Bounds on heavy axions with an X-ray free electron laser

Jack W. D. Halliday[1, 2, 3,†], G. Marocco[4], K. A. Beyer[5], C. Heaton[1], M. Nakatsutsumi[6], T. R. Preston[6], C. D. Arrowsmith[1], C. Baehtz[7], S. Goede[6], O. Humphries[6], A. Laso Garcia[7], R. Plackett[1], P. Svensson[1], G. Vacalis[1], J. Wark[1], D. Wood[1], U. Zastrau[6], R. Bingham[3, 8], I. Shipsey[1], S. Sarkar[1], and Gianluca Gregori[1]

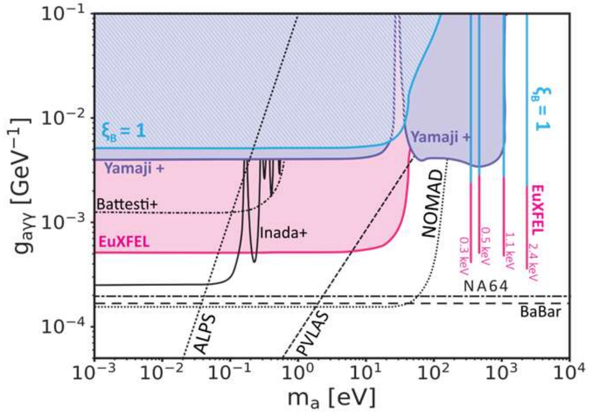
[1]University of Oxford; [2]Imperial College London; [3]STFC, Rutherford Appleton Laboratory; [4]Lawrence Berkley National Lab; [5]Max-Planck-Institut für Kernphysik Saupfercheckweg; [6]European XFEL; [7]Helmholtz Zentrum Dresden-Rossendorf; [8]University of Strathclyde

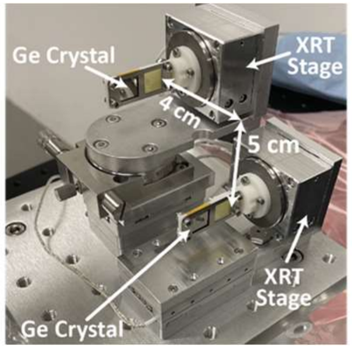
[†]jack.halliday@stfc.ac.uk – jackhalliday.github.io

This project was begun with the collaboration of Prof. Ian Shipsey – to whom this talk is dedicated

Overview

- Introduction to Axions / ALPS and overview of search strategies
- Experimental design for a light shining through a wall experiment with an X-Ray Free Electron Laser (XFEL)
- Experimental data for initial experiments on EuXFEL & expected detection threshold for future experiments
- Fundamental physics with high powered lasers





Axions solve the strong CP problem and are a cold dark matter candidate

- Postulated to explain the *absence* of CP-violation in strong interactions (which would otherwise generate an electric dipole moment for the neutron)

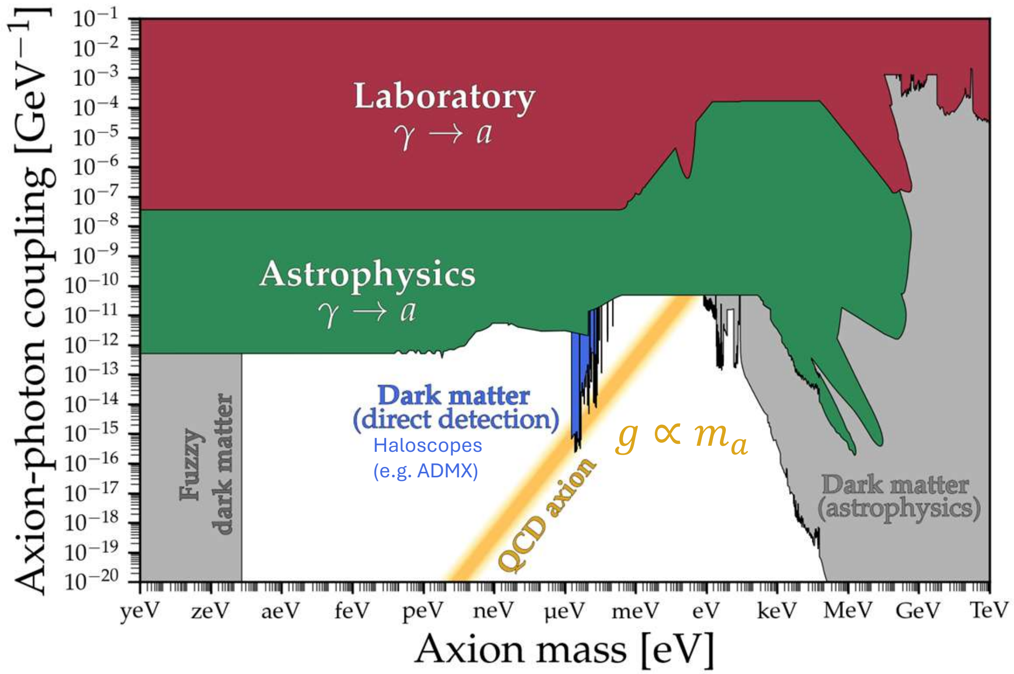
- The axion arises from spontaneous breaking of the Peccei-Quinn (PQ) symmetry

- Axion-like particles (ALPs) also arise generically in string theory

- Light axions ($m_a \sim 10^{-6} - 10^{-4}$ eV) are a natural candidate for cold dark matter

- Laboratory searches target this light axion window

F. Chadha-Day *et al*. "*Axion dark matter: What is it and why now?*" *Sci. Adv.* **8**, eabj3618 (2022).



Credit: Ciaran O'Hare (github.com/cajohare/AxionLimits)

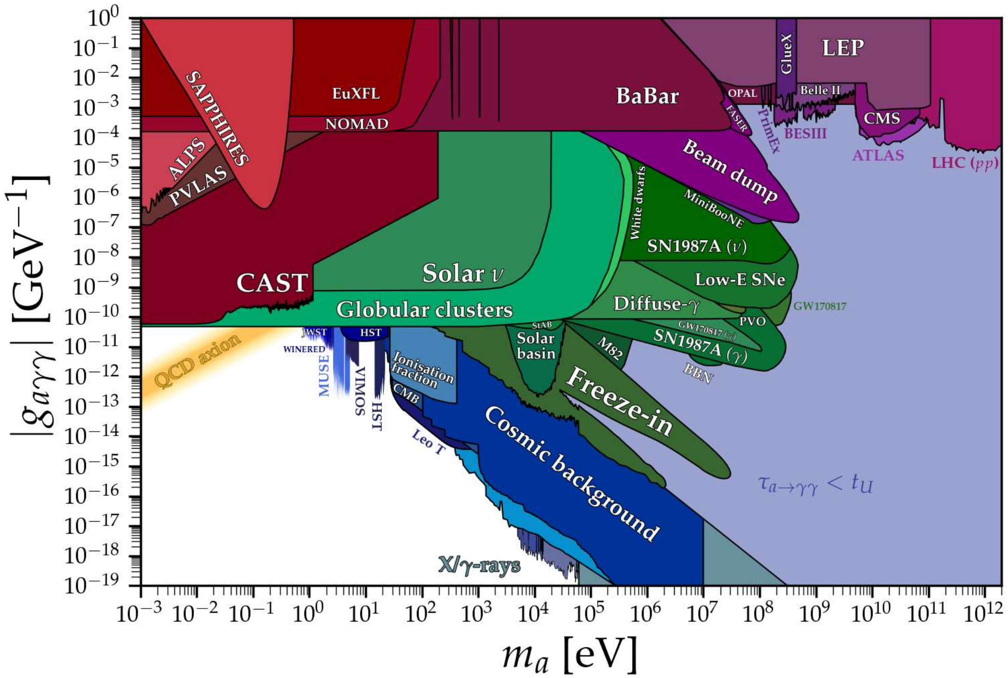
There has recently been interest in heavier axions, which motivates our study

- Recent suggestion: When PQ symmetry is broken *after* cosmological inflation, axions are also produced in the decay of axion domain walls

- Taking this additional contribution into account requires axions which make up dark matter to have a **mass above 10^{-2} eV**

A. Ringwald & K. Saikawa, "*Axion dark matter in the post-inflationary Peccei-Quinn symmetry breaking scenario*". Phys. Rev. D **93**, 085031 (2016).

K. Beyer & S. Sarkar, "*Ruling out light axions: The writing is on the wall*", SciPost Phys. **15**, 003 (2023).



Credit: Ciaran O'Hare (github.com/cajohare/AxionLimits)

There are a variety of complimentary axion search strategies

Increased sensitivity

Astrophysical

Axion losses in stellar models

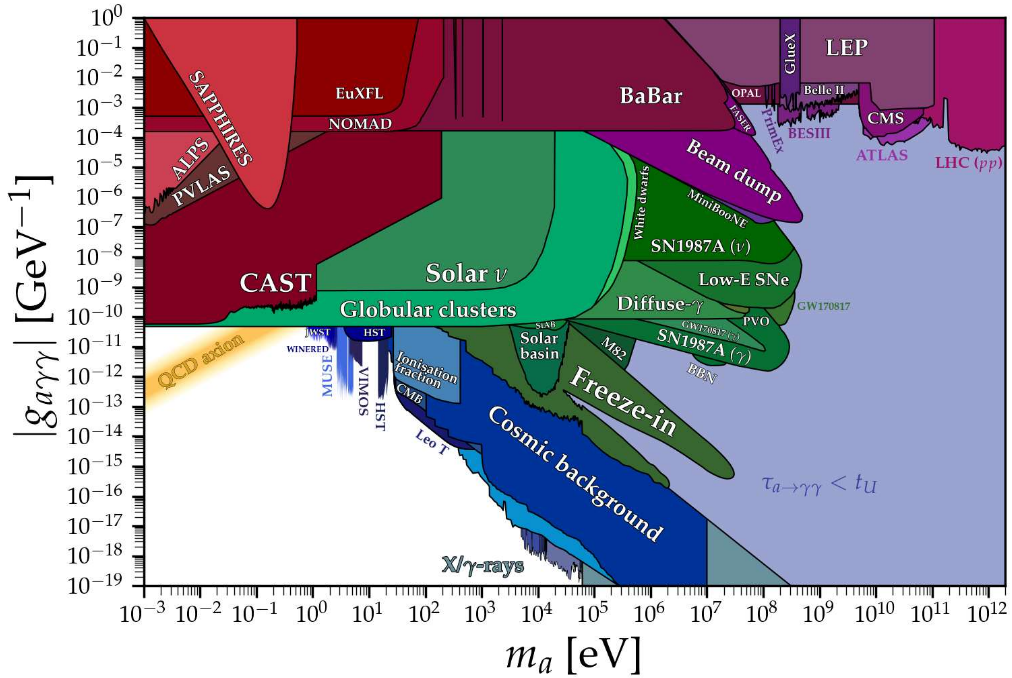
Lab detection

Helio/haloscopes

Lab generation

Generate (& detect) terrestrially

Reduced model dependence



Credit: Ciaran O'Hare (github.com/cajohare/AxionLimits)

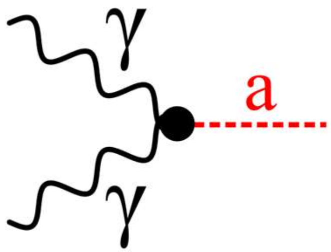
Yannis K. Semertzidis & SungWoo Youn, "*Axion dark matter: How to see it?*", Sci. Adv. **8**, eabm9928 (2022).

A. Caputo & G. Raffelt, *"Astrophysical Axion Bounds: The 2024 Edition"*, Proceedings of Science **454**, 041 (2024).

For $m_a \leq 1$ eV, stringent bounds are imposed by the CERN Axion Telescope (CAST)

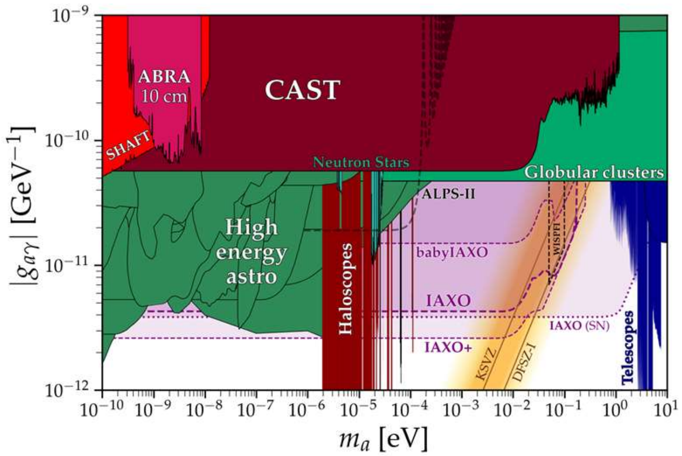
- Helioscope: Axions produced in the sun convert into photons via the Primakoff effect:

$$\mathcal{L}_{\text{axion}} = g_{a\gamma\gamma}\boldsymbol{E}\cdot\boldsymbol{B}a$$



- Geometry suppresses sensitivity to axions with $m_a > 1$ eV

- Some model dependence due to high temperature and ω_p in Solar plasma



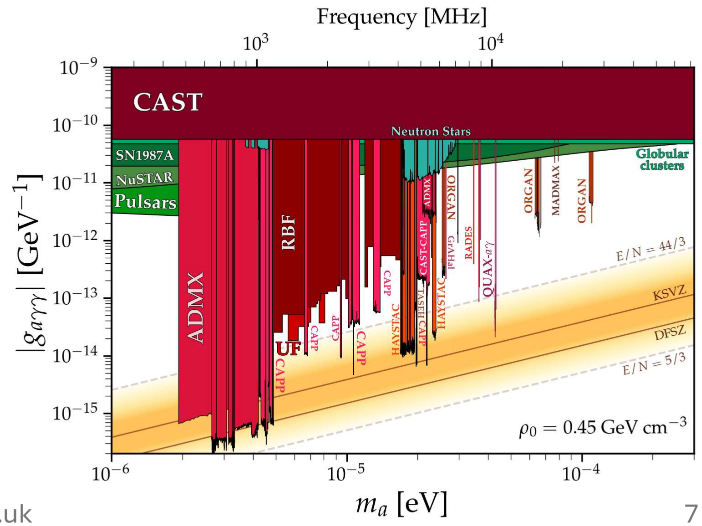
CAST Collaboration, "*New CAST limit on the axion–photon interaction*" Nature Phys. **13**, 584 (2017).

J. Jaeckel *et al*. "*Need for purely laboratory-based axionlike particle searches",* Phys. Rev. D **75,** 013004 (2007)

In the $m_a \sim \mu\text{eV}$ region, haloscopes, including ADMX, probe down to the QCD axion coupling.

- Axion Dark Matter eXperiment (ADMX), at the University of Washington
- Haloscope: Axions from (local) Galactic halo convert to microwaves in cavity surrounded by a super-conducting electromagnet ($\text{B} = 7.6$ T)
- Tune cavity Q-factor to scan through different axion masses
- Probes down to QCD axion couplings for $m_a = 1 - 100$ μeV

C. Goodman *et al*, "*Axion dark matter experiment around 3.3 eV with Dine-Fischler-Srednicki-Zhitnitsky discovery ability*", Phys. Rev. Lett. – in press (2025)



There are a variety of complimentary axion search strategies

Increased sensitivity

Astrophysical

Axion losses in stellar models

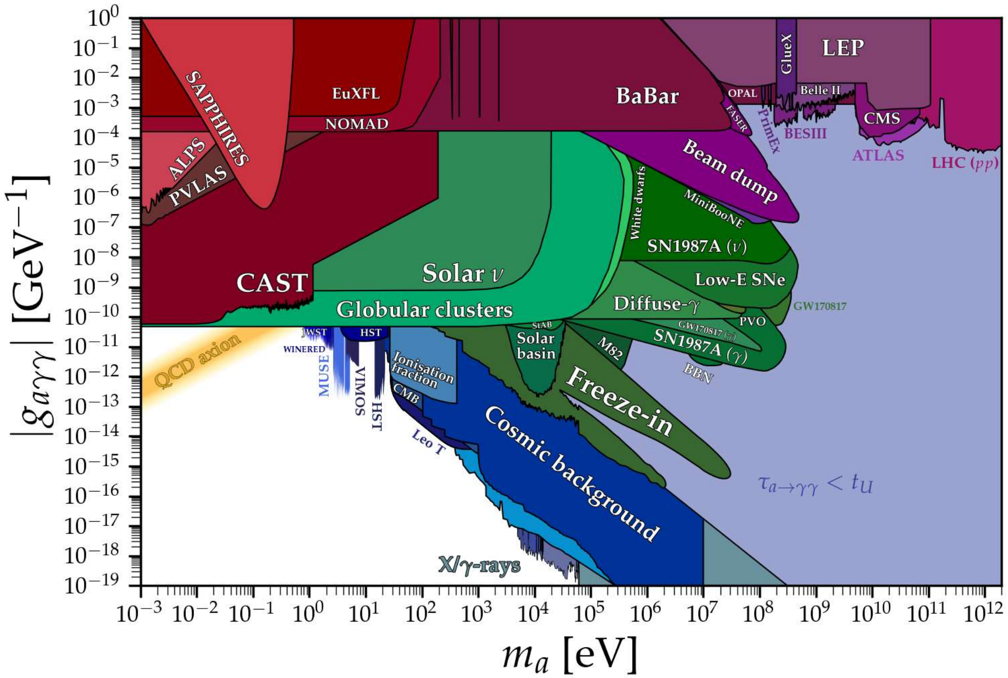
Lab detection

Helio/haloscopes

Lab generation

Generate (& detect) terrestrially

Reduced model dependence



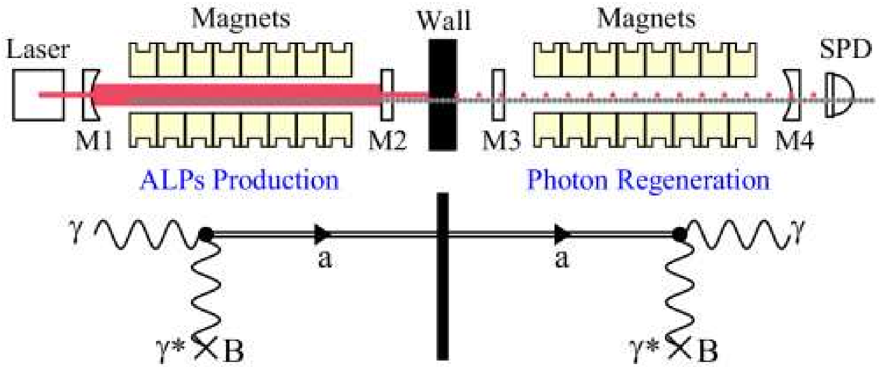
Yannis K. Semertzidis & SungWoo Youn, "*Axion dark matter: How to see it?*", Sci. Adv. **8**, eabm9928 (2022).

Credit: Ciaran O'Hare (github.com/cajohare/AxionLimits)

A. Caputo & G. Raffelt, *"Astrophysical Axion Bounds: The 2024 Edition",* Proceedings of Science **454**, 041 (2024).

Light shining through wall (LSW) experiments are an example of direct detection searches

- Generating axions in the lab removes any model dependence
- But sensitivity suppressed by a factor of $\sim g_{a\gamma\gamma}$
- No sensitivity to $m_a > 1\ \mathrm{eV}$ (uses visible photons)
- Photo taken from ALPS-II, current experiment at DESY which recycles magnets from HERA



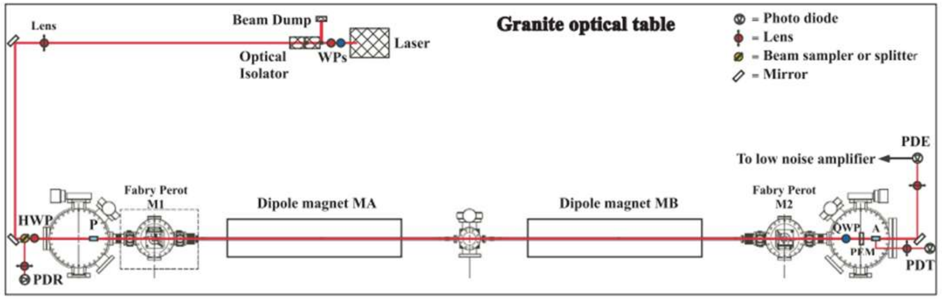
C. Robilliard *et al, "No light shining through a wall",* Phys Rev. Lett. **99**, 190403 (2007).

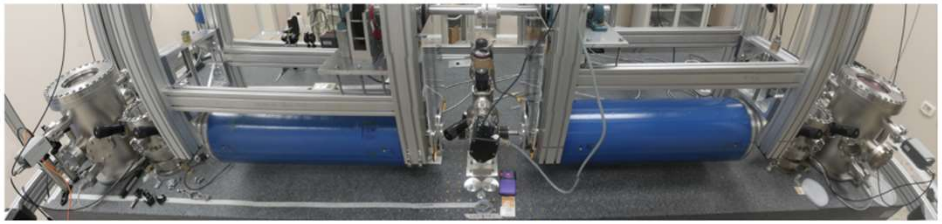
Klaus Ehret *et al, "New ALPS results on hidden-sector lightweights",* Phys. Lett. B **689**, 149 (2010).

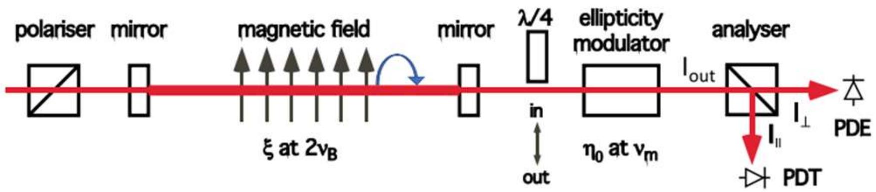
QSQAR collaboration*, "New exclusion limits on scalar and pseudoscalar axionlike particles"* Phys. Rev. D**92**, 092002 (2015).

Vacuum birefringence experiments, including PVLAS, are also sensitive to axions generated terrestrially

- Vacuum birefringence – induce ellipticity in laser propagating though magnetic field
- Sensitive to a variety of QED effects (e.g. photon-photon scattering)
- Axion-photon coupling can also induce ellipticity
- Sensitive to <u>generation</u> of axions, *not* <u>regeneration</u> of photons:
    - $P_{a\leftrightarrow\gamma} \propto (g_{a\gamma\gamma})^{1}$ – Vacuum birefringence
    - $P_{a\leftrightarrow\gamma} \propto (g_{a\gamma\gamma})^{2}$ – LSW experiments
- Hard limit on sensitivity due to QED effects



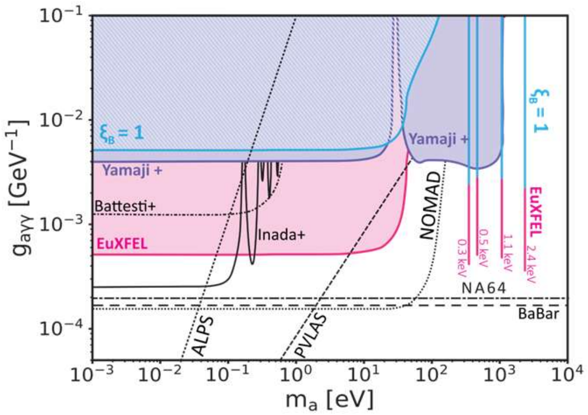


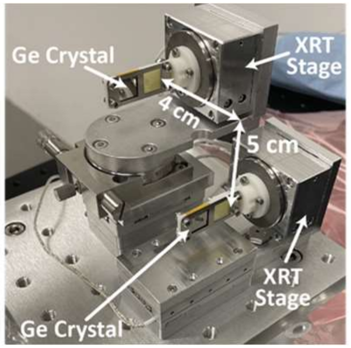


A. Ejlli *et al*, "*The PVLAS experiment: A 25 year effort to measure vacuum magnetic birefringence*" Phys. Rep. **871**, 1 (2020).

Overview

- Introduction to Axions / ALPS and overview of search strategies
- **Experimental design for a light shining through a wall experiment with an X-Ray Free Electron Laser (XFEL)**
- Experimental data for initial experiments on EuXFEL & expected detection threshold for future experiments
- Fundamental physics with high powered lasers



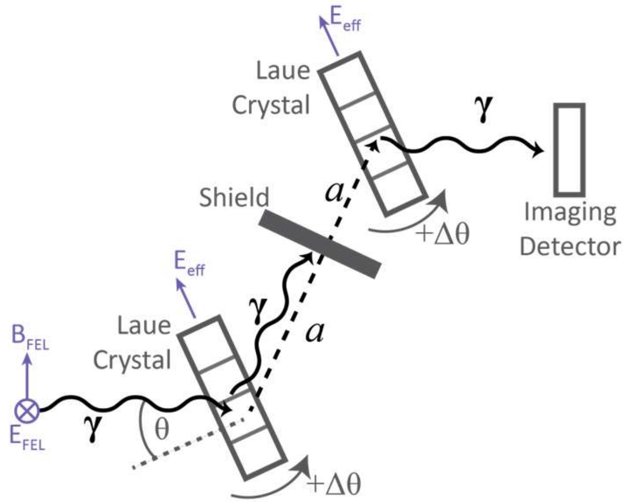


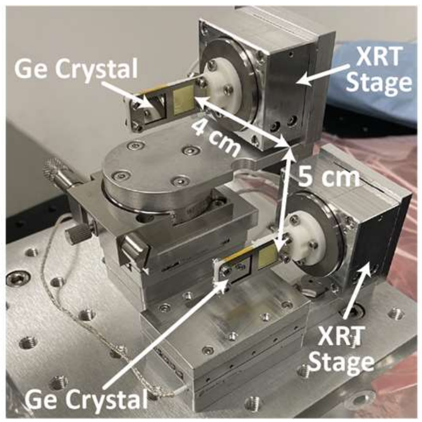
X-ray light sources enable the extension of light shining through wall experiments to search for heavy axions

- Again, rely on Primakoff effect:

$$\mathcal{L}_{\text{axion}} = g_{a\gamma\gamma}\boldsymbol{E}\cdot\boldsymbol{B}a$$

- Magnetic field from (σ-polarised) X-Ray beam
- Electric field normal to planes in crystalline material:
  - $E{\sim}10^{11}$ V/m (equivalent to $B{\sim}1$ kT)
  - Germanium crystal, thickness $L = 500$ μm
  - Path integrated field ~ 25 Tm
- Previous experiments [Yamaji+] on Synchrotron light source
- Our study is the *first* to use an X-Ray Free Electron Laser



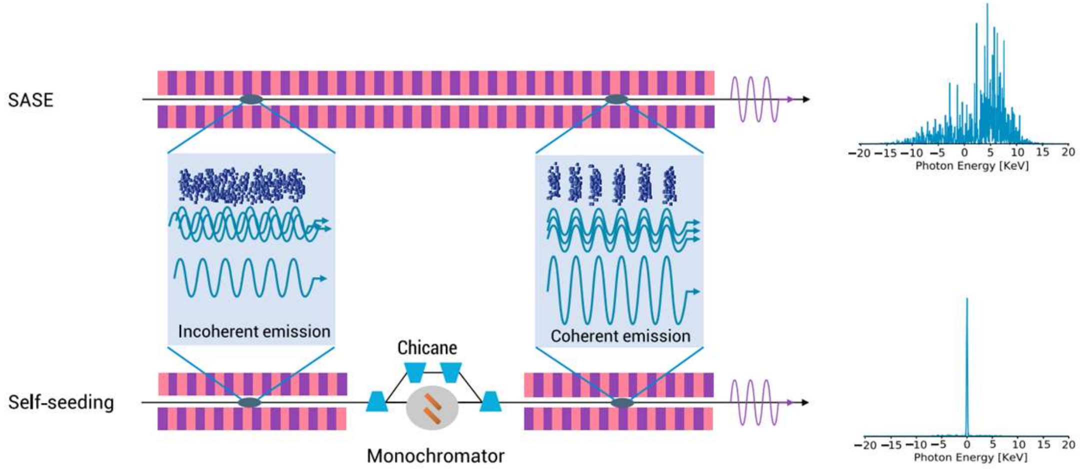


Buchmüller & Hoogevee, *"Coherent production of light scalar or pseudoscalar particles in Bragg scattering",* Phys. Lett. B **267**, 2 (1990).

T. Yamaji *et al, "Search for Axion like particles using Laue-case conversion in a single crystal"* Phys. Lett. B **782**, 526 (2018).

An X-ray Free Electron Laser (XFEL) is a light source, based on an electron LINAC

- LINAC accelerates electrons to $\sim 10\ \mathrm{GeV}$
- Undulators (periodic dipole magnets) cause beam to oscillate → radiation emitted
- **SASE** (Self Amplified Spontaneous Emission):
  - Emitted radiation interacts with beam causing microbunching
  - Microbunching amplifies light (coherent emission)
  - Very short pulse-duration ($\sim 10\ \mathrm{fs}$)
- **Self-seeding –** monochromator in beamline imposes preferred length scale for microbunches → monochromatic light
- Extremely high "brilliance"
  $(N_\gamma/\ \mathrm{s}\ /\mathrm{mm}^2/\ \mathrm{mrad}^2/\ 0.1\%\mathrm{BW})$
- High impact science from drug discovery to fundamental physics with access decided in extremely competitive calls



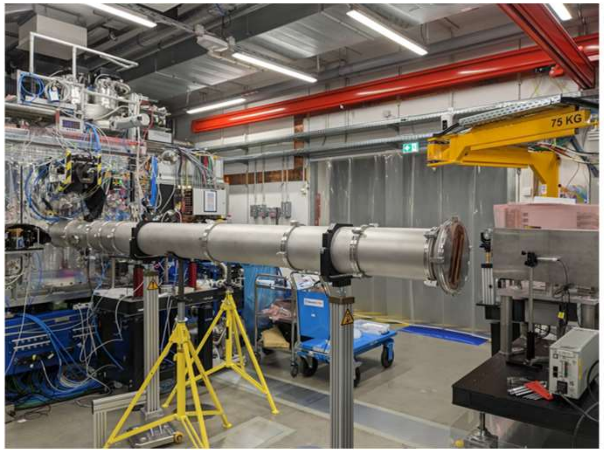
Facility	Beam energy [GeV]	Photon energy [eV]	Pulse duration [fs]
EuXFEL	8.5 – 17.5	240 – 25,000	3-150
LCLS-II	4 – 15	200 – 25,000	1 – 500
SACLA	5.1 – 8.9	400 – 12,800	2 – 10

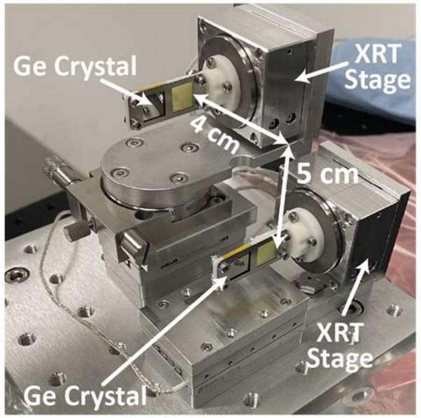
N. Huang *et al.* *"Features and futures of X-ray free-electron lasers"*. The Innovation. **2**, 2 (2021)

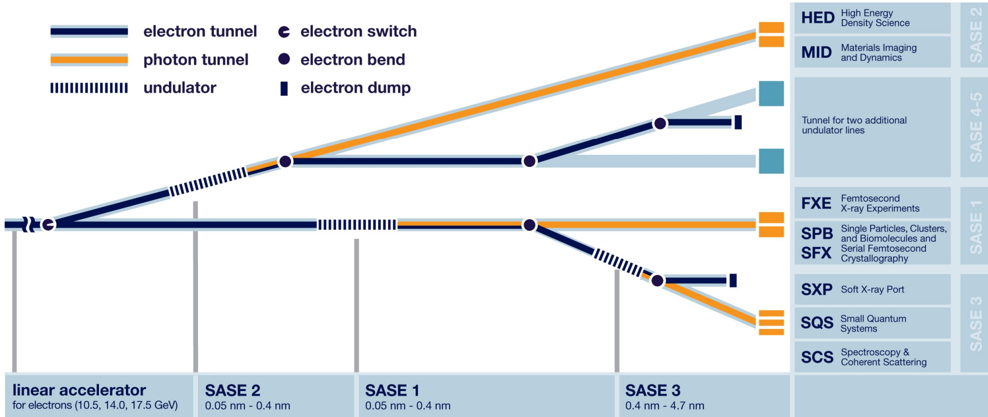
Configuration used for our experiment at The European XFEL (EuXFEL)

- Our search used the HED/HiBEF instrument at EuXFEL, Hamburg
- Three days (72 hrs) of beamtime awarded over the Easter bank holiday 2023
- Operated in a self-seeded mode

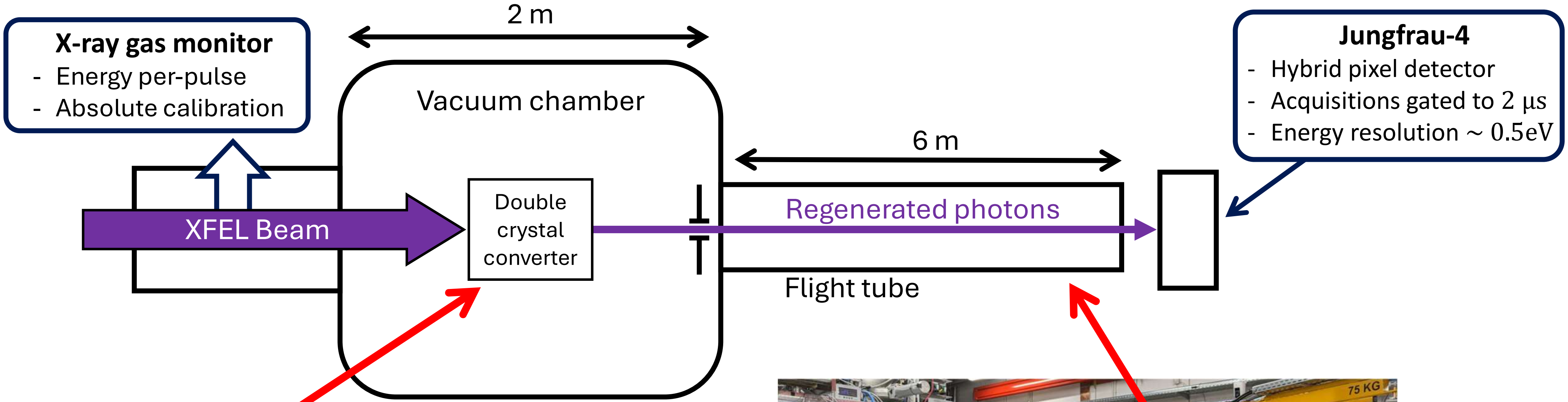
Parameter	Value
Photon energy / wavelength	9.8 keV / 1.3 Å
Bandwidth ($\Delta E/E$)	5×10^{-5}
Pulse duration	$\sim 10^{-1}$ s
Repetition rate	10 Hz
Spot size	$400 \times 400\ \mu m^2$
Total photons on-target	1×10^{17}

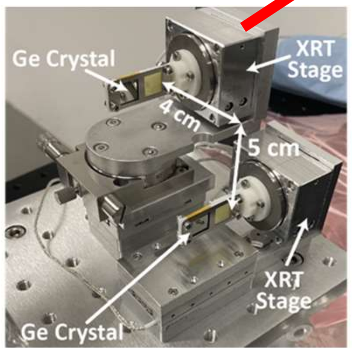


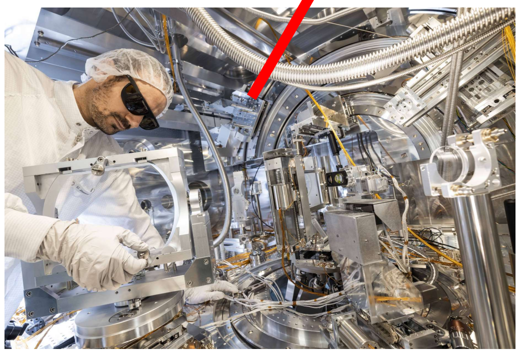


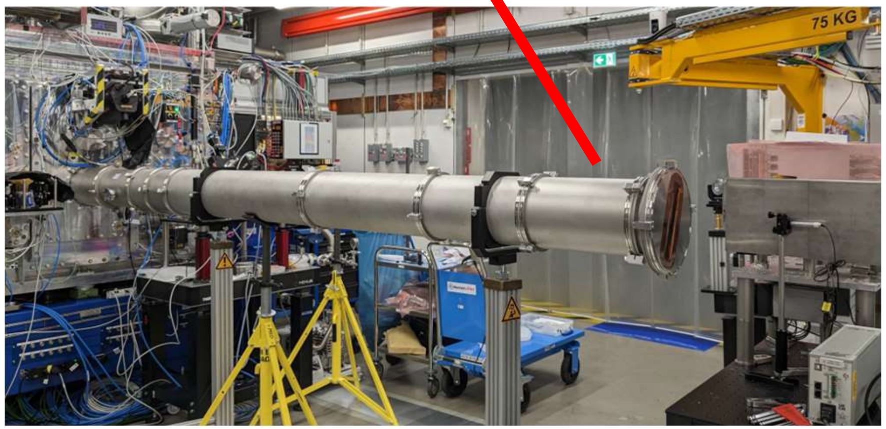


The HED/HiBEF instrumentation enabled an accurate constraint to be placed on input / output X-ray flux









The sensitivity of the scan is a function of the interaction-length, while the mass range is tuned by adjusting θ

Probability of conversion:

$$P(a \leftrightarrow \gamma) = \left(\frac{1}{4} g_{a\gamma\gamma} E_{\text{eff}}\, L_{\text{eff}} \cos\theta_B\right)^2$$

$$L_{\text{eff}} = 2L_{\text{att}}^B \left[1 - \exp\left(-\frac{L_x}{2L_{att}^B}\right)\right], \quad L_x = \ell/\cos(\theta_B + \Delta\theta)$$

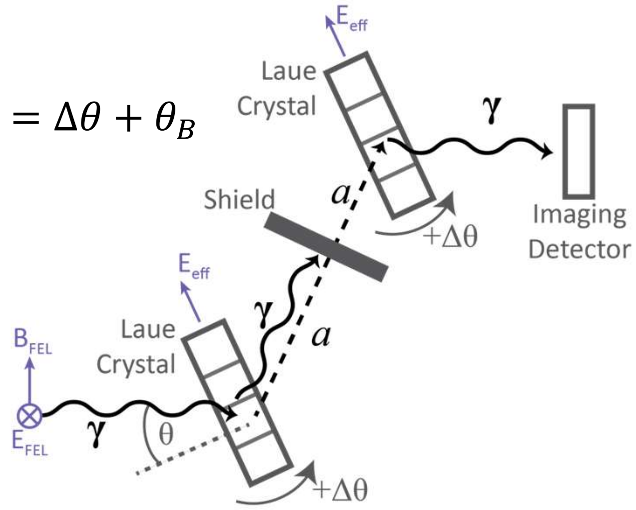
Mass range for the search:

$$|m_a^2 - m_\gamma^2| \le \frac{4k_\gamma}{L_{eff}} \qquad (\Delta\theta = 0)$$

$$m_a = \sqrt{m_\gamma^2 + 2q_T k_\gamma \cos(\theta_B)\Delta\theta} \qquad (\Delta\theta \neq 0)$$

T. Yamaji *et al, "Theoretical calculation of coherent Laue-case conversion between x-rays and ALPS for an X-ray LSW experiment"* Phys. Rev. D **96**, 15001 (2017).

$$\theta = \Delta\theta + \theta_B$$



Reciprocal lattice spacing (q_T)	6.2 keV
Plasma frequency (m_γ)	44 eV
Crystalline field (E_{eff})	7.3×10^{10} V/m
Attenuation length (L_{att}^B)	1.5 mm
Crystal thickness (ℓ)	500 μm

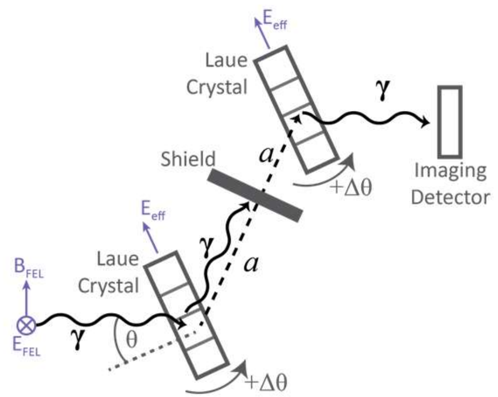
The Borrmann effect increases the extinction length by $\sim 10^3$ for Laue-case diffraction

- Conversion probability - $P(a \leftrightarrow \gamma) \propto L_{ext}^2$
- **Laue-case diffraction**
  - Scatter off planes normal to crystal surface
  - Anomalous extinction, $L_{\text{ext}}^{B} = 1500\ \mu\text{m}$
  - Due to the Borrmann effect
- **Bragg case diffraction**
  - Scatter off planes parallel to crystal surface
  - Extinction, $L_{\text{ext}} = 1\ \mu\text{m}$
- In both cases diffraction follows Bragg's law

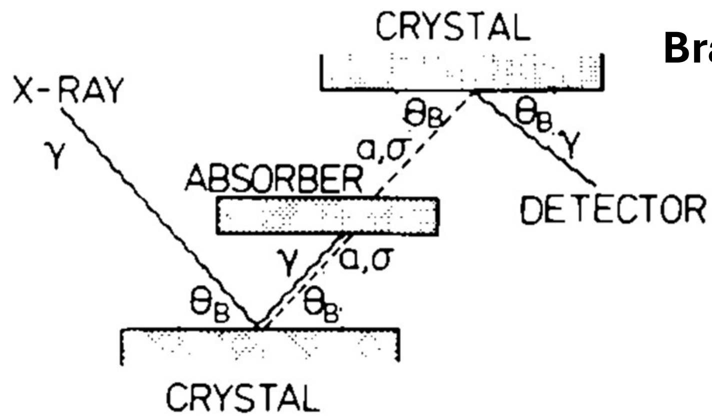
$$2d\sin\theta_B = n\lambda$$

G. Borrmann *"Über Extinktionsdiagramme von Quarz"* Physikalische Zeitschrift **42** (1941).

Laue



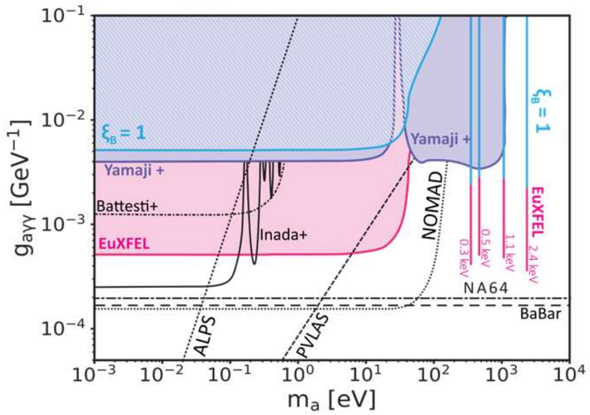
Bragg

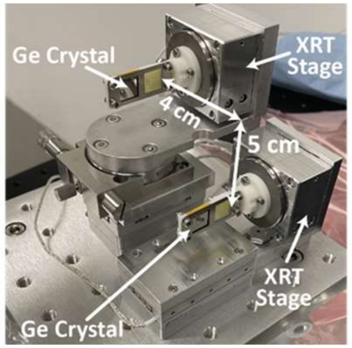


[Buchmüller & Hoogeveen – 1990]

Overview

- Introduction to Axions / ALPS and overview of search strategies
- Experimental design for a light shining through a wall experiment with an X-Ray Free Electron Laser (XFEL)
- **Experimental data from initial experiments on EuXFEL & expected detection threshold for future experiments**
- Fundamental physics with high powered lasers



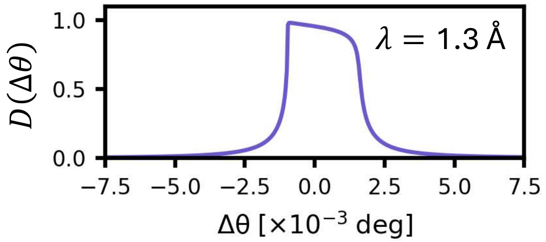


The measured rocking curve is a convolution of the XFEL spectrum and the Darwin curve of the crystal

- Measured transmission given by

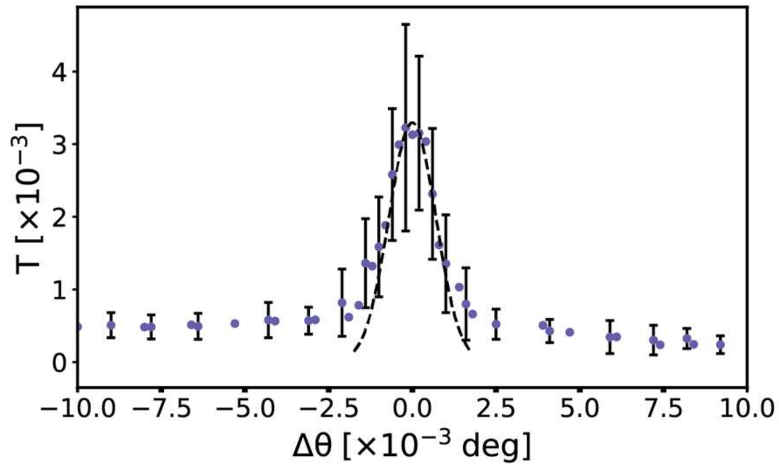
$$T(\Delta\theta) = \frac{\int F^*(\lambda) d\lambda}{\int F(\lambda) d\lambda} = \frac{\int D(\lambda, \Delta\theta) F(\lambda) d\lambda}{\int F(\lambda) d\lambda}$$

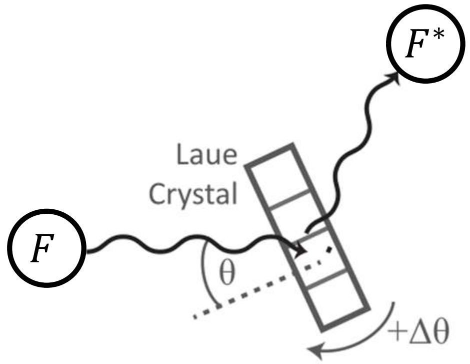
- The Darwin curve, $D(\lambda, \Delta\theta)$ is the transmission through the crystal for a given wavelength as a function of $\Delta\theta = \theta - \theta_B$



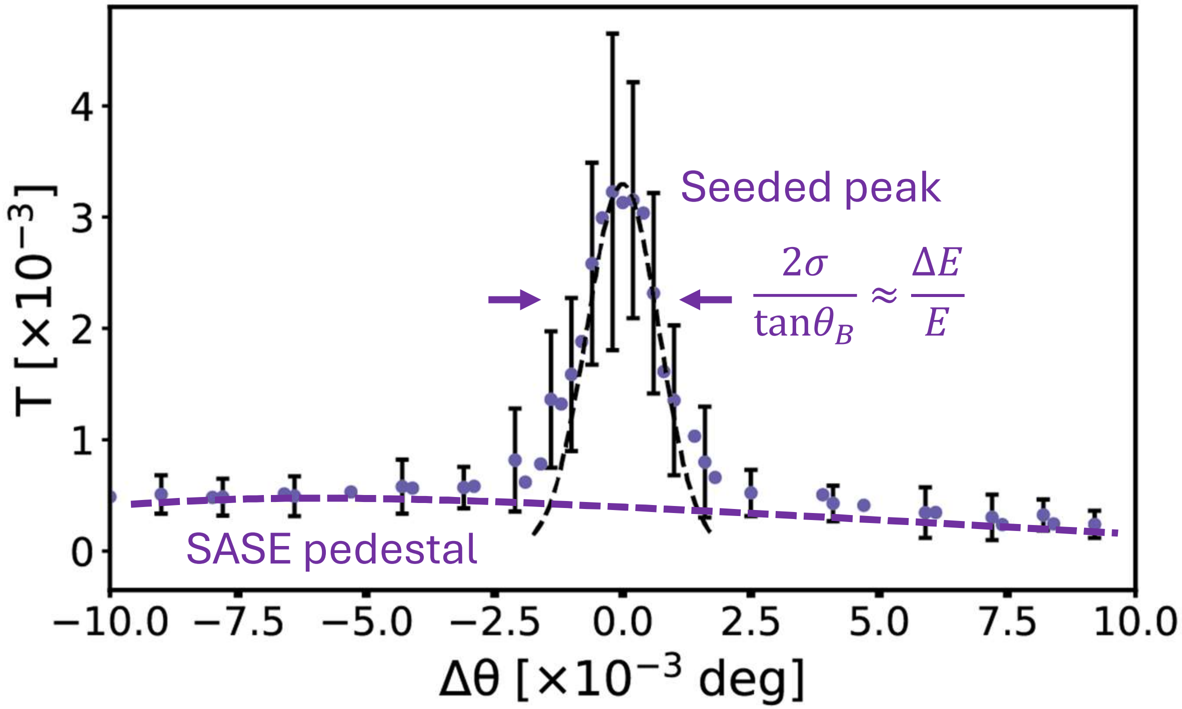
- Curve centred on $\theta_B \rightarrow 2d\sin\theta_B = \lambda$

Photon energy / wavelength	9.8 keV / 1.3 Å
Bandwidth ($\Delta E/E$)	5×10^{-5}





The bandwidth of the XFEL limits the angular resolution in a rocking curve measurement



Photon energy / wavelength	9.8 keV / 1.3 Å
Bandwidth ($\Delta E/E$)	5×10^{-5}

It was found that the X-ray pulse duration was short compared to the timescale associated with diffraction

- For a transform limited pulse:

$$\Delta\theta_{RC}\Delta t \approx \frac{\lambda_x \tan\theta_B}{c}$$

- The timescale for diffraction is:

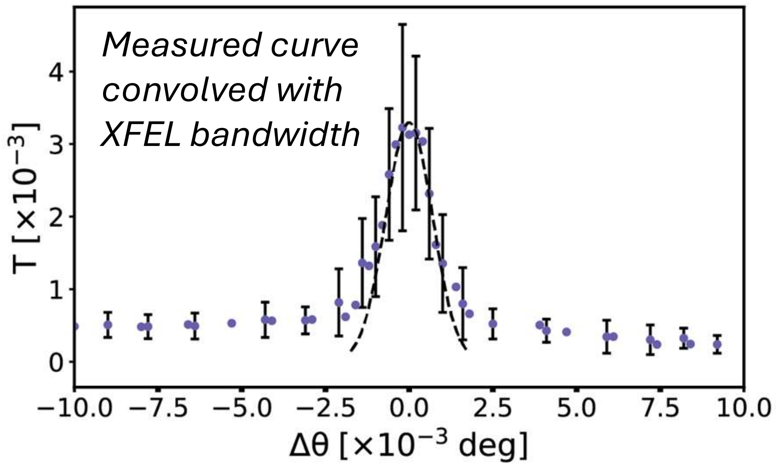
$$\Delta t = (2\ell/c)\tan\theta_B \sin\theta_B$$

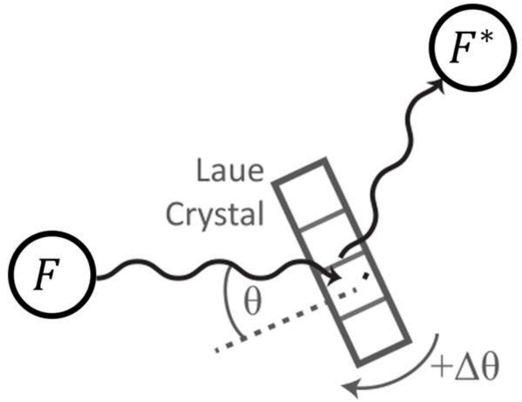
- Consider 2 measures of angular spread of transmitted X-rays:

Rocking curve width ($\Delta\theta_{RC}$): *Actual* angular spread

Darwin width ($\Delta\theta_D$): *Predicted* spread using a theory that neglects time dependence

Wark & Lee, "*Simulations of femtosecond X-ray diffraction from unperturbed and rapidly heated single crystals*", Journal of Applied Crystallography **32** (1999).





It was found that the X-ray pulse duration was short compared to the timescale associated with diffraction

- For a transform limited pulse:

$$\Delta\theta_{RC}\Delta t \approx \frac{\lambda_x \tan\theta_B}{c}$$

- The timescale for diffraction is:

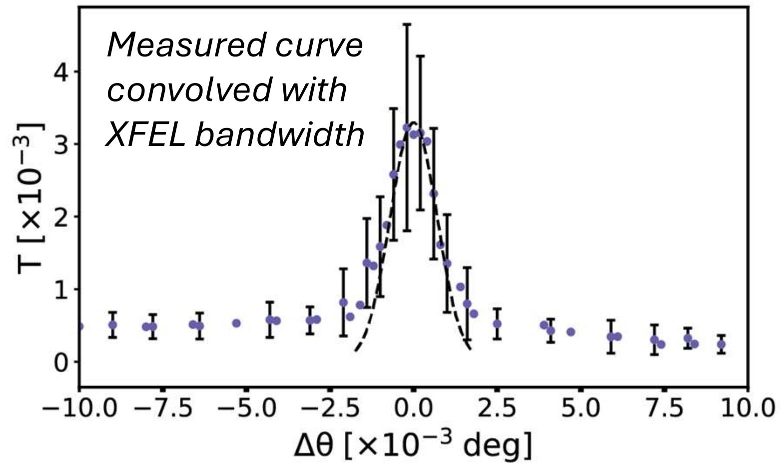
$$\Delta t = (2\ell/c)\tan\theta_B \sin\theta_B$$

- Consider 2 measures of angular spread of transmitted X-rays:

Rocking curve width ($\Delta\theta_{RC}$): *Actual* angular spread

Darwin width ($\Delta\theta_D$): *Predicted* spread using a theory that neglects time dependence

Wark & Lee, "*Simulations of femtosecond X-ray diffraction from unperturbed and rapidly heated single crystals*", Journal of Applied Crystallography **32** (1999).



$L^B_{att} \gg \ell \Rightarrow$ Expect full transmission

Attenuation length (L^B_{att})	1.5 mm
Crystal thickness (ℓ)	500 μm

Rocking curve width ($\Delta\theta_{RC}$)	~0.4 μrad
Darwin width ($\Delta\theta_D$)	44 μrad

Probability for photon-axion conversion should account for the short pulse-duration

- There is an implicit assumption in theory from Yamaji *et al*. that $\Delta\theta_{RC} = \Delta\theta_D$
- Expression for conversion probability modified to account for this:

$$P(a \leftrightarrow \gamma) \approx \left(\frac{1}{4} g_{a\gamma\gamma} \xi_B E_{\text{eff}}\, L_{\text{eff}} \cos\theta_B\right)^2$$

$$\xi_B = \Delta\theta_D / \Delta\theta_{RC}$$

- Time-bandwidth product – expression accurate to factor of order unity
- Physical interpretation: increase in refractive index of crystal

T. Yamaji *et al, "Theoretical calculation of coherent Laue-case conversion between x-rays and ALPS for an X-ray LSW experiment"* Phys. Rev. D **96**, 15001 (2017).

Rocking curve width ($\Delta\theta_{RC}$): *Actual* angular spread of transmitted X-rays

Darwin width ($\Delta\theta_D$): *Predicted* spread using a theory that neglects time dependence

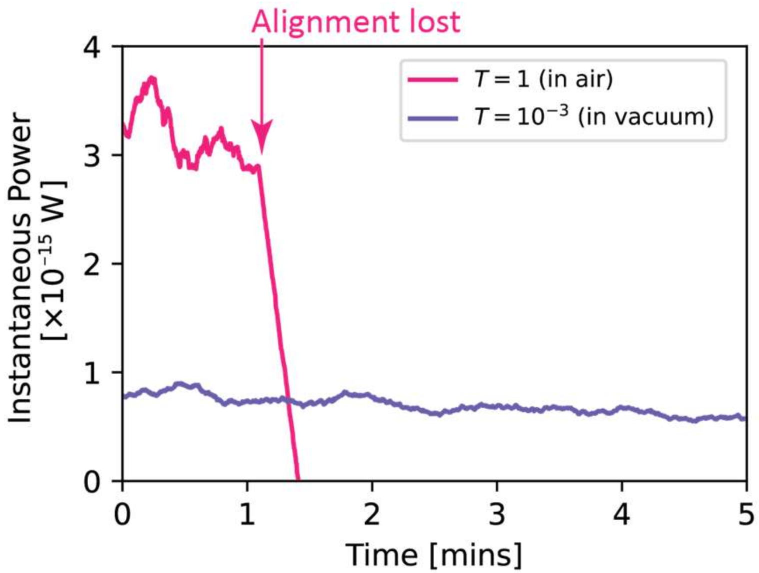
Rocking curve width ($\Delta\theta_{RC}$)	~ 0.4 µrad
Darwin width ($\Delta\theta_D$)	44 µrad
ξ_B	~ 110

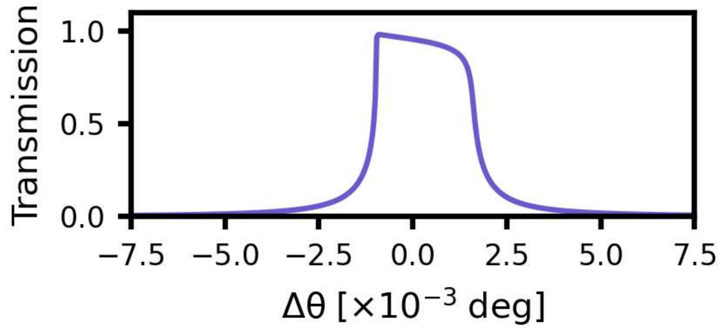
"Brilliance" of XFEL amplifies conversion probability

$(N_\gamma /\ \text{mm}^2 /\ \text{s} / \text{mrad}^2 /\ 0.1\%\text{BW})$

Crystal heating was a major problem during the beamtime

- Maintaining alignment is challenging as the heat load distorts the crystal lattice
- Rocking curve width for Germanium $\sim 5 \times 10^{-3}$ degrees
- Able to maintain stable alignment with 10^3 attenuation applied to the XFEL beam
- Limited search to acquiring data at θ_B and at 4 discrete values of $\Delta\theta$
- In future experiments can mitigate:
  - Active cooling of conversion crystal





Due to crystal heating, a multi-step process was used for data collection

- Process for data collection:
  1. Stages tuned to θ_B and radiation shield removed. ~2 mins characterization data collected.
  2. Stages tuned to search angle & radiation shield returned. ~10 mins search data collected.
  3. Stages tuned back to θ_B and shield removed. ~2 mins characterization data collected.
- For the i'th run calculate:

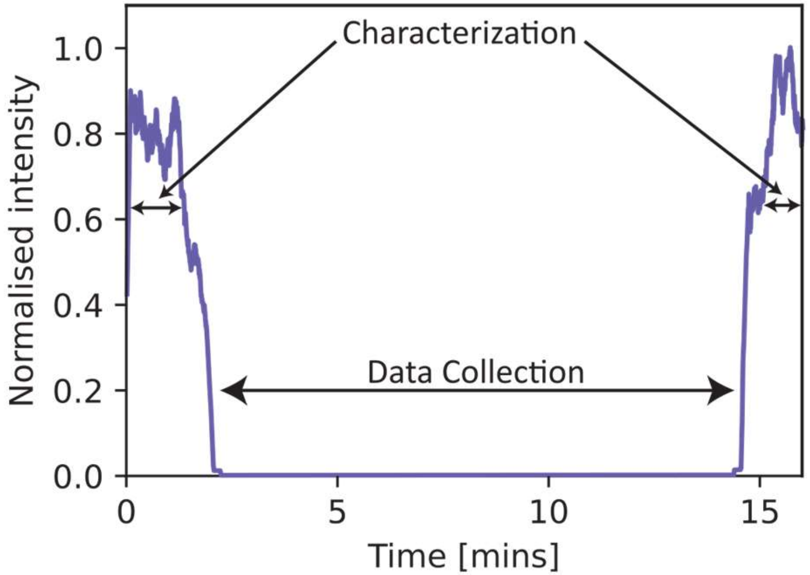
$$\eta_i = \frac{1}{T_{\text{Ge}}^2} \frac{E_i^{\text{JF,ch}}}{E_i^{in,ch}}$$

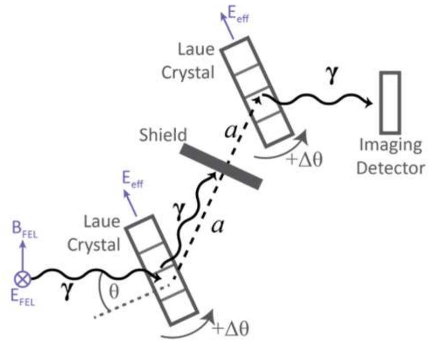
- Then, for all runs at a given $\Delta\theta$:

$$\eta N_{in} = \sum_i \eta_i E_i^{\text{in,aq}} / k_\gamma$$

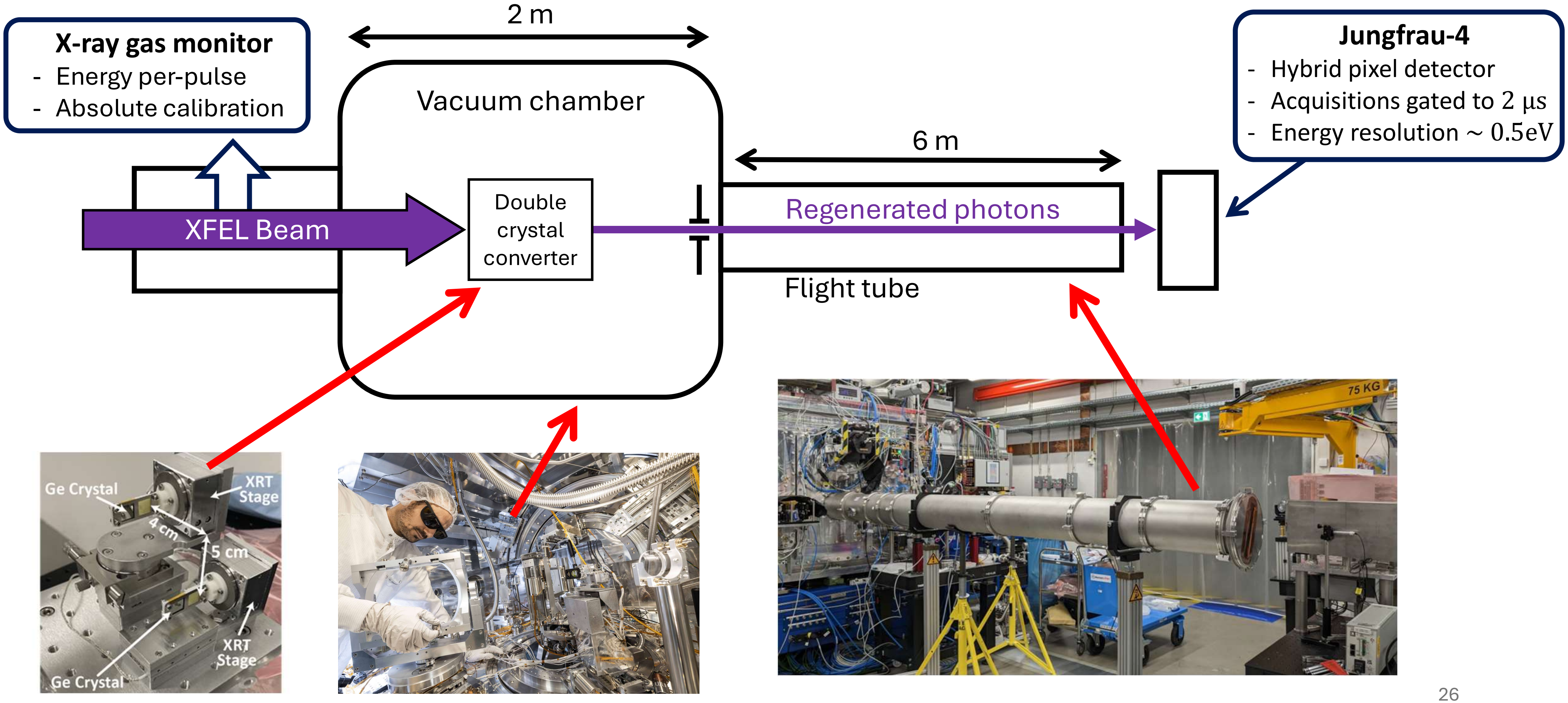
- Measured probability of axion generation then:

$$P(a \leftrightarrow \gamma)^2 = \frac{N_{\text{det}}}{\eta N_{in}}$$

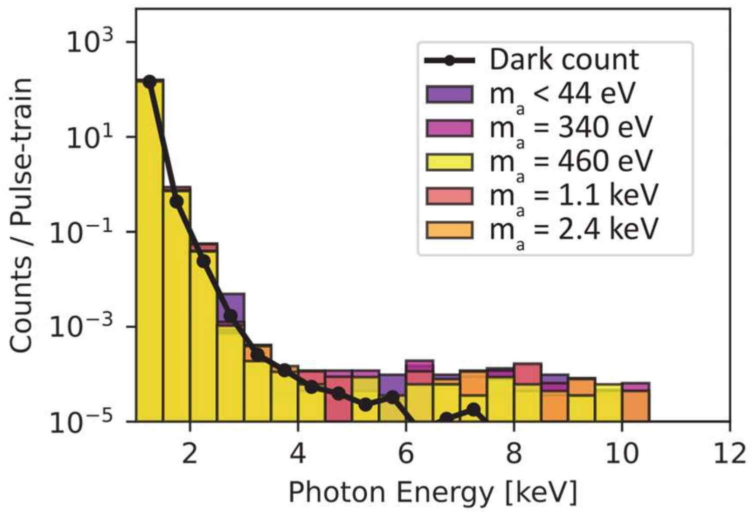


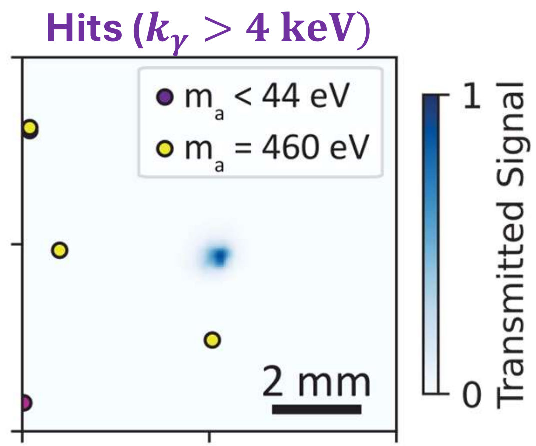


The HED/HiBEF instrumentation enabled an accurate constraint to be placed on input / output X-ray flux



No events consistent with axion production were detected during the experiment





Events consistent with axion production have $k_\gamma = 9.8\ \mathrm{keV}$.

They should also lie on the X-Ray spot observed on the Jungfrau when in characterisation mode.

Bounds on the coupling strength were calculated taking a 90% confidence interval

- Measured probability of axion generation:

$$P(a \leftrightarrow \gamma)^2 = \frac{N_{\text{det}}}{\eta N_{in}}$$

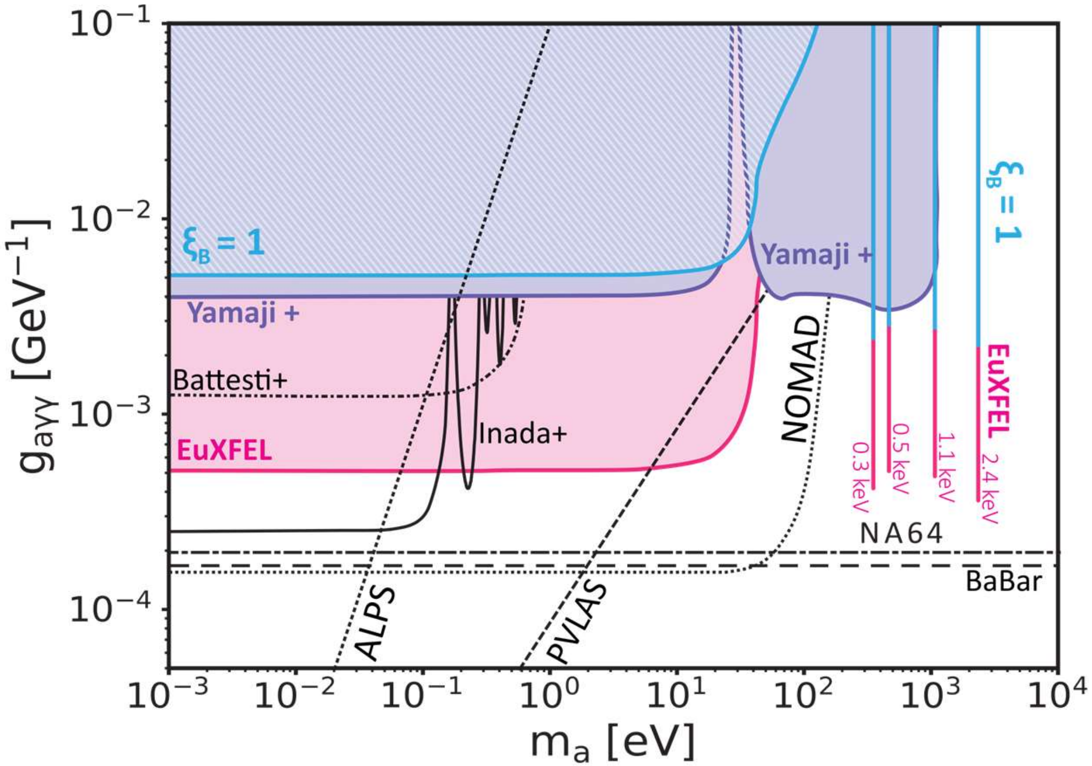
- Take $N_{\text{det}} = 2.3$ events (assume events Poisson distributed, 90 % confidence interval with $\mu = 0$)
- Upper bound on axion-photon coupling:

$$g_{a\gamma\gamma} < \left(\frac{1}{4} E_{\text{eff}} L_B \xi_B \cos\theta_B\right)^{-1} P(a \leftrightarrow \gamma)^{1/2}$$

$\Delta\theta$ [mrad]	m_a [eV]	N_{in} ($\times 10^{16}$)	$g_{a\gamma\gamma}$ [Gev^{-1}]
0.0	< 44	2.6	3.91×10^{-4}
1.0	3.4×10^2	2.4	3.10×10^{-4}
1.8	4.6×10^2	1.6	3.87×10^{-4}
10.0	1.1×10^3	1.7	3.69×10^{-4}
50.0	2.4×10^2	1.5	2.76×10^{-4}

T. Junk *“Confidence level computation for combining searches with small statistics”* Nucl. Instrum. Meth. A **434,** 435 (1999)

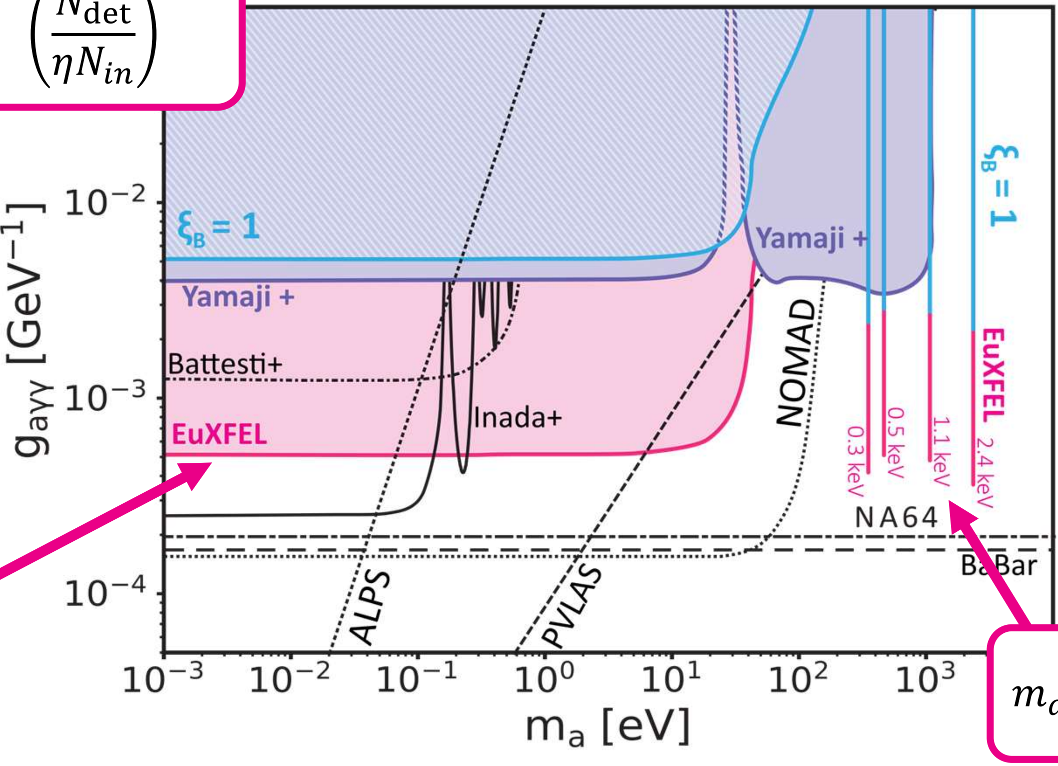
Bounds obtained in this work are competitive with existing studies - using a complementary search technique



J. W. D. Halliday *et al. "Bounds on heavy axions with an X-ray free electron laser"* Accepted for publication in Phys. Rev. Lett. (2025)

Bounds obtained in this work are competitive with existing studies - using a complementary search technique

$$g_{a\gamma\gamma} < \left(\frac{1}{4}E_{\text{eff}}L_B\xi_B\cos\theta_B\right)^{-1}\left(\frac{N_{\text{det}}}{\eta N_{in}}\right)^{-1/4}$$



$$|m_a^2 - m_\gamma^2| \leq \frac{4k_\gamma}{L_{eff}}$$

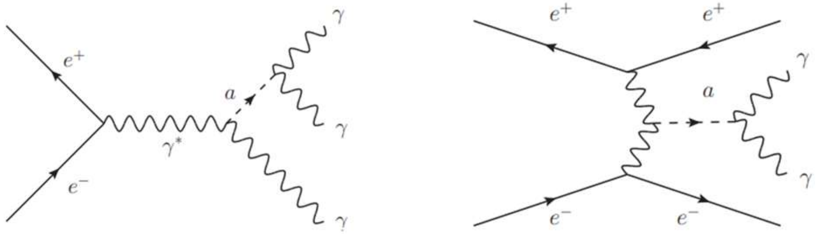
$$m_a = \sqrt{m_\gamma^2 + 2q_T k_\gamma \cos(\theta_B)\Delta\theta}$$

J. W. D. Halliday *et al.* *"Bounds on heavy axions with an X-ray free electron laser"* Accepted for publication in Phys. Rev. Lett. (2025)

Our bounds are complimentary to those imposed from searches using accelerators

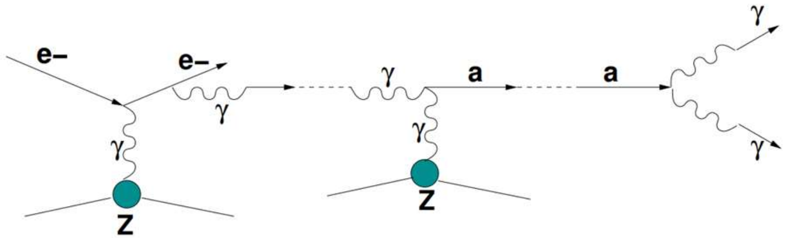
BaBar

- Bounds on dark photon detection [BaBar Collaboration] recast by Dolan *et al*. to obtain a bounds on $g_{a\gamma\gamma}$.
- Probes spontaneous axion decay



NA64

- Target-Bremsstrahlung photons generated by 100 GeV electrons in NA64 beam dump produce axions via the Primakoff effect
- Search for both spontaneous axion decay and missing energy in calorimetry data



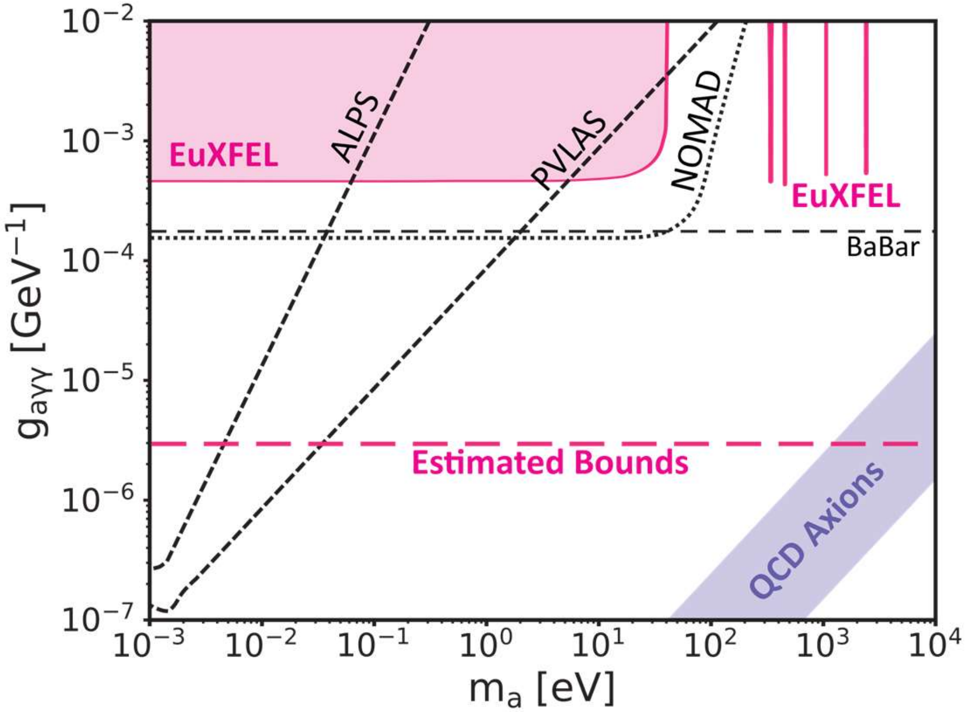
The BaBar collaboration. *"Search for Invisible Decays of a Dark Photon Produced in e⁺e⁻ Collisions at BaBar"*, Phys. Rev. Lett. **119** (2017)

J. Dolan *et al*. *"Revised constraints and Belle-II sensitivity for visible and invisible axion-like particles"*, J. High Energ. Phys. 94 (2017)

The NA64 Collaboration, "Search for Axionlike and Scalar Particles with the NA64 Experiment" Phys. Rev. Lett. **125** (2020)

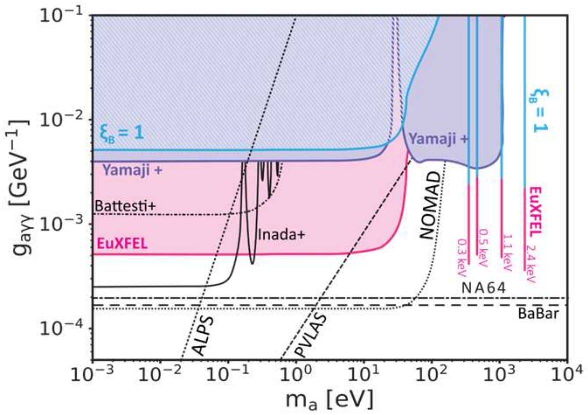
From experience gained in these initial experiments, we see a pathway to search for QCD axions with $\sim 1 - 10\ \mathrm{keV}$ mass

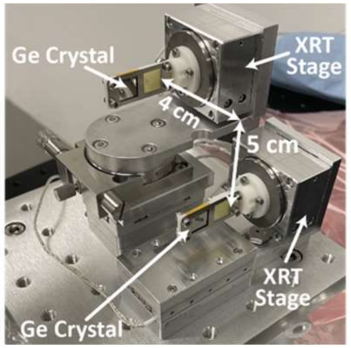
- First results encouraging, but the experiment is far from perfect – major issue with crystal heating
- Negating these issues – increase flux on-target by a factor 3×10^5:
  - Attenuated X-ray beam by 10^3
  - Used 1 out of 300 pulses per train
  - Improve by actively cooling the conversion crystal
- Thicker germanium crystals ($\ell \sim L^B_{\mathrm{ext}} = 1.5\ \mathrm{mm}$)
- Together these imply in an $\sim 150 \times$ increase in detection threshold
  ($g_{a\gamma\gamma} > 2 \times 10^{-6}\ \mathrm{GeV}^{-1}$)
- Projected to be sensitive to QCD axions with $m_a \sim 1 - 10\ \mathrm{keV}$



Overview

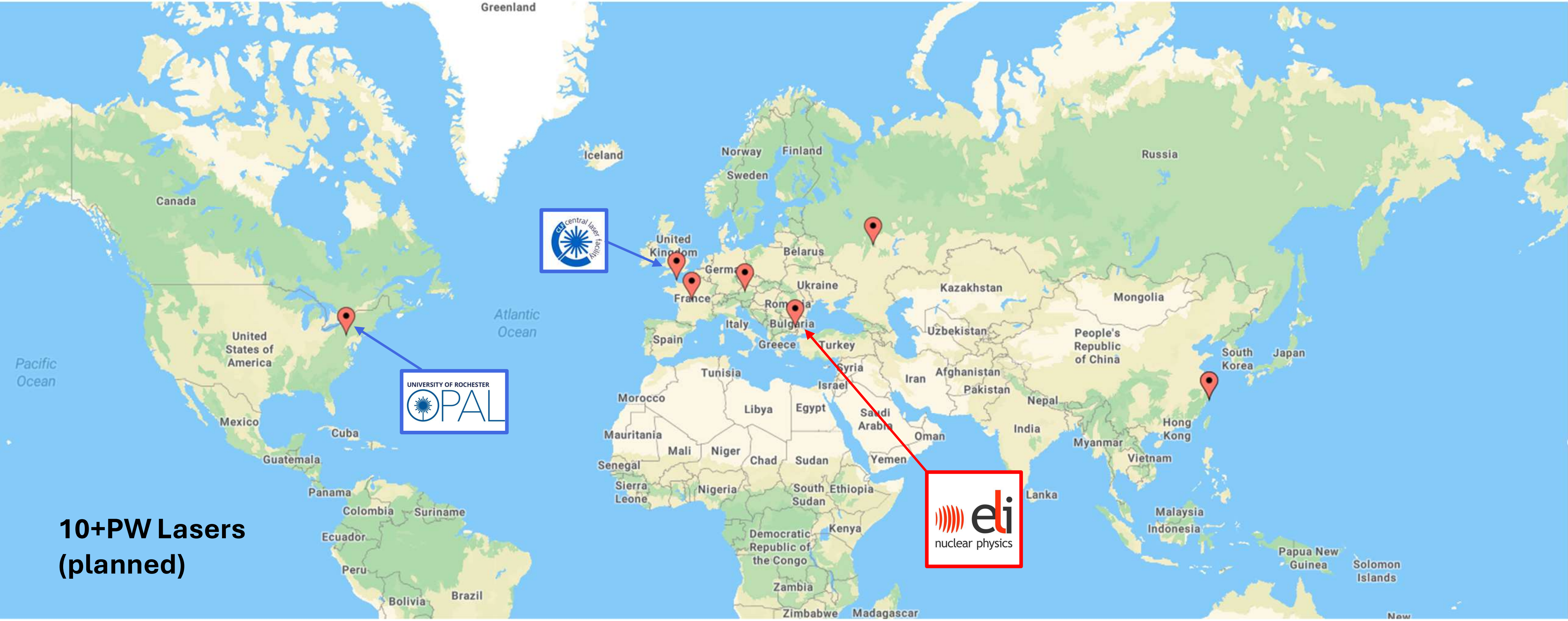
- Introduction to Axions / ALPS and overview of search strategies
- Experimental design for a light shining through a wall experiment with an X-Ray Free Electron Laser (XFEL)
- Experimental data from initial experiments on EuXFEL & expected detection threshold for future experiments
- **Fundamental physics with high powered lasers**



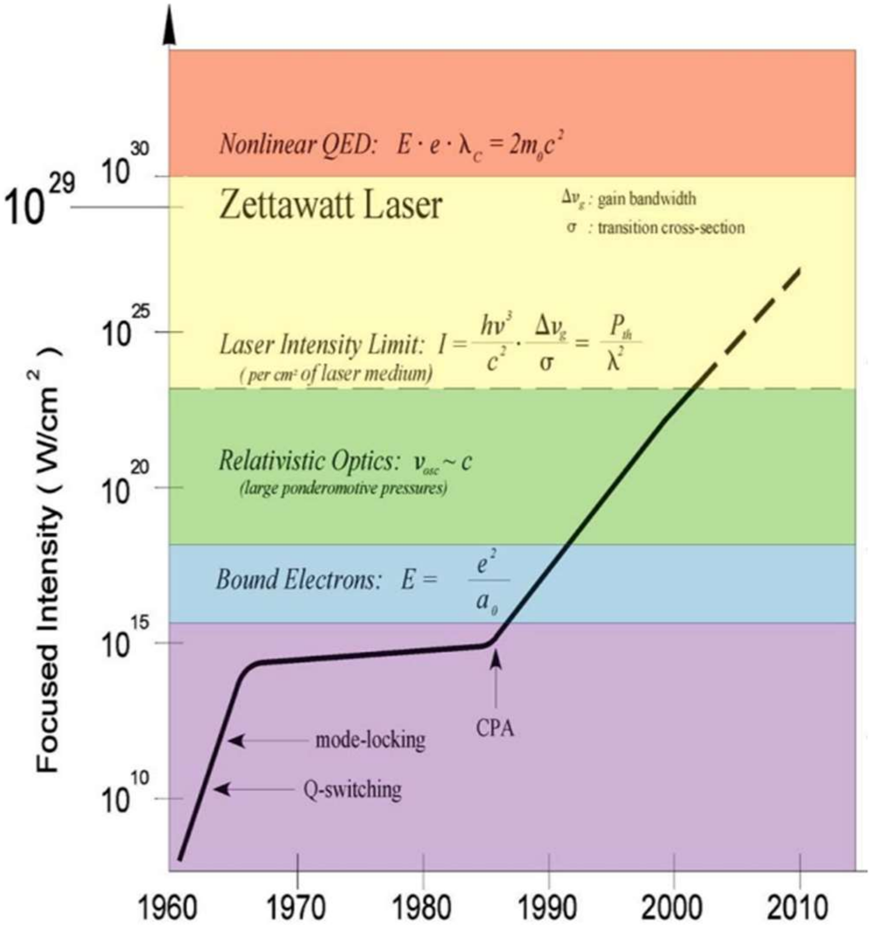


There are a growing number of PW and multi-PW laser systems globally

There are a growing number of PW and multi-PW laser systems globally



High-powered lasers have enormous potential for fundamental physics applications



- Progress enabled by CPA technique
- Normalized vector potential (peak electron velocity / m_e)

$$a_0 = \frac{eE}{mc\omega} = 0.6 \left(\frac{I}{10^{18} W/cm^2}\right)^{1/2} \left(\frac{\lambda}{\mu m}\right)$$

- $a_0 > 1$ implies relativistic motion for the electron
- Quantum non-linearity parameter (peak E_L/E_{crit})

$$\eta = \frac{2\, a_0^2 \hbar\omega}{mc^2} = 0.18 \left(\frac{I}{10^{23} W/cm^2}\right) \left(\frac{\lambda}{\mu m}\right)$$

- $\eta > 1$ means that pair production is important

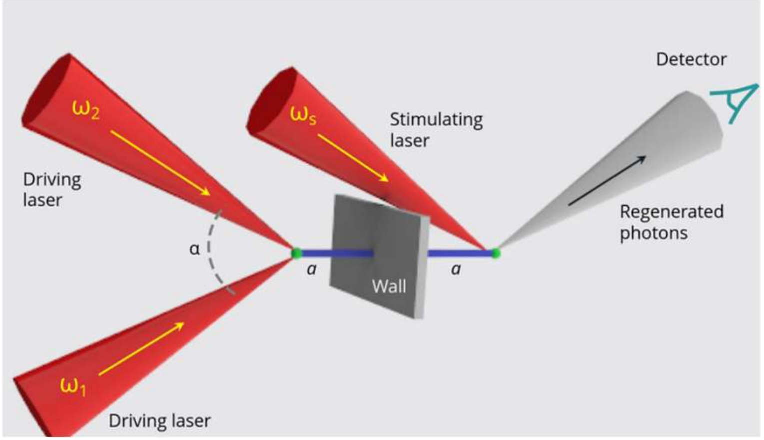
E.g. consider an axion search, with optical lasers, sensitive to $m_a = 1 - 10$ eV

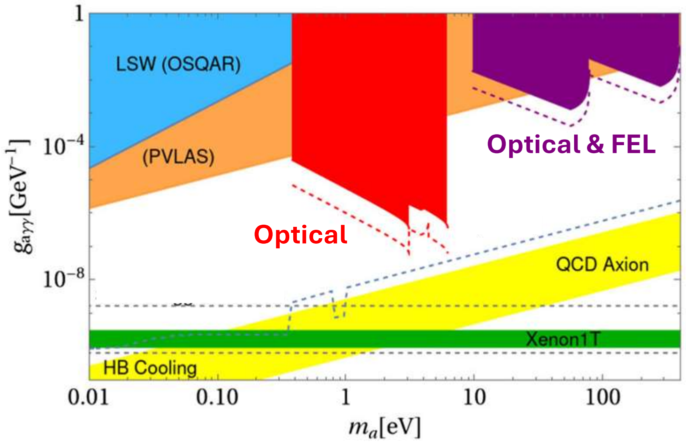
- Two short-pulse laser beams with orthogonal polarisation made to cross in a vacuum
- Axion production:
$$L_{\text{axion}} = g_{a\gamma\gamma} \boldsymbol{E} \cdot \boldsymbol{B} a$$
- Regenerate photons via inverse process, using a 3rd beam
- Projected bounds are for Aton-4 laser (Extreme Light Infrastructure, Czechia):

$$g_{a\gamma\gamma} \geq 3.5 \times 10^{-7}\ \text{GeV}^{-1} \left(\frac{1.5\ \text{kJ}}{E_1}\right)^{\frac{1}{4}} \left(\frac{1.5\ \text{kJ}}{E_2}\right)^{\frac{1}{2}} \left(\frac{d}{10\ \text{cm}}\right)^{\frac{1}{4}} \left(\sqrt{1 - \left(\frac{m_a}{3.08\ \text{eV}}\right)^2}\right)^{\frac{1}{4}} \left(\frac{3.08\ \text{eV}}{m_a}\right)^2 \left(\frac{R_\gamma}{\text{day}^{-1}}\right)^{\frac{1}{4}},$$

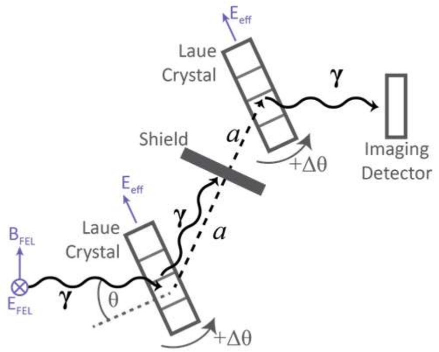
- Potential to adapt for studies on Vulcan-2020 or EPAC

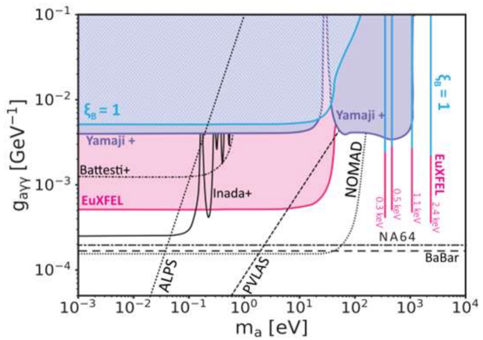
K. Beyer *et al. "Light-shining-through-wall axion detection experiments with a stimulating laser"* Phys. Rev. D **105**, 035031 (2022)

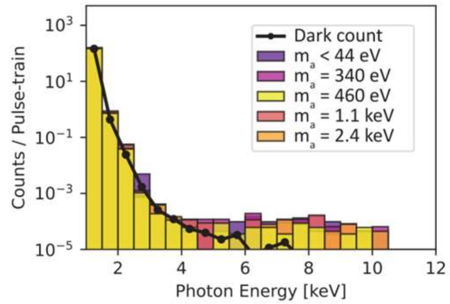


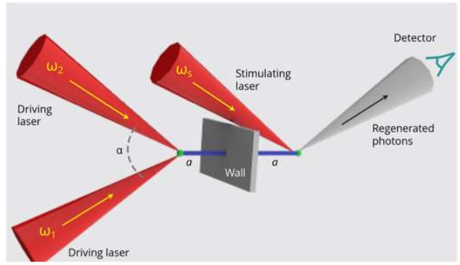


Conclusions









- Axions explain the absence of CP-violation in strong interactions and are a cold dark matter candidate
- We performed a search for heavy axions on EuXFEL
- Had issues with crystal heating, but we could still impose competitive bounds on $g_{a\gamma}$ using a search technique complimentary to collider-based experiments
- There is a pathway to searching for QCD axions with $m_a \sim 1 - 10\ \mathrm{keV}$ using the same technique
- High powered lasers have potential uses for fundamental physics, including in axion searches.

J. W. D. Halliday *et al. "Bounds on heavy axions with an X-ray free electron laser"* Accepted for publication in Phys. Rev. Lett. (2025)