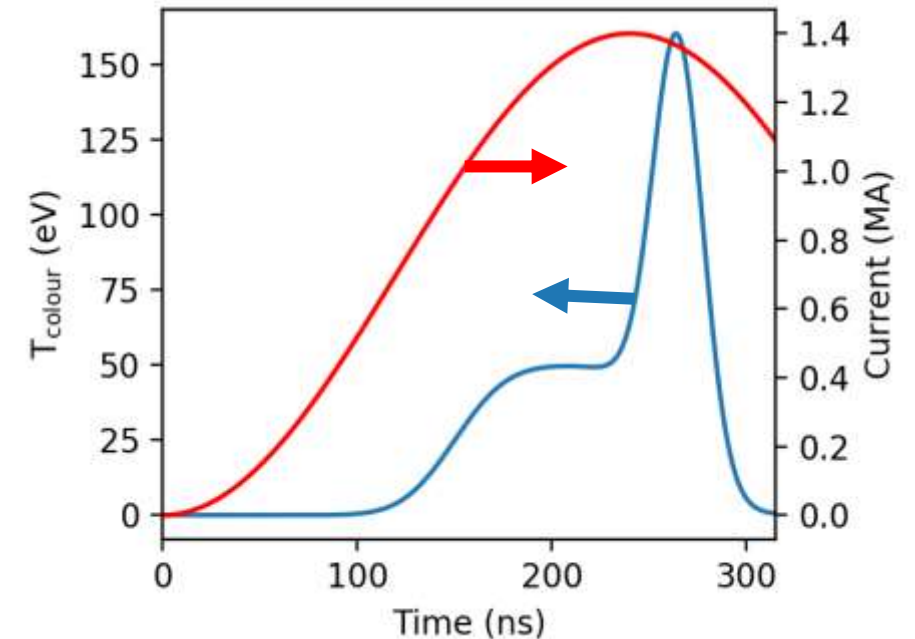




$$\kappa_{ve} \propto \frac{Zn_e^2 \ln(\Lambda) T_e^{-\frac{3}{2}}}{\sqrt{\omega^2 \left(1 - \frac{\omega_p^2}{\omega^2}\right)}}$$

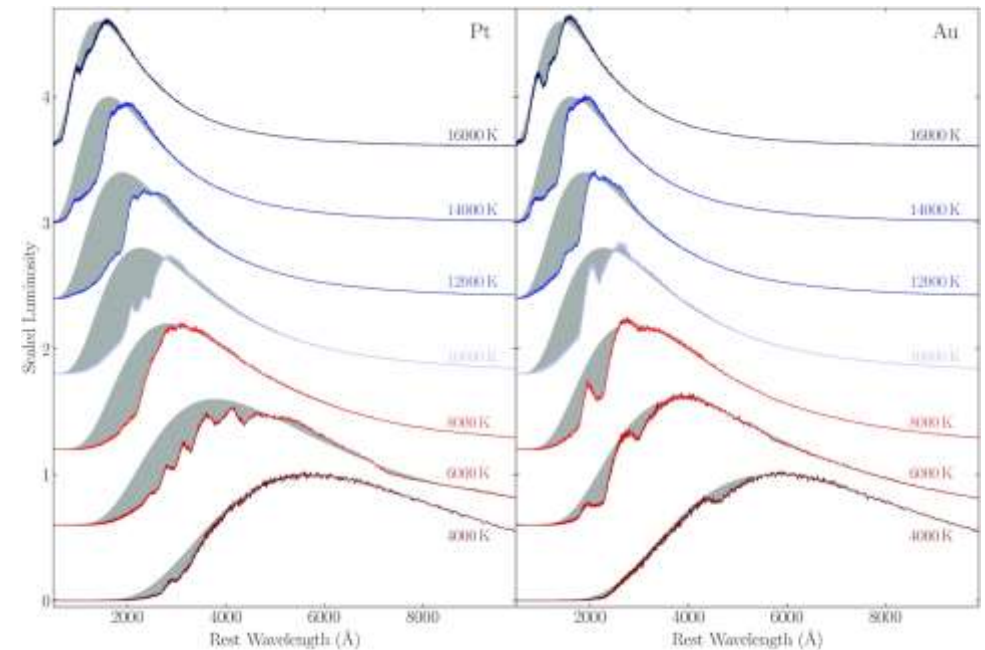
N. R. L. plasma physics formulary

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# Visible Emission Features of High Z Elements

- Binary neutron star mergers primary candidate for nucleosynthesis of high Z elements (i.e. Gold, Platinum)
- Constrain composition of outflowing material using observed spectra
- Potential to verify atomic modelling efforts [Gillanders+, MNRAS 2022] via experimental measurements
- Temperature range of interest in the range 1 – 10 eV . There are relevant spectral features in the visible.



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